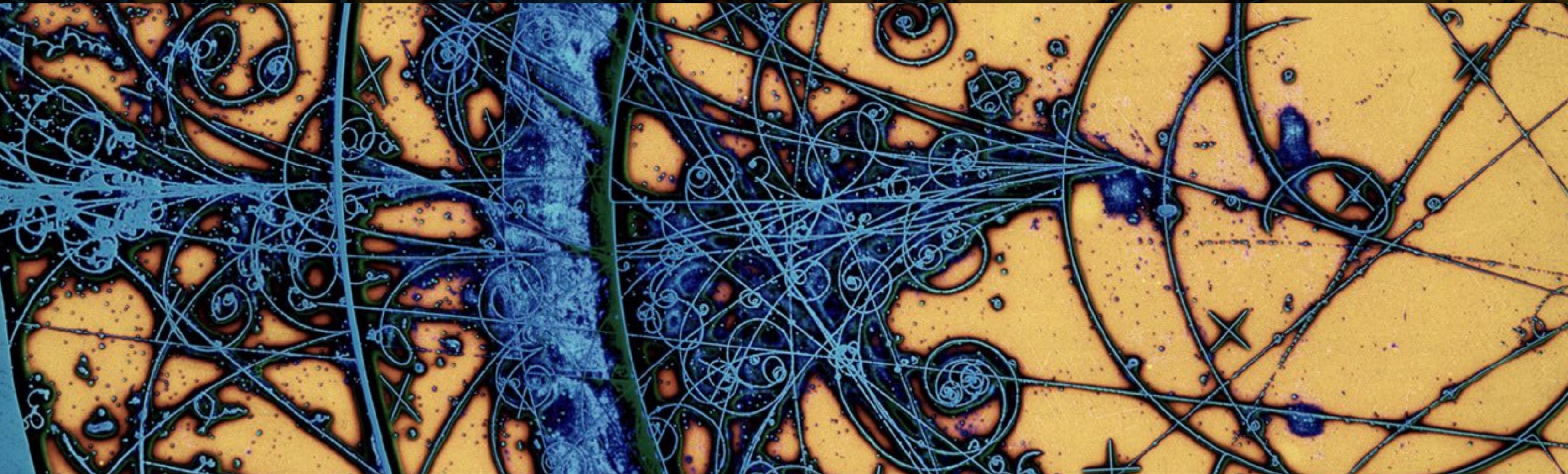


Neutrinos at the Frontiers of Particle Physics: Present and Future

Columbia University, Dec 8th, 2025

BEBC CERN-EX-66954B-1 © 1998-2025 CERN



Matheus Hostert matheus-hostert@uiowa.edu

University of Iowa

IOWA

This talk: some aspects **neutrino phenomenology**

- **Neutrino masses and the Standard Model**
- **Neutrino experiments to search for rare phenomena.**
- **Towards novel neutrino beams in a future μ -collider campus.**





NEAT: Neutrino Experiment and Theory (Workshop)

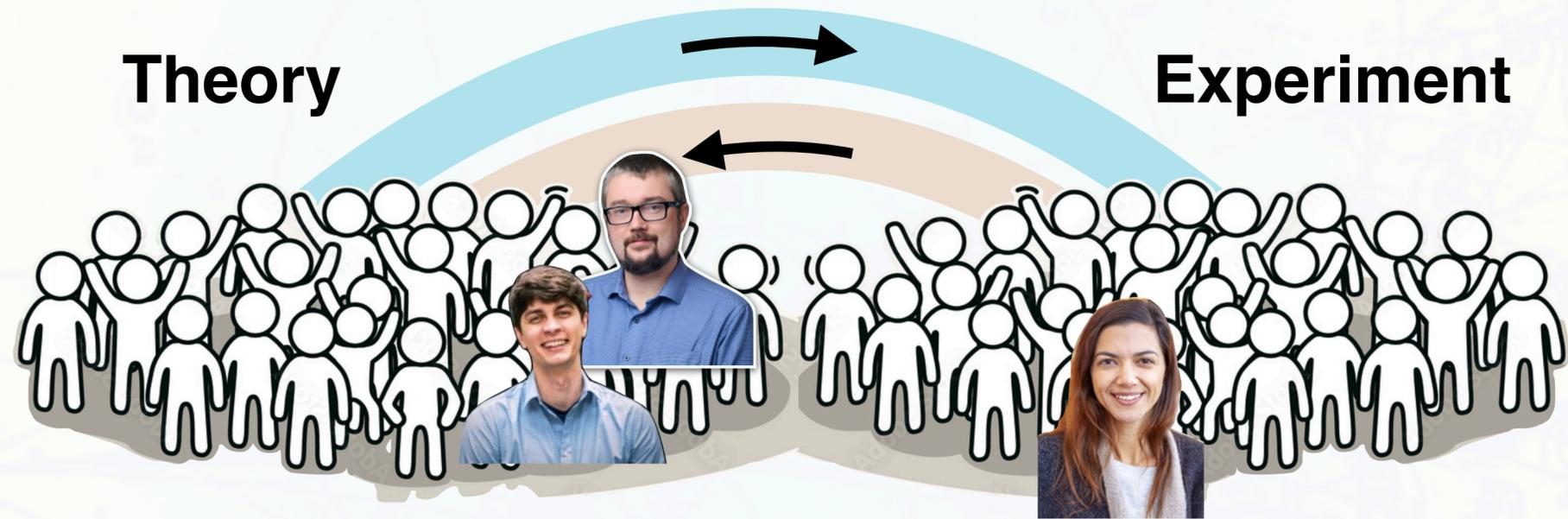
2nd Short-Baseline Experiment-Theory Workshop, April 2nd to 5th 2024

5th New Physics Opportunities at Neutrino Facilities Workshop (NPN 2025)

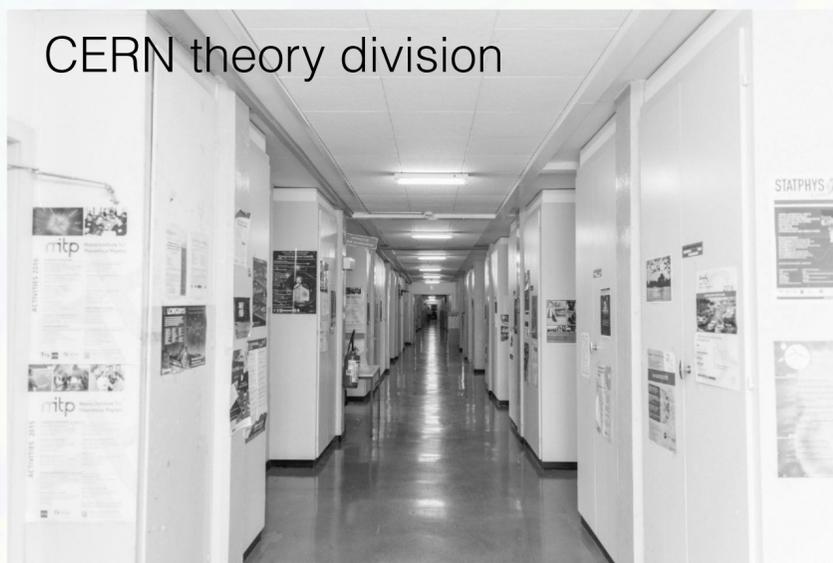
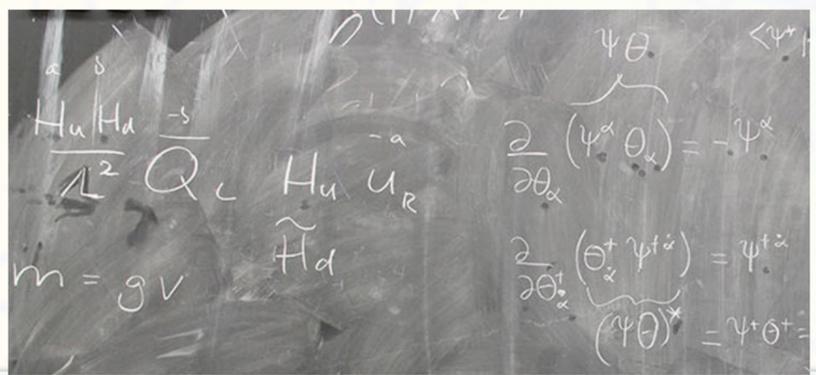


Joint workshops

Phenomenology



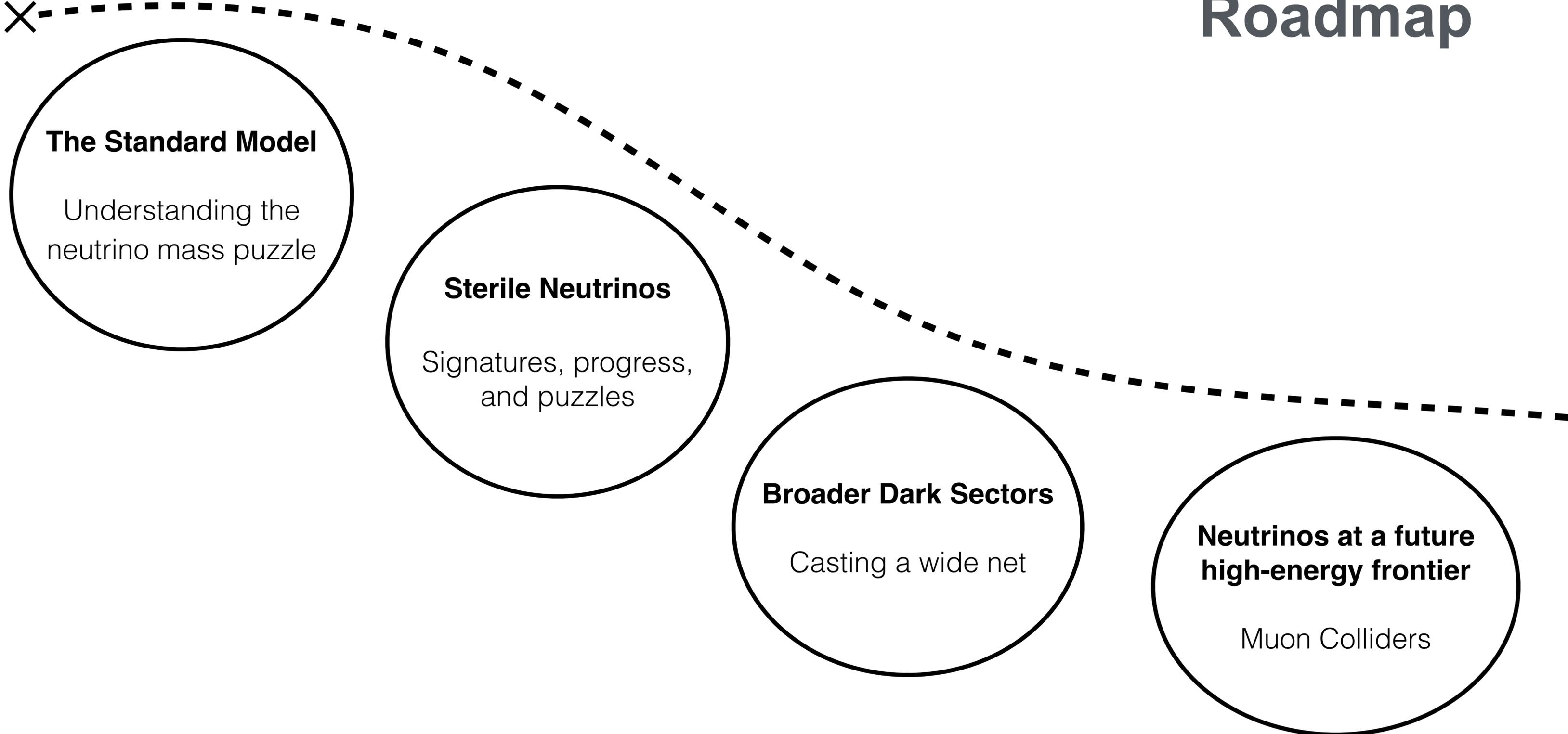
Fermilab Theory division



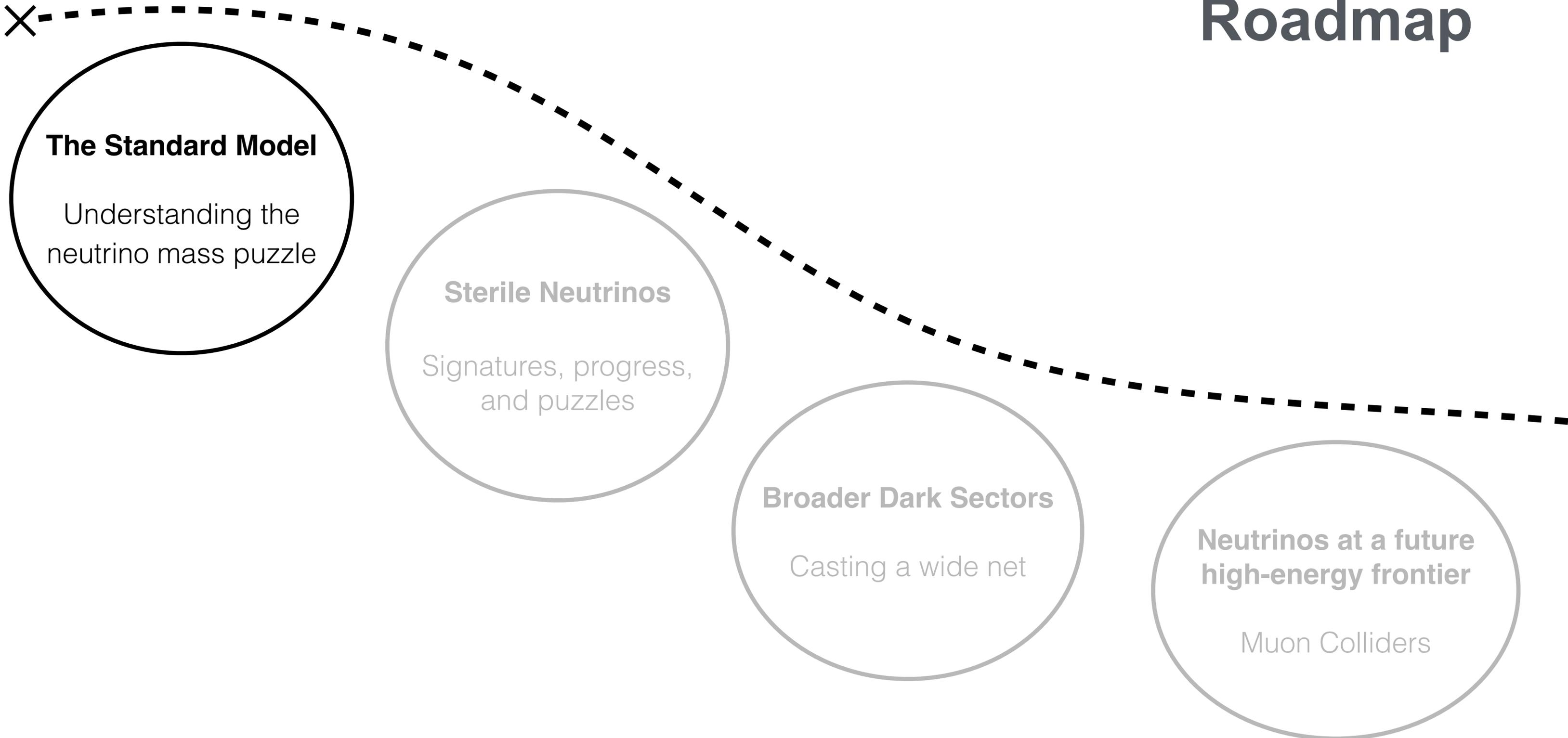
Joint Experimental-Theoretical Physics Seminar

This is the so-called "Wine and Cheese" seminar.
Regular seminars are **Fridays at 3:30 p.m.** in Wilson Hall, One West.

Roadmap



Roadmap



STANDARD MODEL

The SM is not wrong...

But more than “reasonable doubt”
that it is **not** the end of the story.

Incompleteness:

Neutrino masses and mixing

Dark matter

Descriptive rather than predictive:

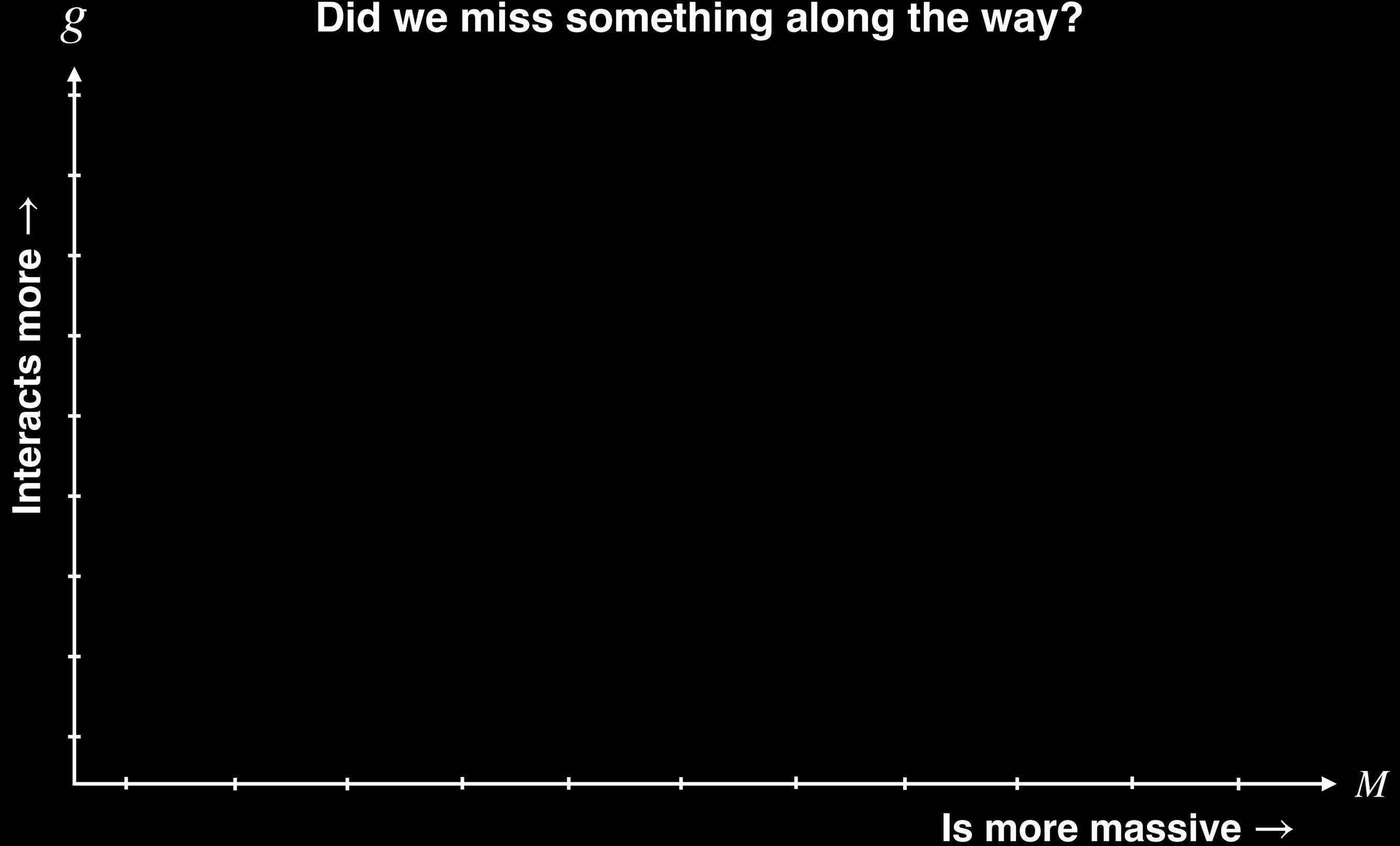
Fermion mass pattern?

Absence of CP violation in strong force?

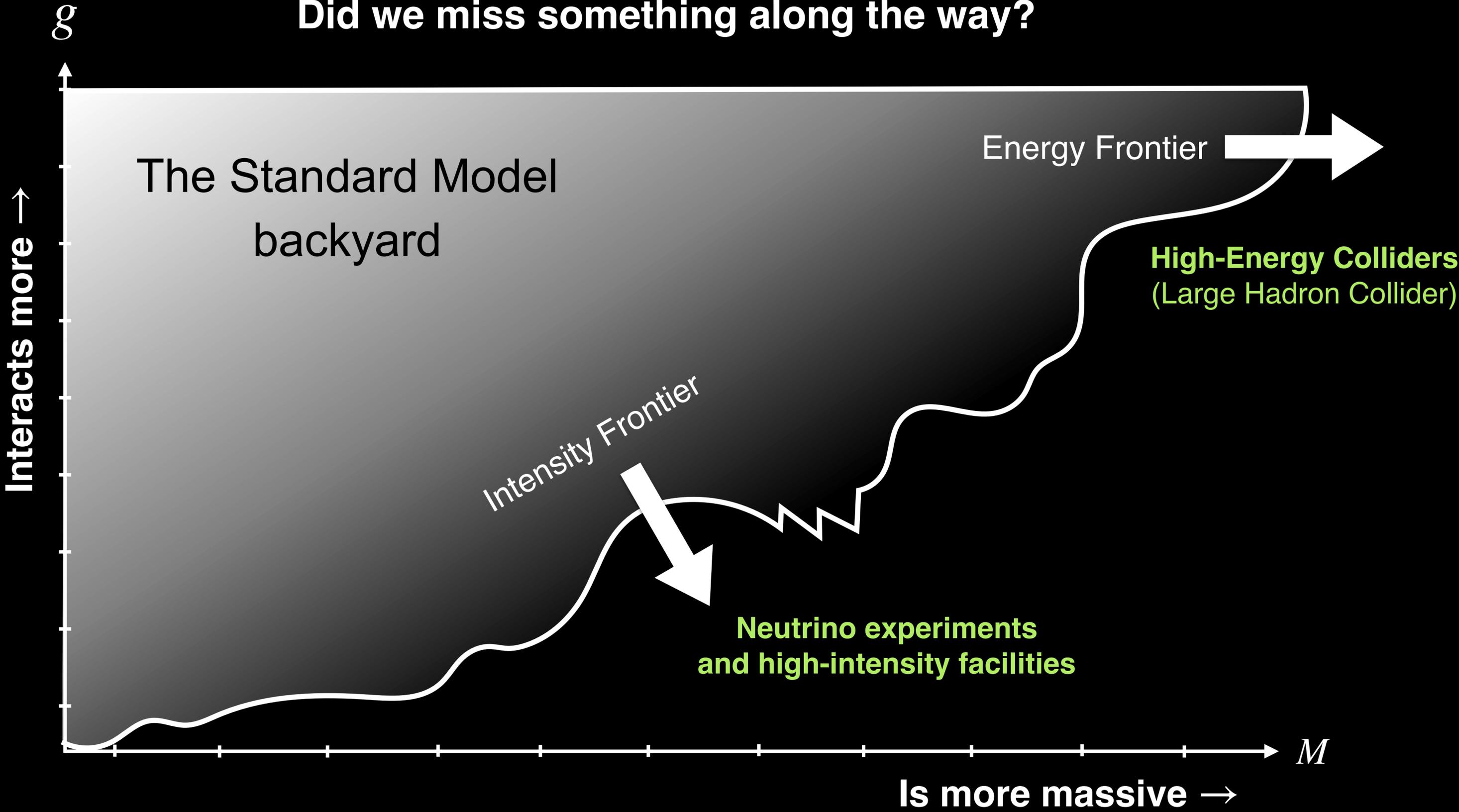
Why is the Higgs mass so small?

Matter-antimatter asymmetry?

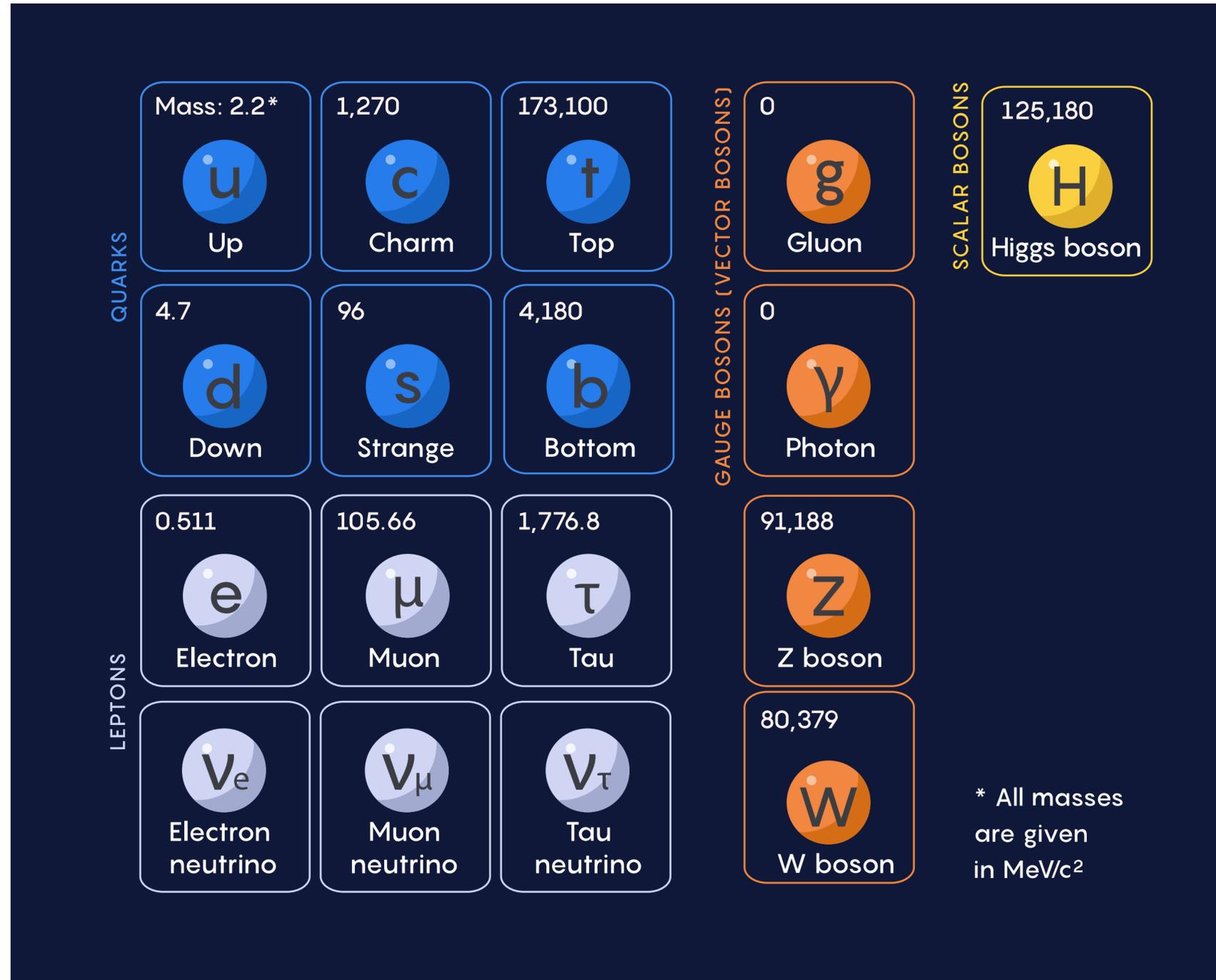
Did we miss something along the way?



Did we miss something along the way?



The Standard Model

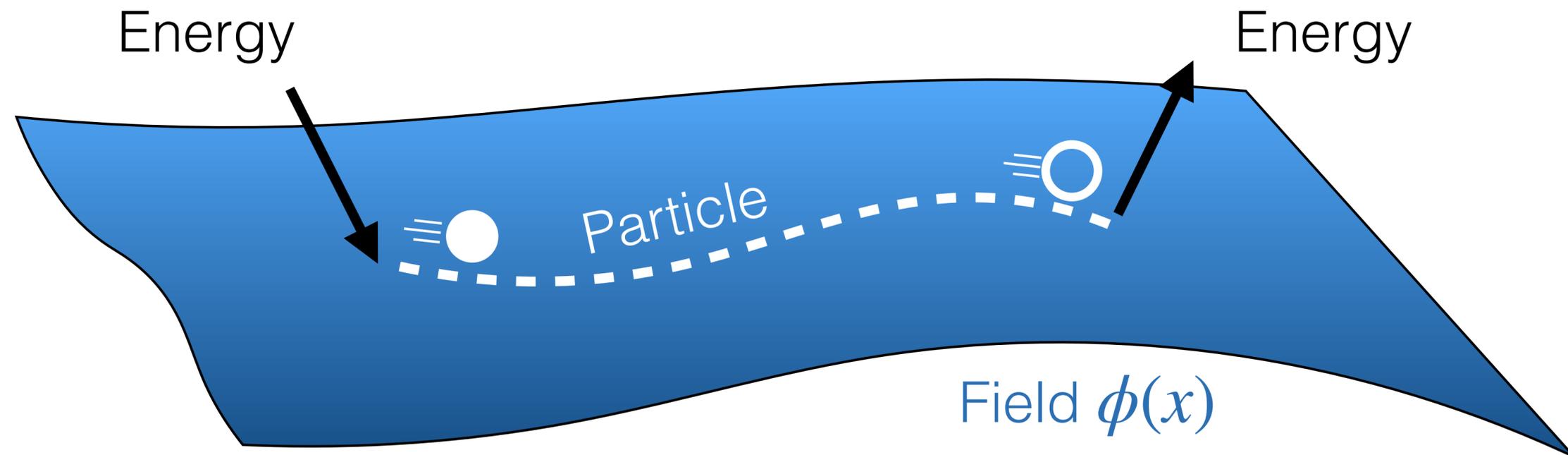


~ 17 particles?



Particles and Fields

Particles are the smallest excitations of “**quantum field**”

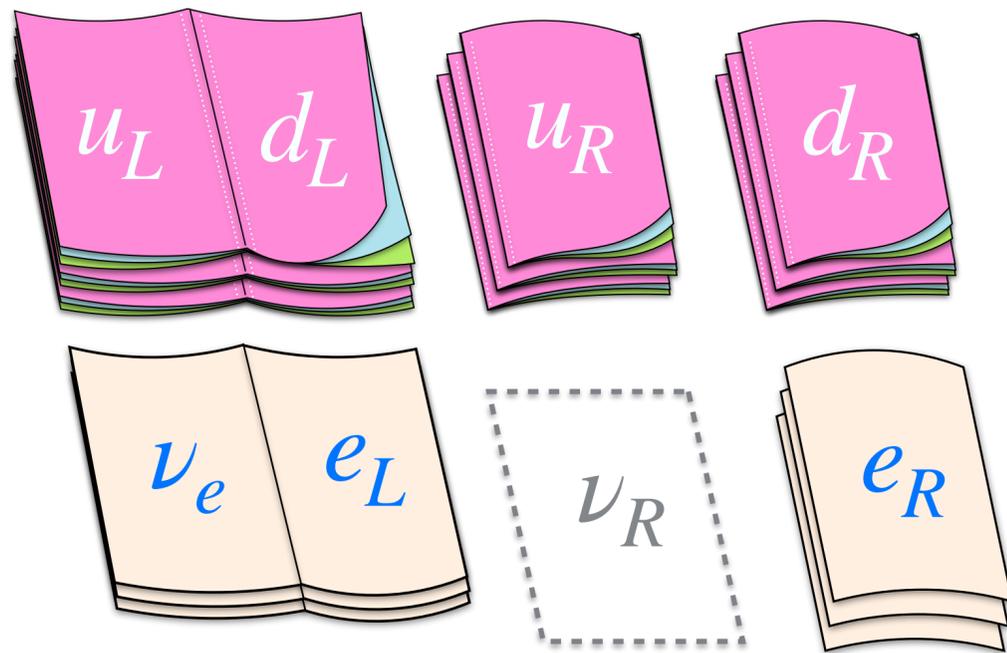


Each particle is associated with its own field

The Standard Model (actually)

More like ~76 different fields...
with a rich symmetry structure.

Matter Content



All Fermions
(Spin-1/2)

The Higgs

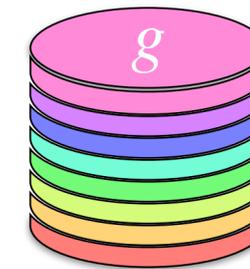


The Higgs Boson
(Spin-0)

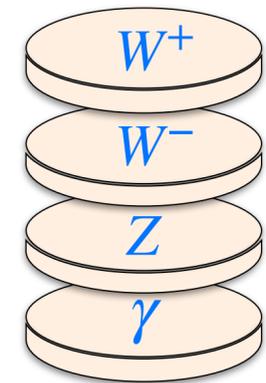
Dynamics

(Symmetries)

Strong Force
 $SU(3)$



Electroweak Force
 $SU(2) \times U(1)$



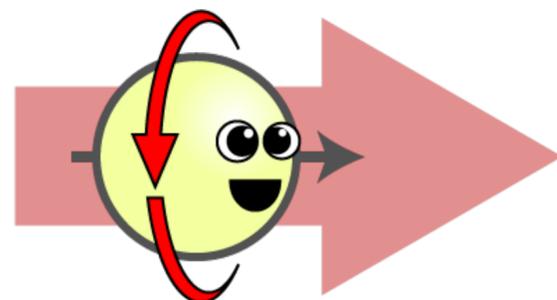
The "two" forces and their mediators
Bosons (Spin-1)

The Standard Model (actually)

More like ~76 different fields...
with a rich symmetry structure.

Chirality or "handedness"

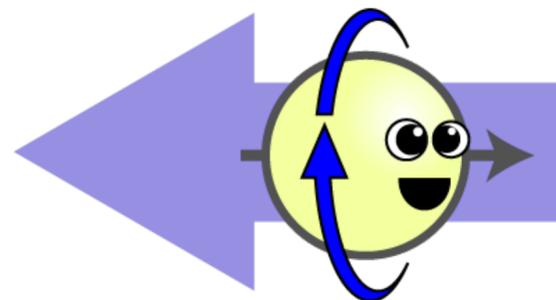
Right-handed $h = + 1/2$



Just another quantum number like charge or spin.

For ultra-relativistic (or massless) particles,
it coincides with **helicity**:

Left-handed $h = - 1/2$



$$h = \frac{\vec{s} \cdot \vec{p}}{|\vec{p}|} = \pm \frac{1}{2}$$

**This turns out to be an extremely important in
the Standard Model.**

Weak Force
U(1)



mediators

<https://www.quantumdiaries.org/2011/08/23/the-spin-of-gauge-bosons-vector-particles/>



Weak Interactions are “Left-Handed”

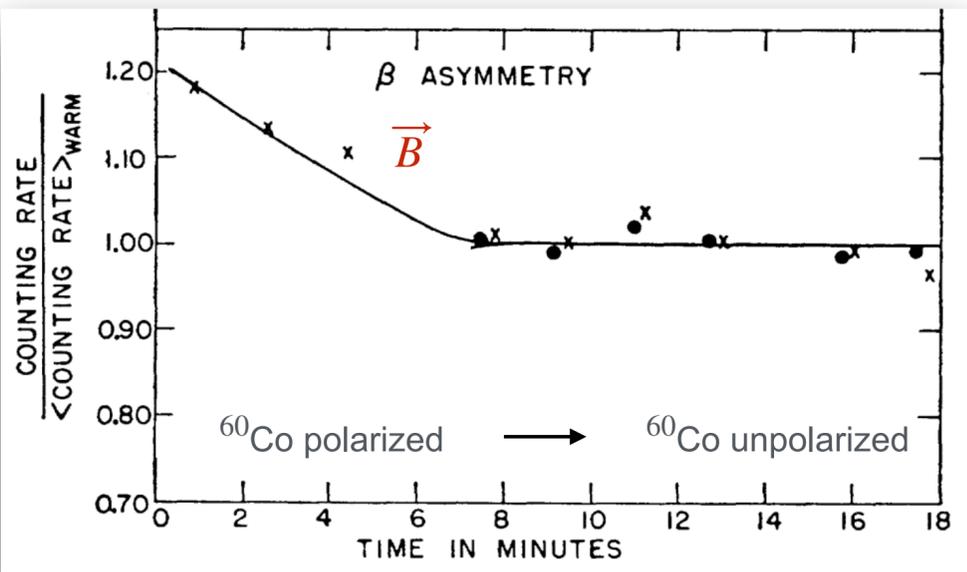
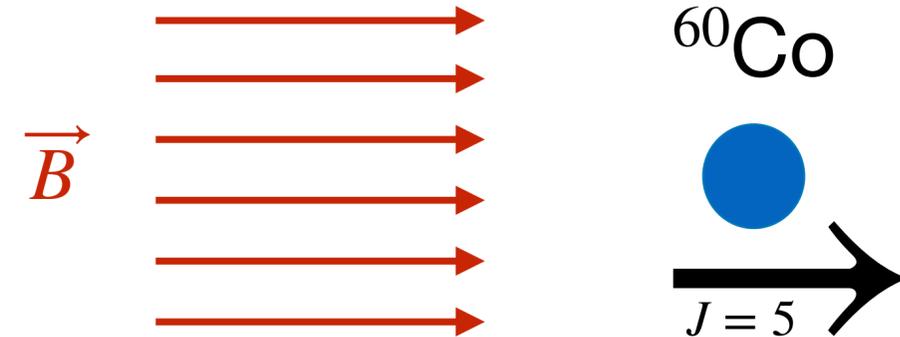
Experimental Test of Parity Conservation in Beta Decay*
 C. S. WU, *Columbia University, New York, New York*
 AND
 E. AMBLER, R. W. HAYWARD, D. D. HOPPES, AND R. P. HUDSON,
National Bureau of Standards, Washington, D. C.
 (Received January 15, 1957)



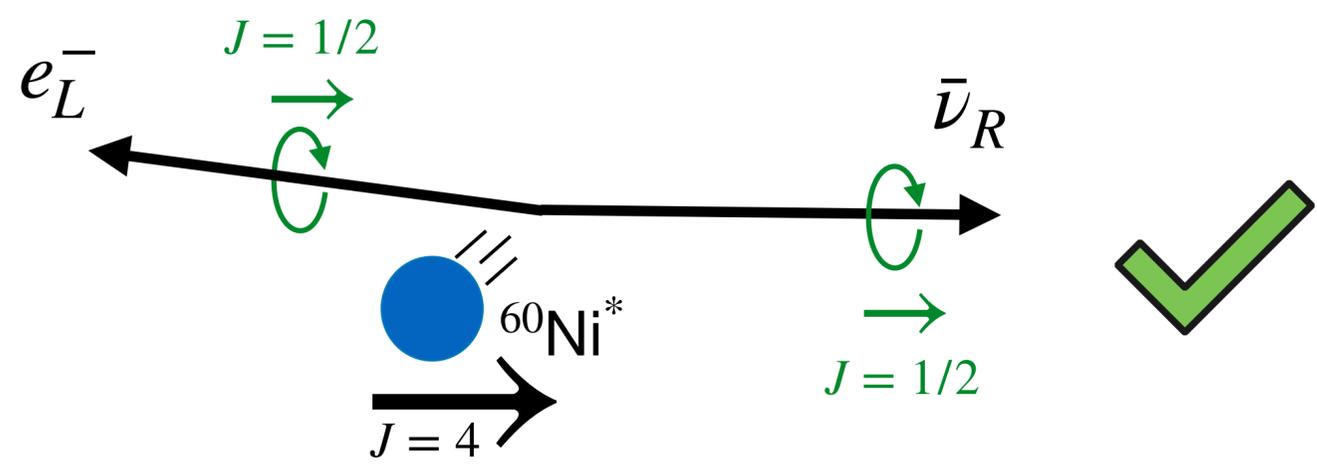
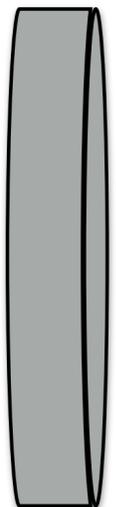
C.S. Wu, “Madame” Wu



$$(J = 5) \longrightarrow \left(J = 4 + \frac{1}{2} + \frac{1}{2} \right)$$



Detector



Weak Interactions are “Left-Handed”

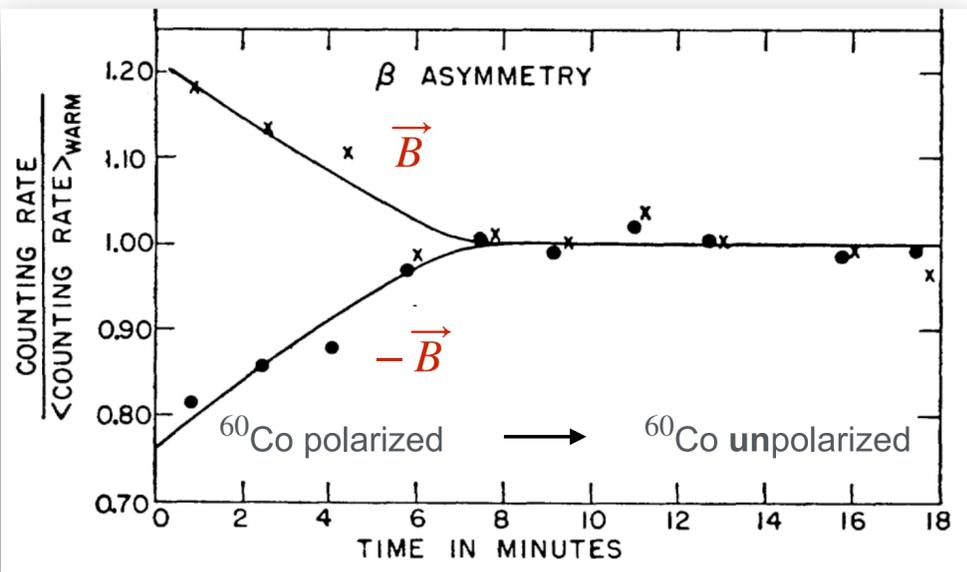
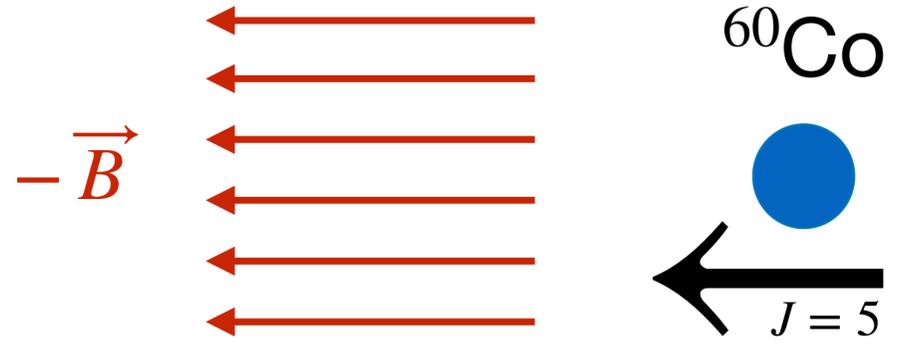
Experimental Test of Parity Conservation in Beta Decay*
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National Bureau of Standards, Washington, D. C.
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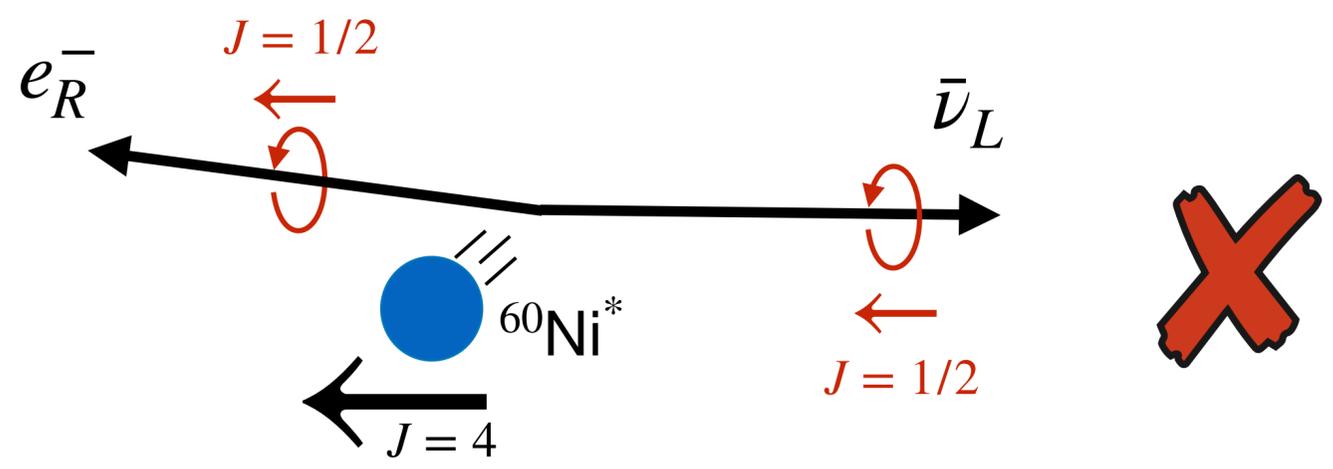
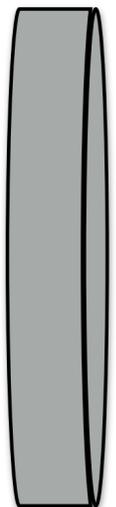
C.S. Wu, “Madame” Wu



$$(J = 5) \longrightarrow \left(J = 4 + \frac{1}{2} + \frac{1}{2} \right)$$

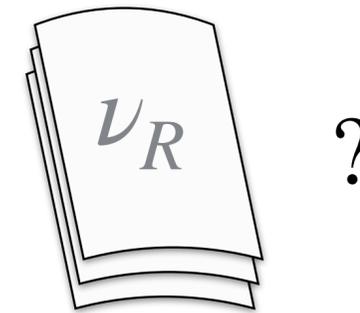
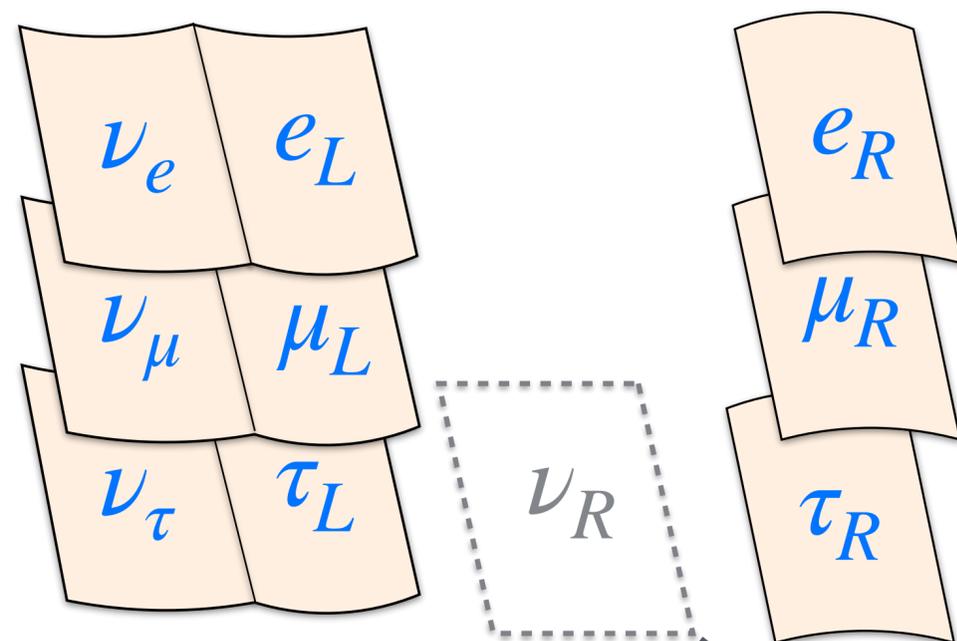


Detector



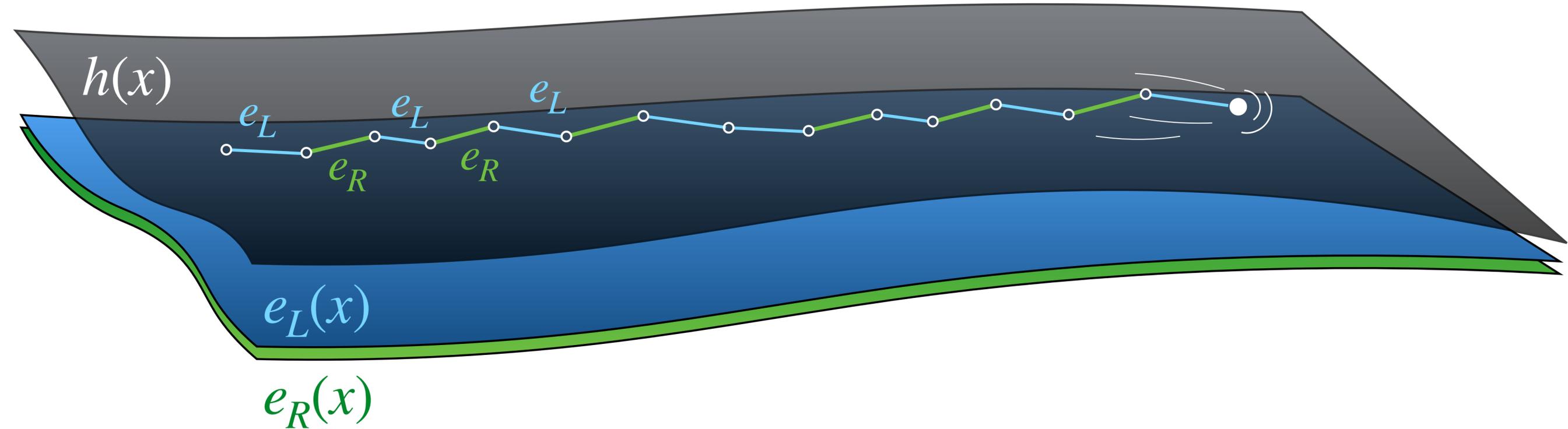
How many neutrinos?

“Sterile” or “right-handed” neutrinos.



If they exist, ν_R neutrinos would not interact with any known force.

Cartoonish picture fermion masses from the Higgs field



Electron-Higgs **interactions** give rise to the electron mass.

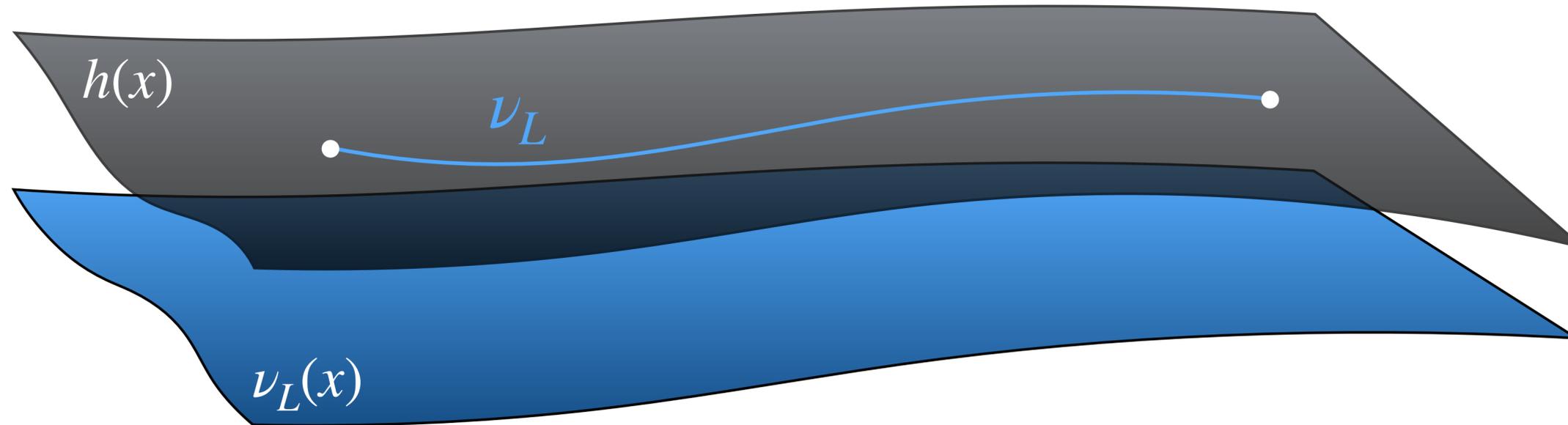
“ e_L ” and “ e_R ” are two sides of the same coin:

The Dirac fermion “ e ”

$$\begin{array}{c} \langle h \rangle \\ \vdots \\ e_L \text{ --- } m_e \text{ --- } e_R \end{array}$$



The problem with neutrinos



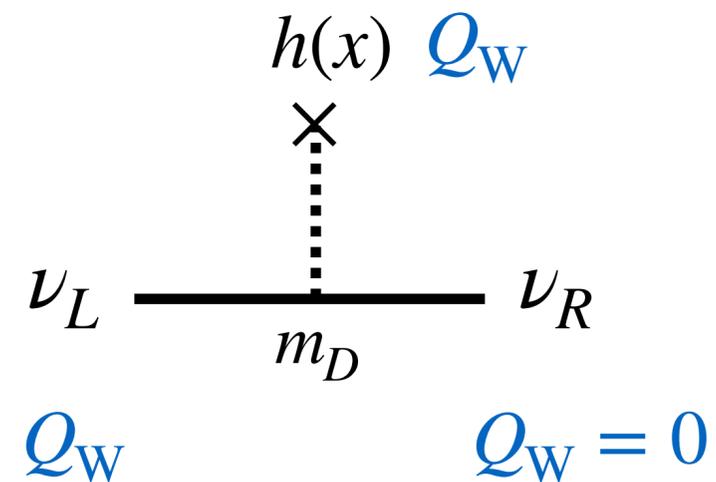
Without ν_R , the Standard Model predicts massless neutrinos.

As you know, this is **not** compatible with the fact that neutrinos oscillate!

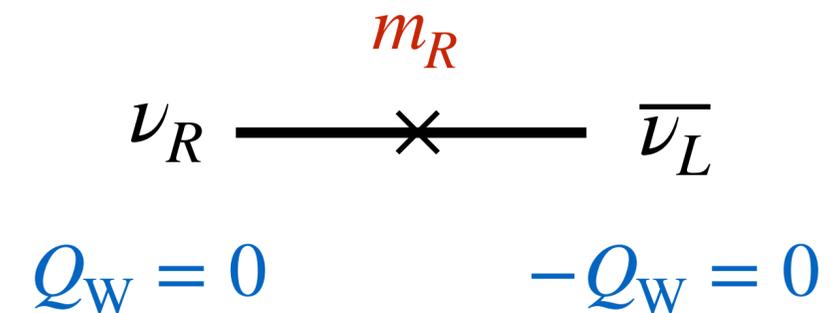


The Ambiguity of Neutrino Masses

Dirac Mass



Majorana Masses



Neutrinos carry no charge, so **Majorana masses are allowed.**

The only fundamental fermion we know of that can do this.

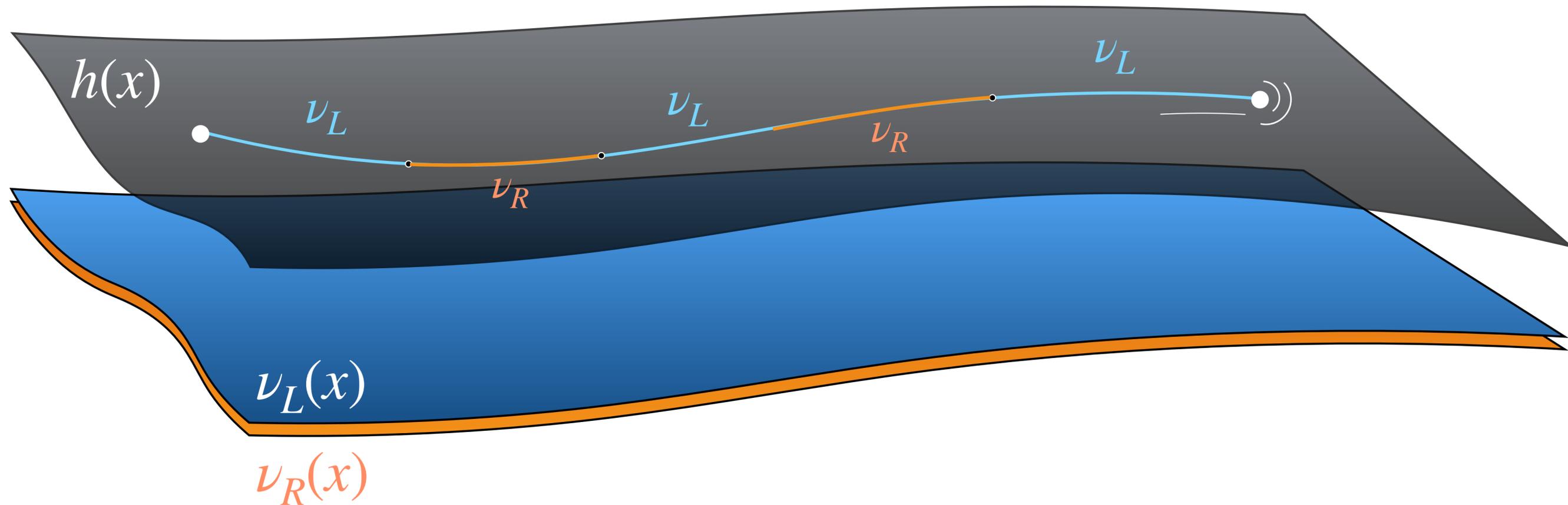
ν_R is “sterile” — it need not carry any (conserved) quantum number.



Dirac Neutrinos

Without a Majorana masses, ν_R is just part of light neutrinos

(“ ν_L and ν_R are two sides of the same coin”)



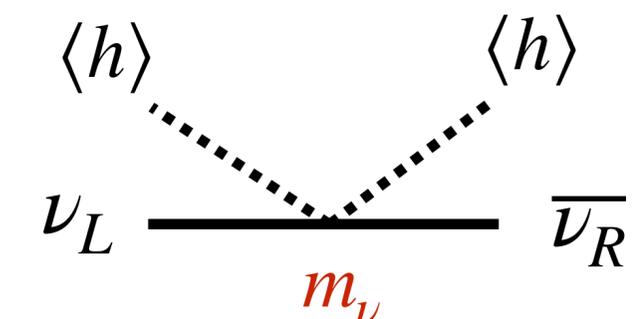
The seesaw mechanism

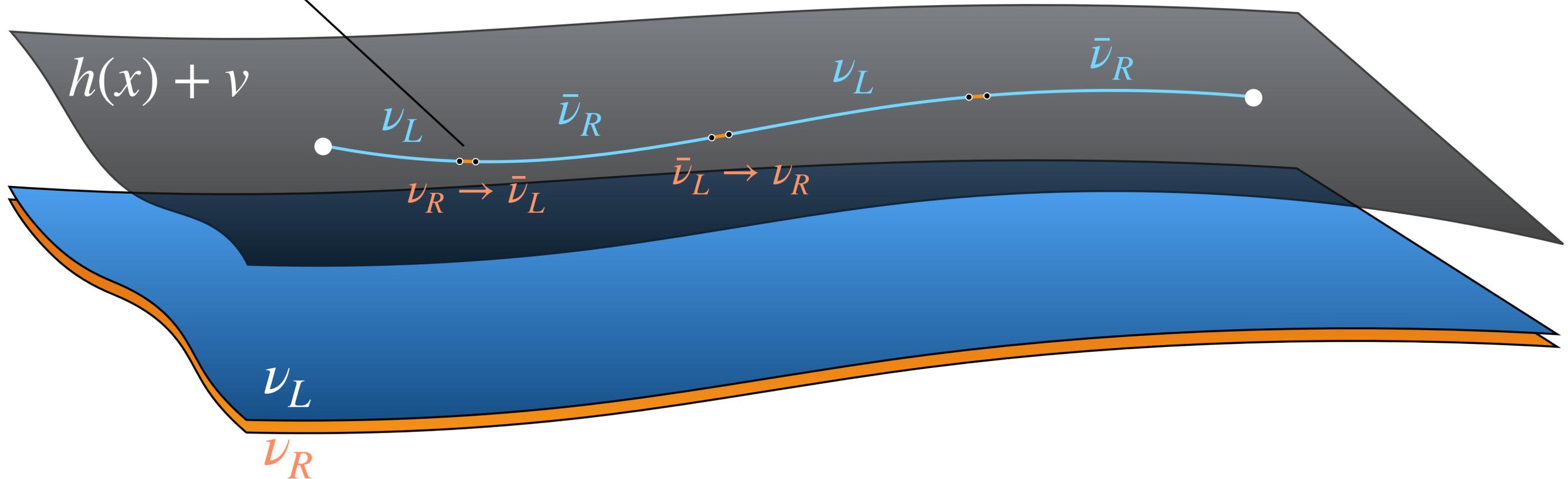
Light Majorana neutrinos become lighter the heavier the heavy Majorana neutrino becomes.

$$\delta t \delta E > \frac{\hbar}{2} \implies \delta t \sim \frac{1}{m_R}$$

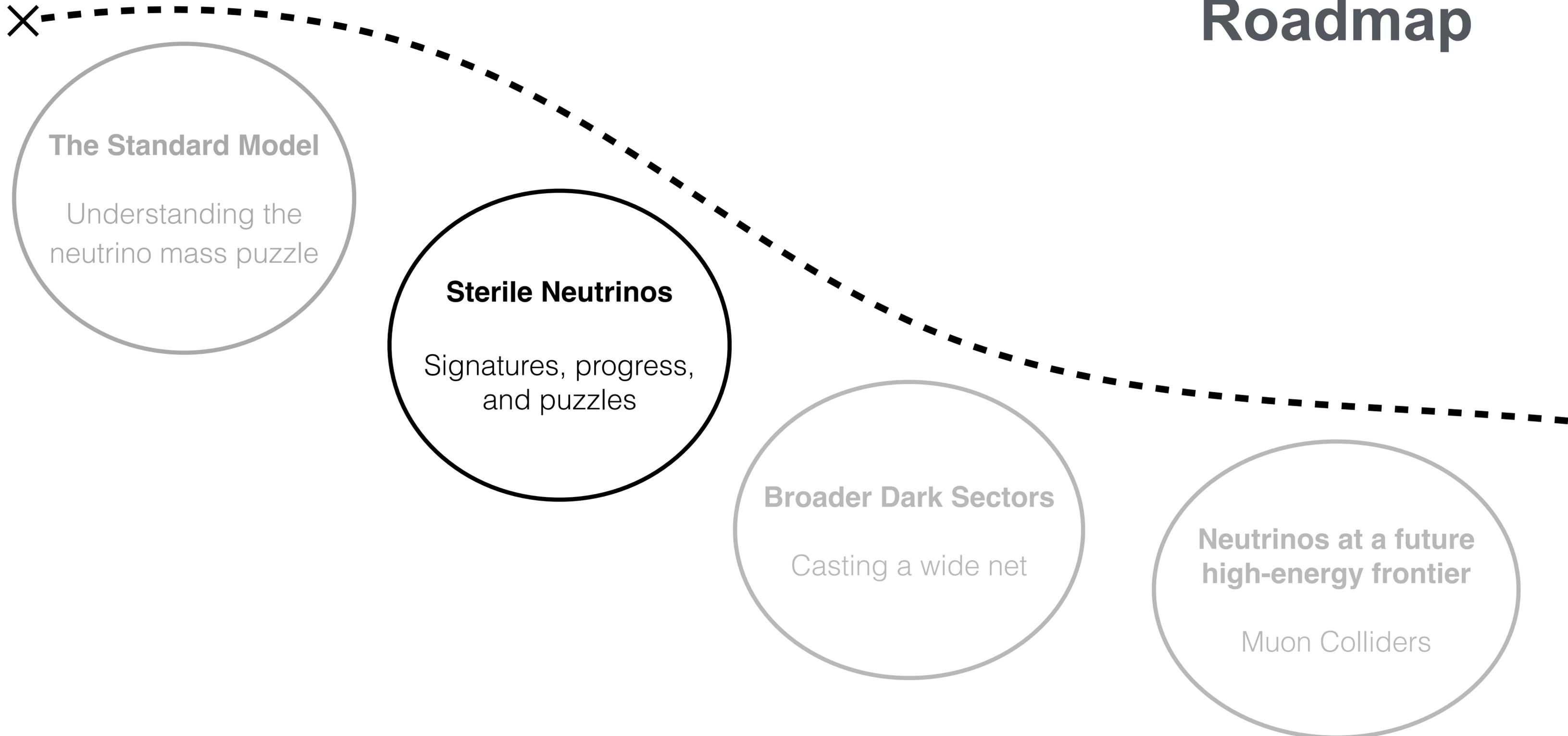
$$m_\nu \sim \frac{m_D^2}{m_R}$$

Effectively:



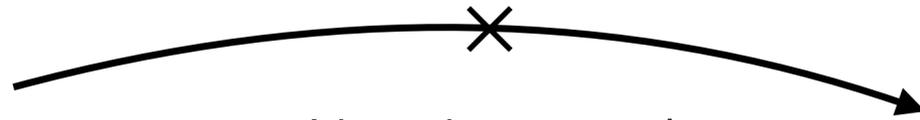


Roadmap



**STANDARD
MODEL**

ν θ N



Neutrino portal

HIDDEN SECTOR

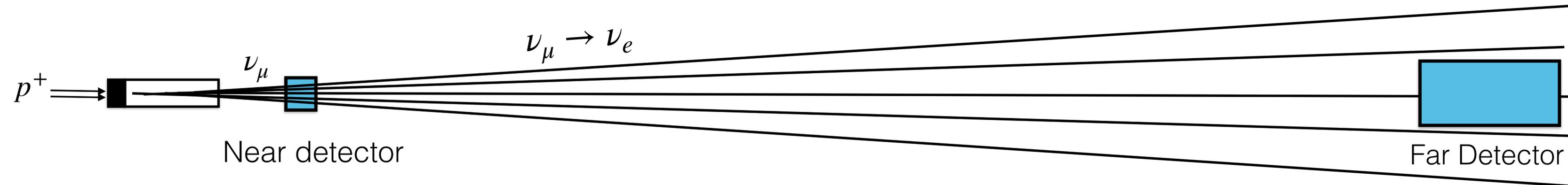
Heavy Neutrinos

New particles carry no SM charge.

New fundamental mass scales.

Light and feebly-coupled particles.

Accelerator Neutrino Experiments

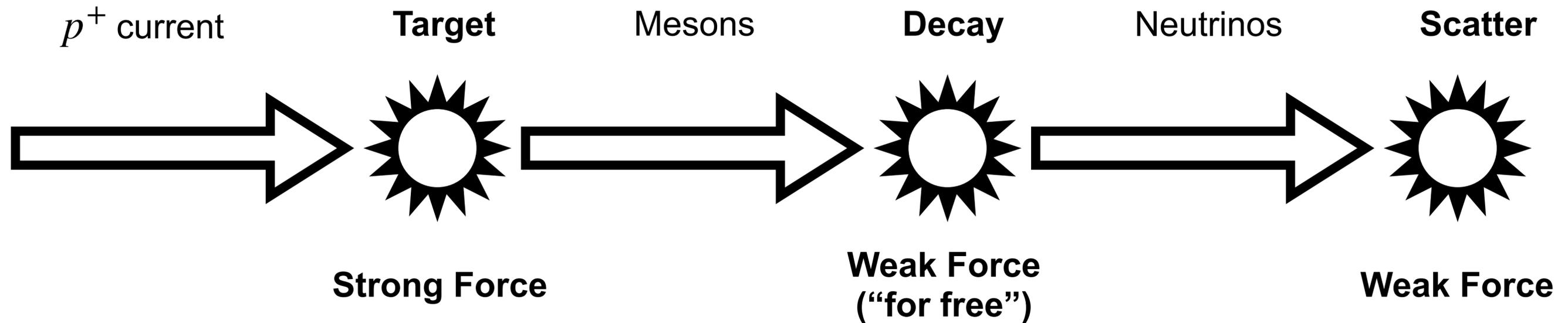
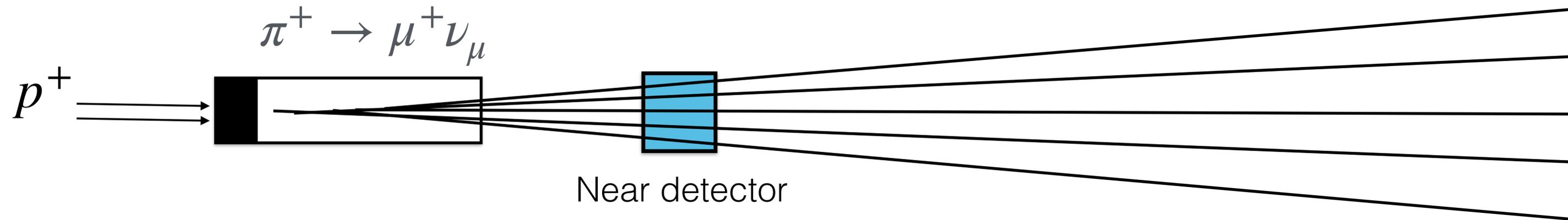


Standard Oscillations



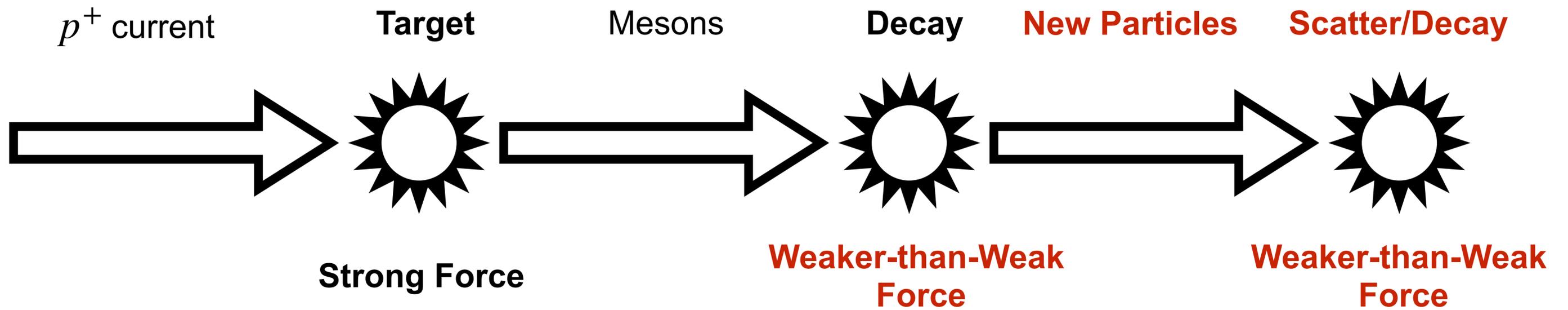
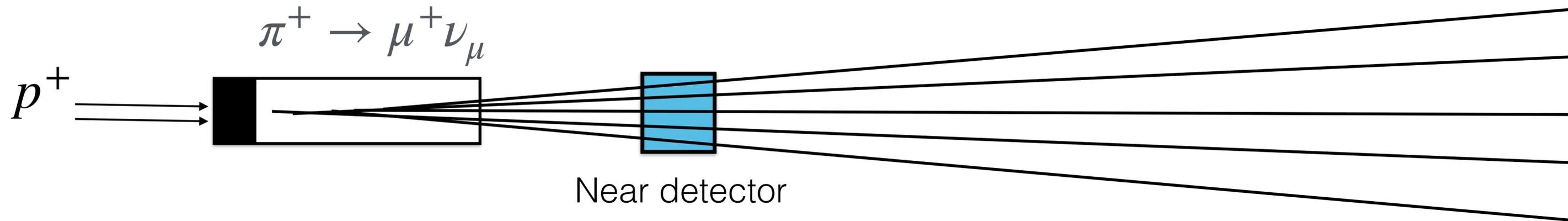
Accelerator Neutrino Experiments

Near detectors / short-baselines



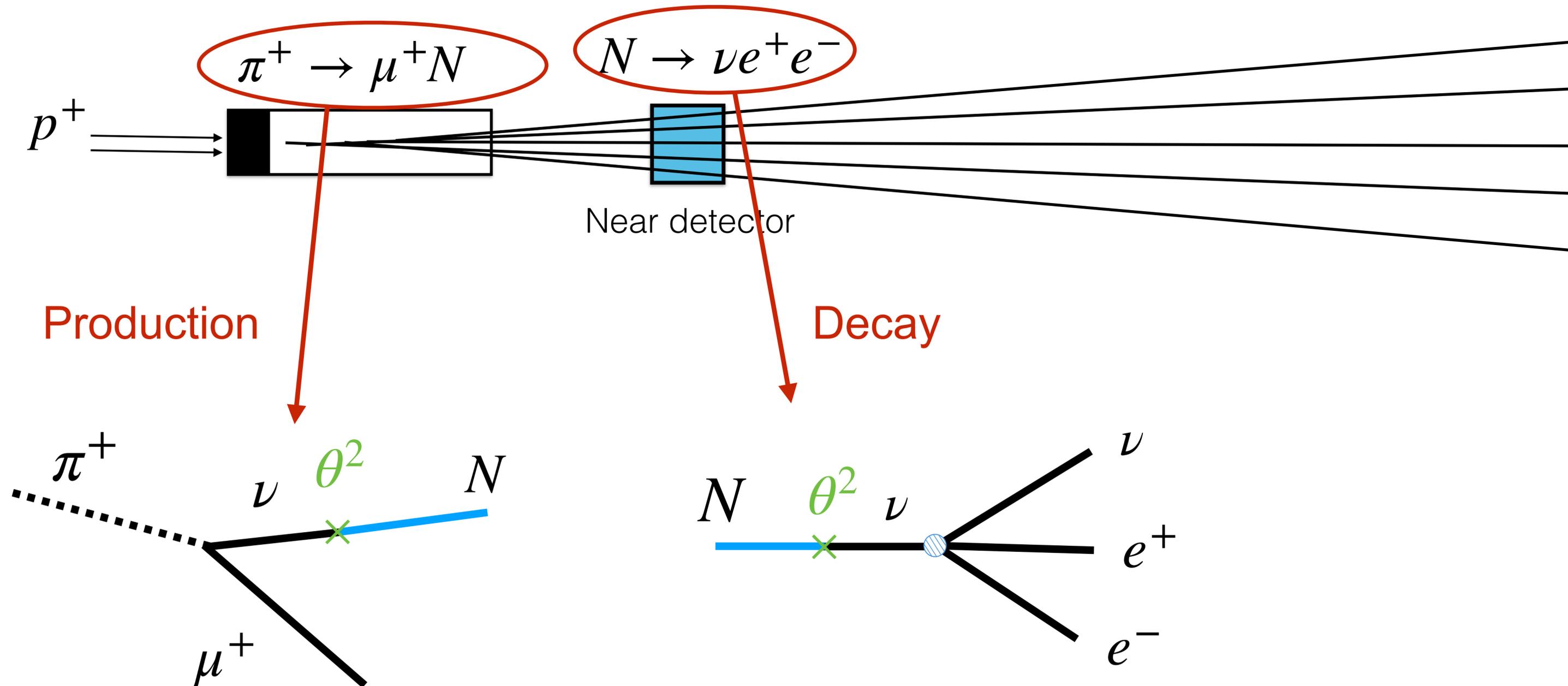
Accelerator Neutrino Experiments

Near detectors / short-baselines



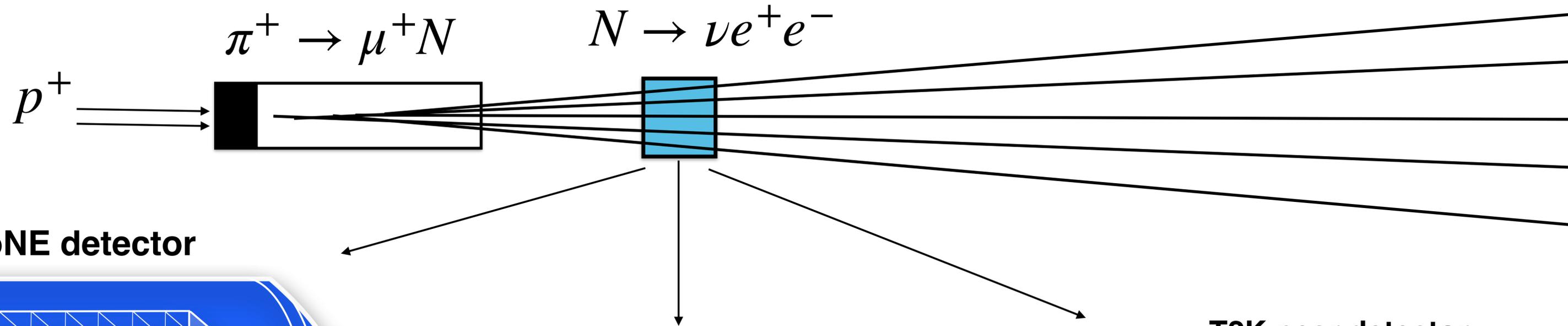
Accelerator Neutrino Experiments

Near detectors / short-baselines

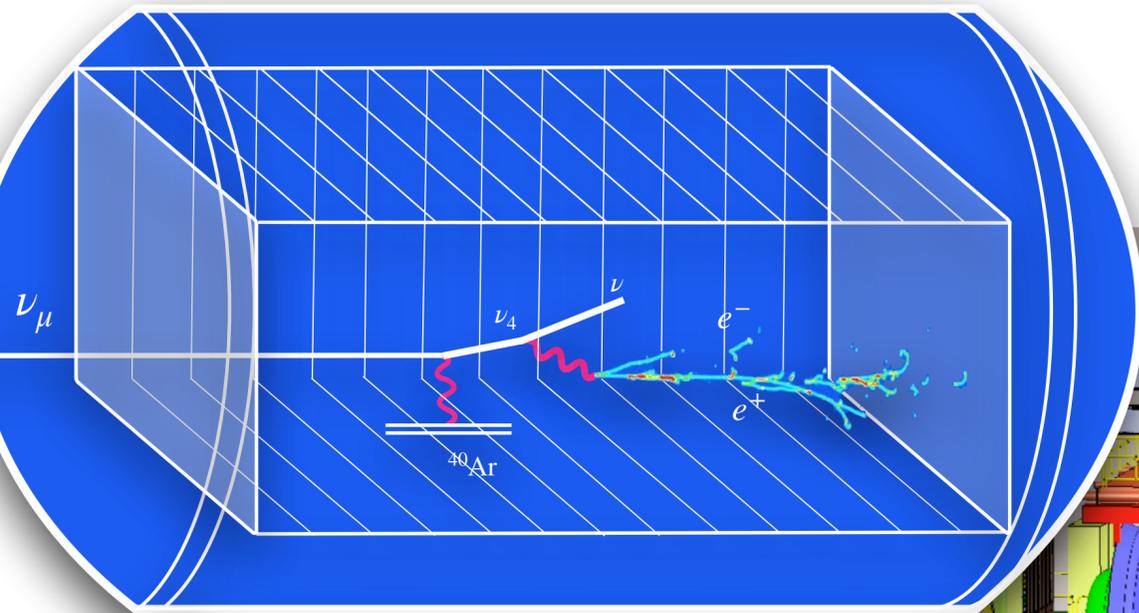


Accelerator Neutrino Experiments

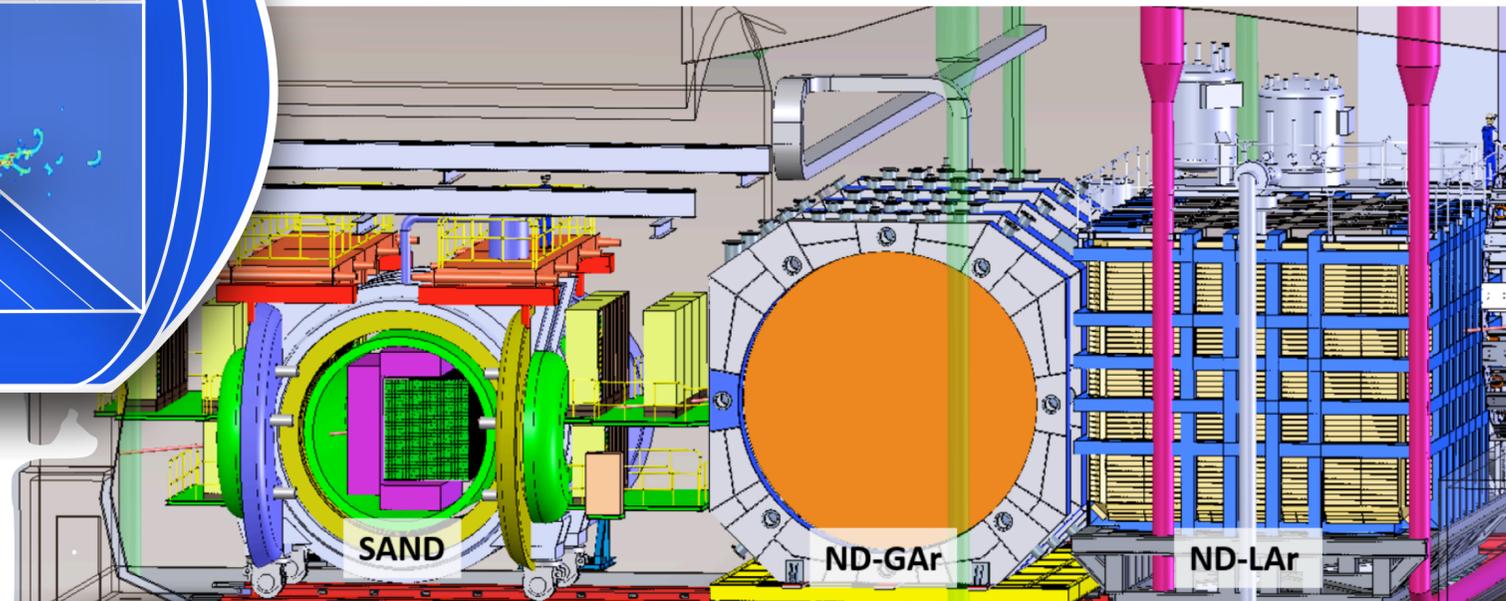
Near detectors / short-baselines



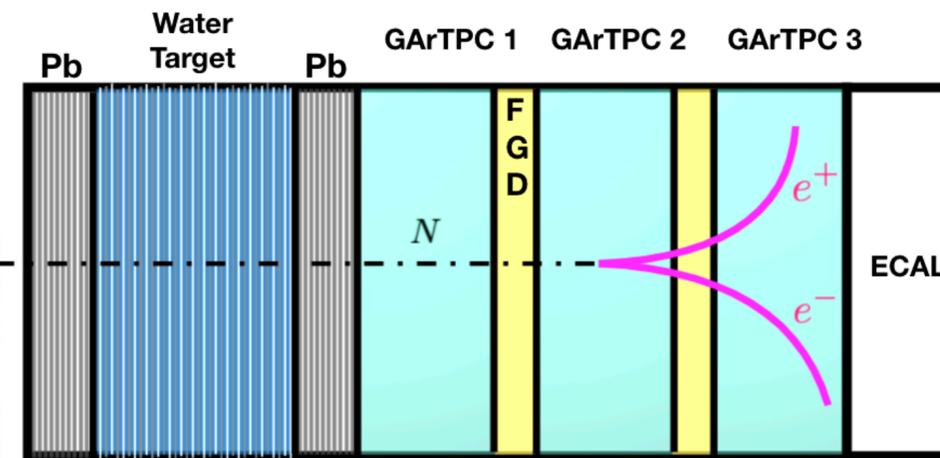
MicroBooNE detector



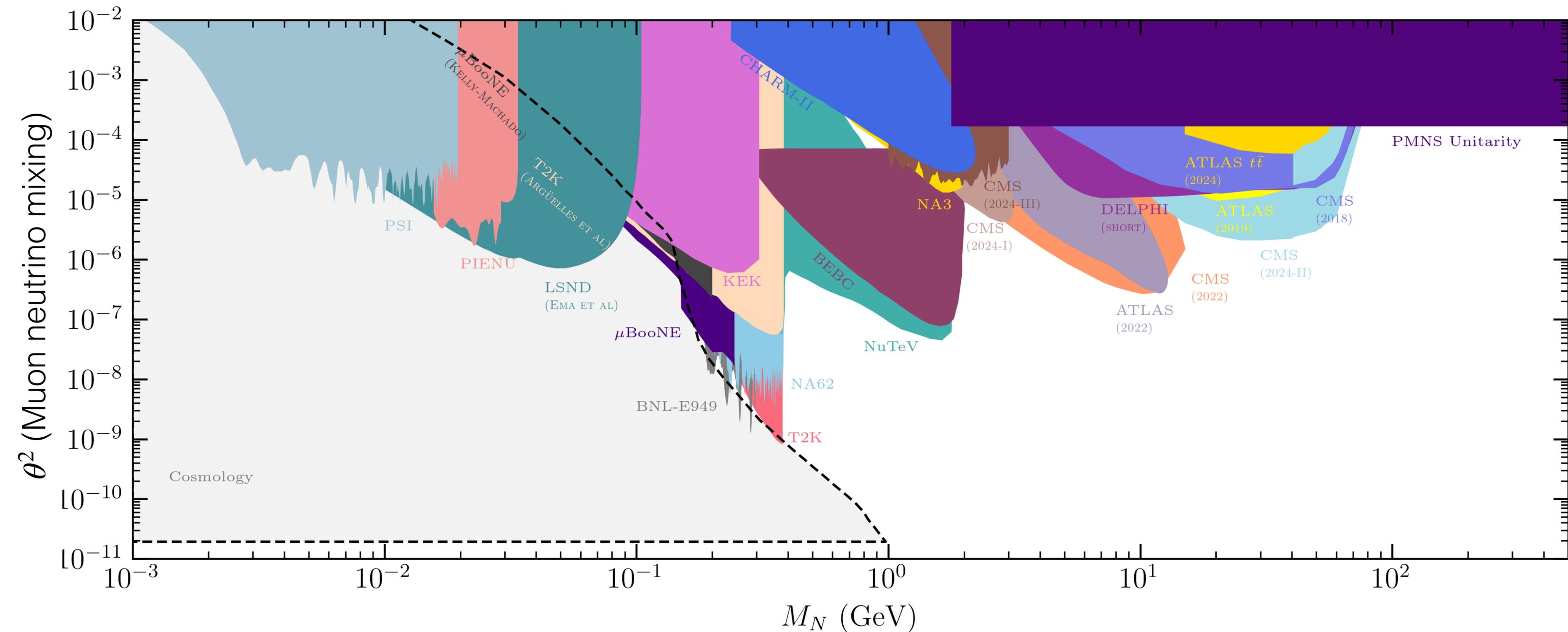
DUNE near detector complex



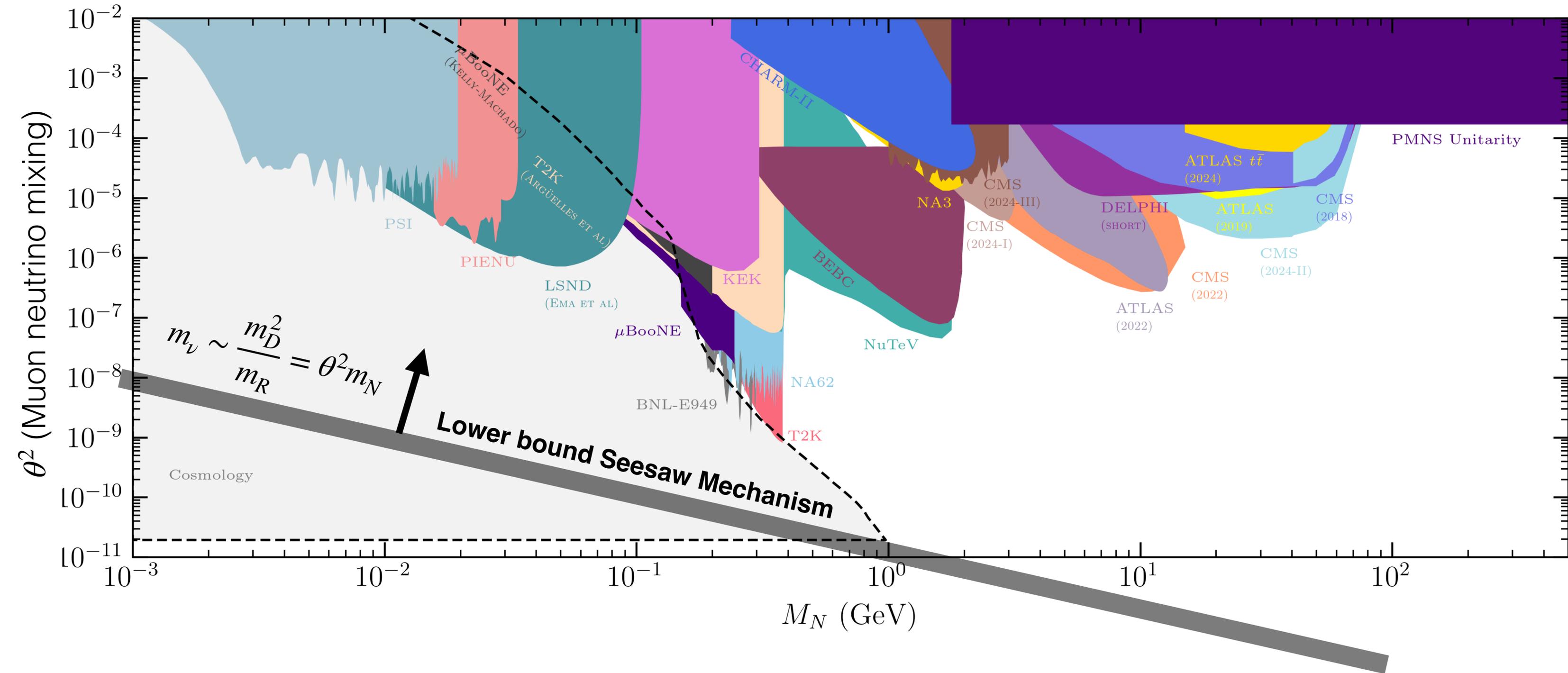
T2K near detector



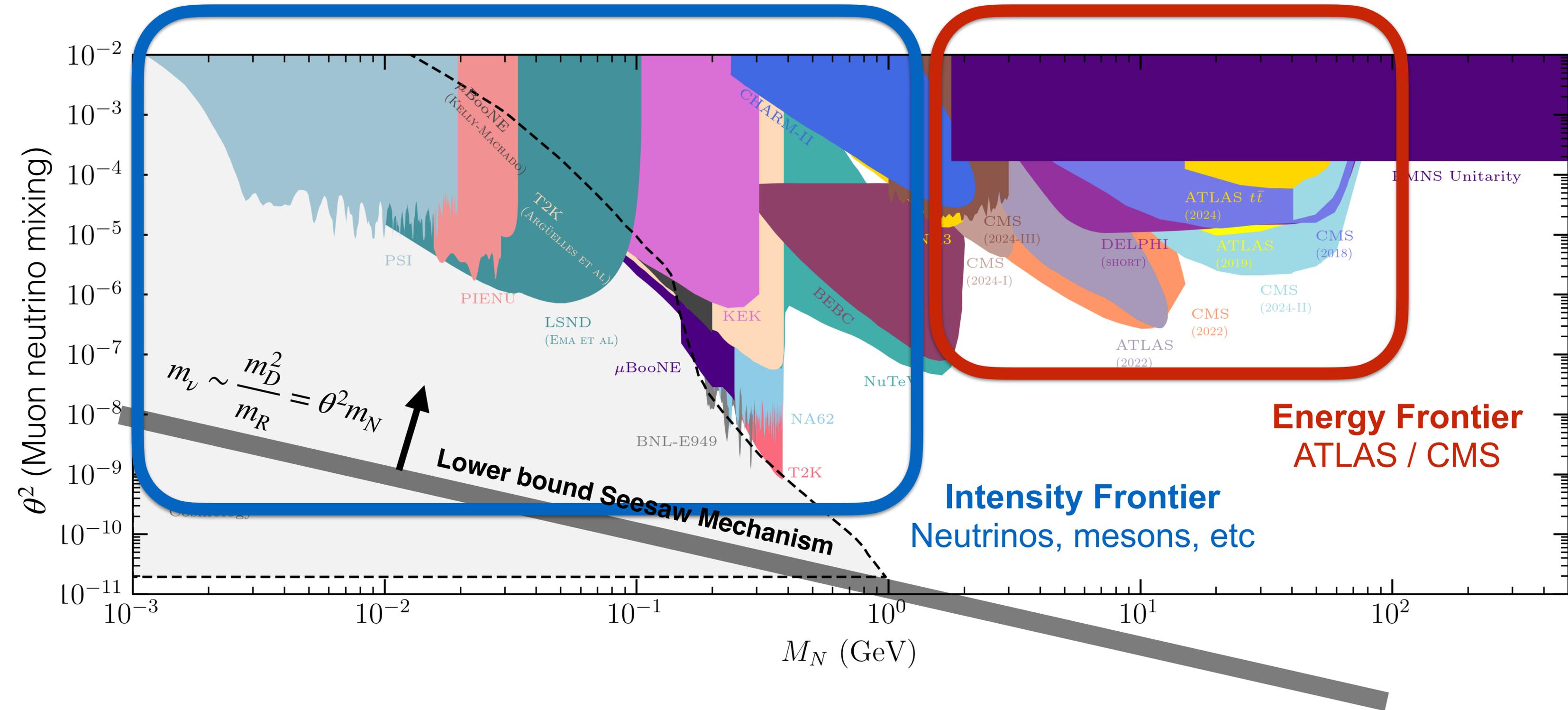
Parameter Space of Hypothetical Heavy Neutrinos



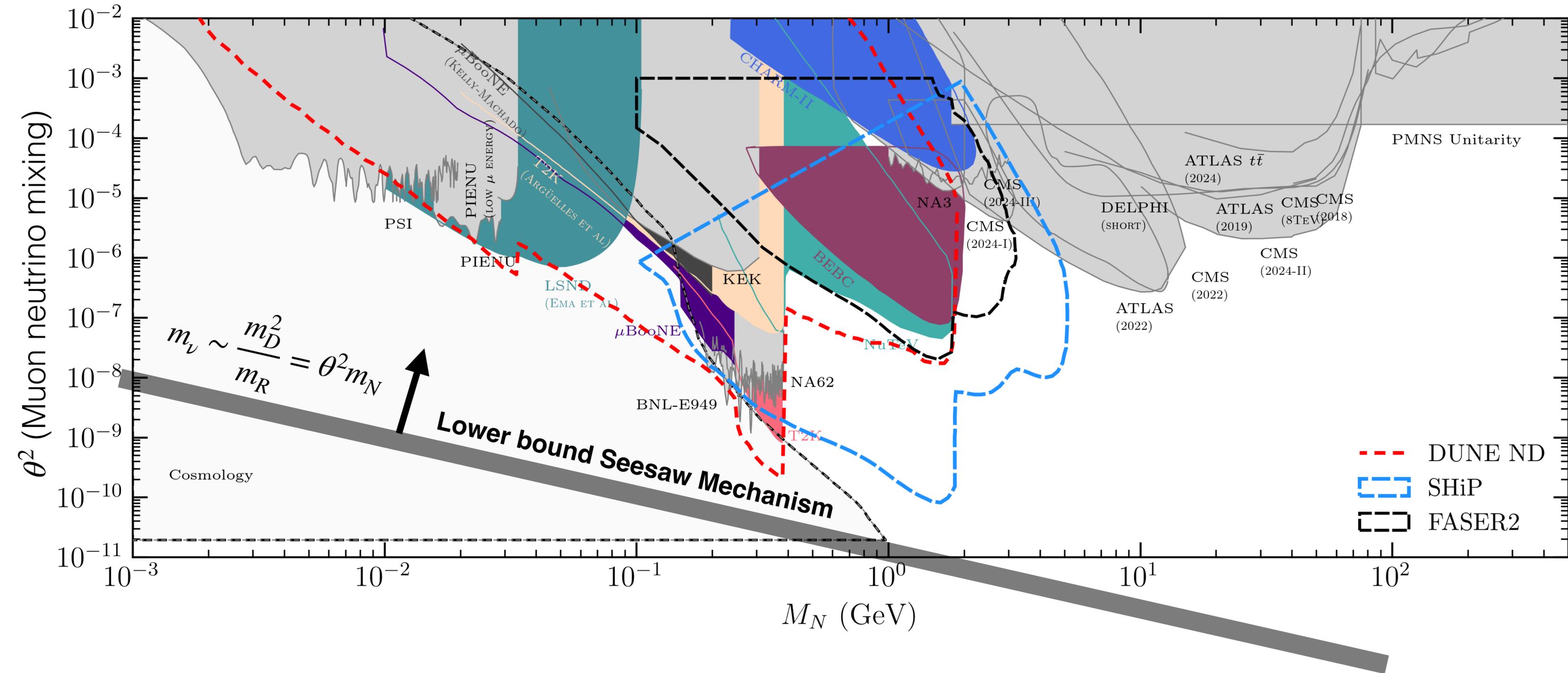
Parameter Space of Hypothetical Heavy Neutrinos



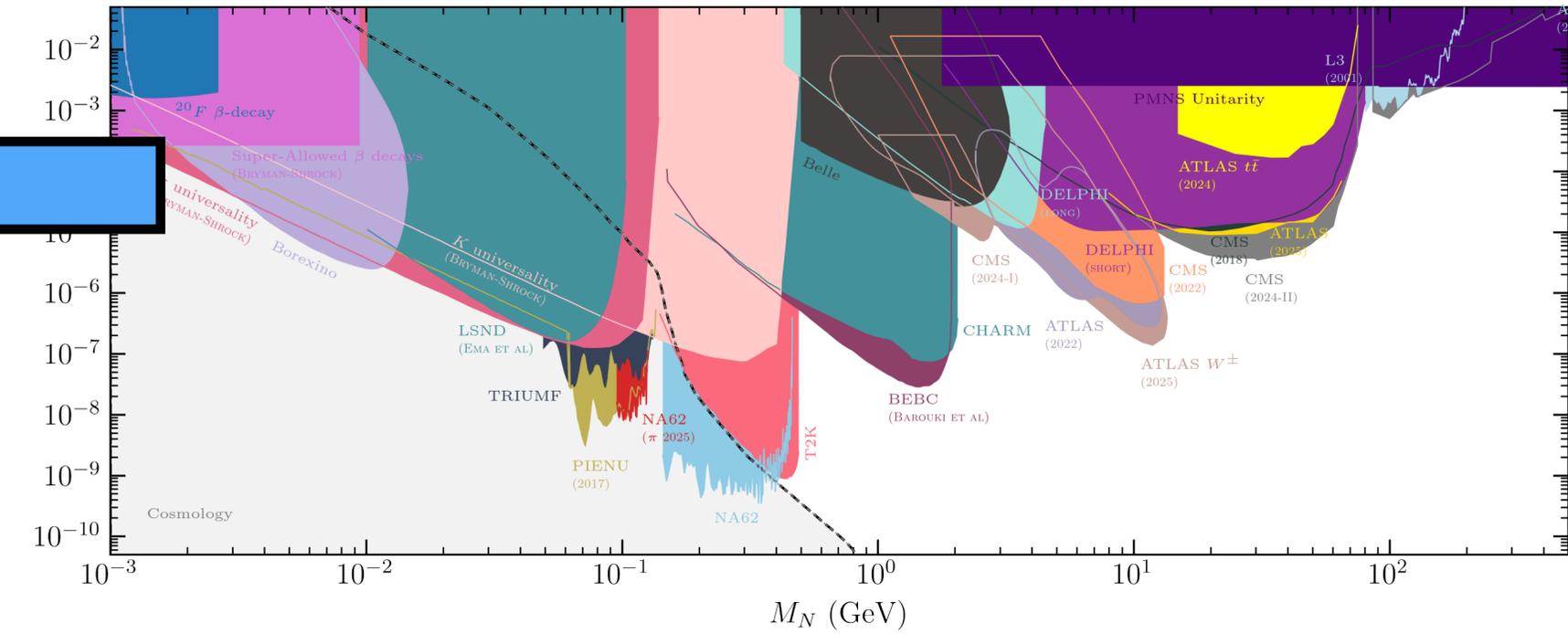
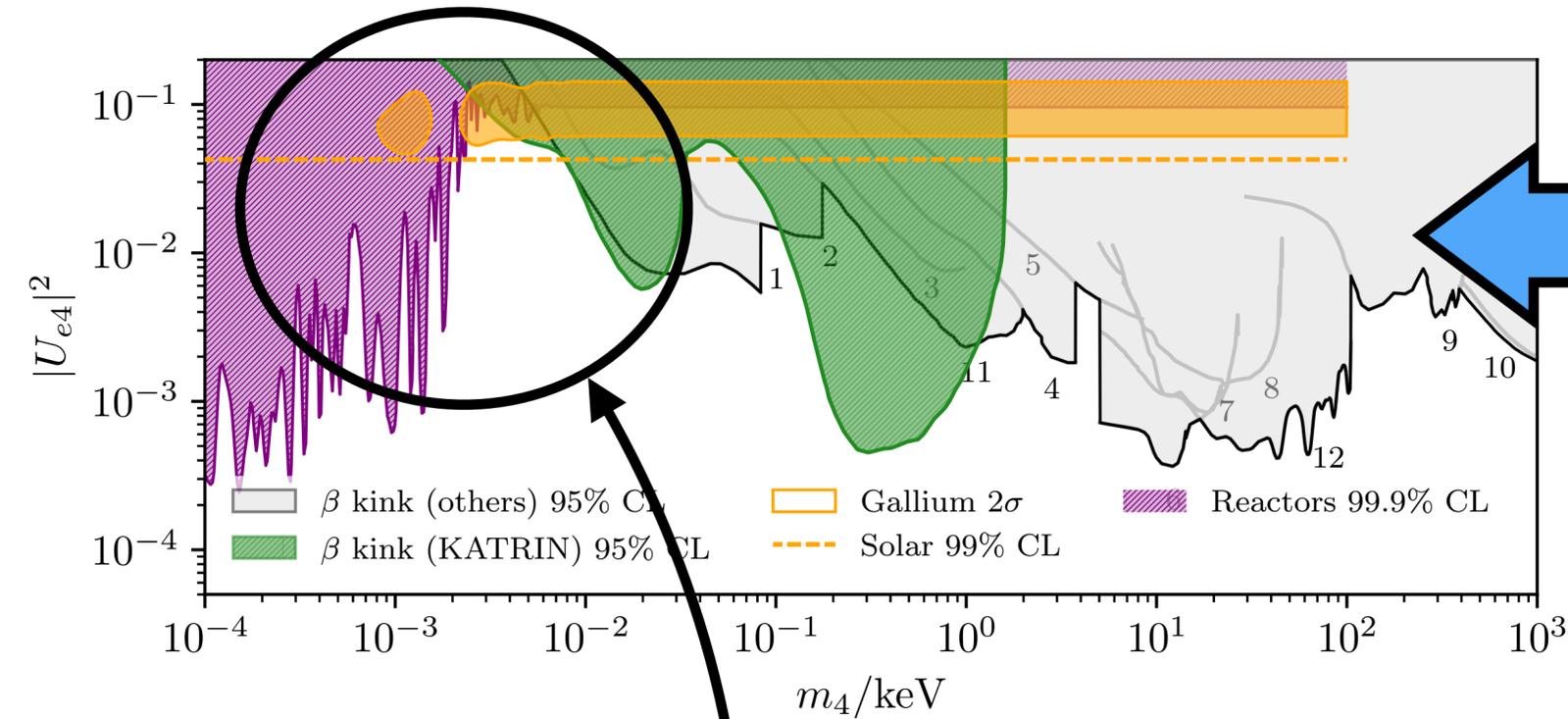
Parameter Space of Hypothetical Heavy Neutrinos



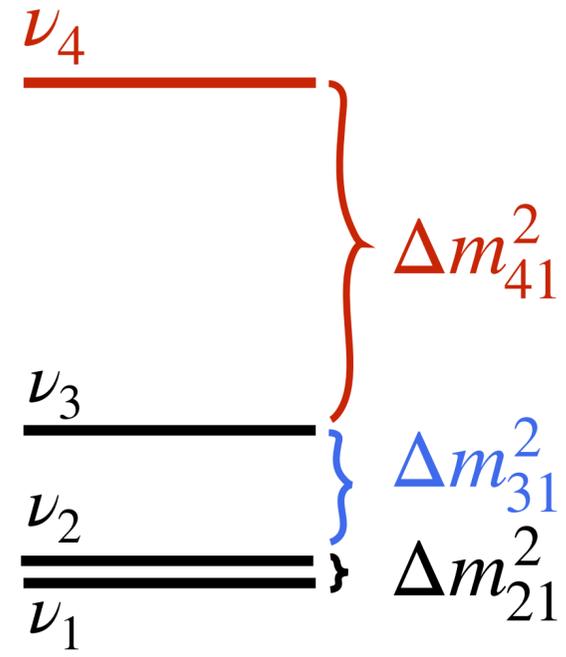
Parameter Space of Hypothetical Heavy Neutrinos



Parameter Space of Hypothetical Heavy Neutrinos



eV-scale sterile-neutrino driven oscillations
"3+1 model"



Article | [Open access](#) | Published: 03 December 2025

Search for light sterile neutrinos with two neutrino beams at MicroBooNE

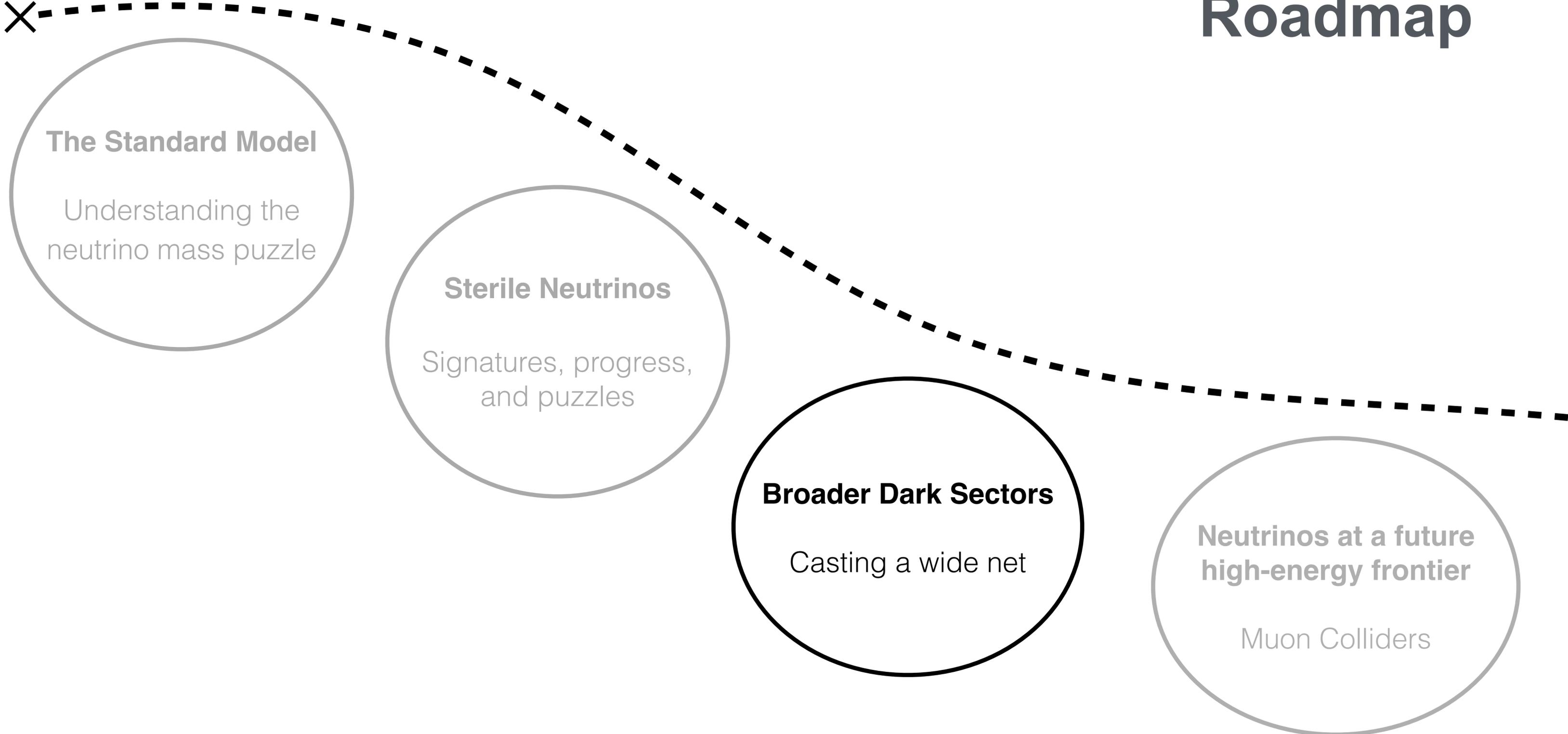
[The MicroBooNE Collaboration](#)

[Nature](#) 648, 64–69 (2025) | [Cite this article](#)

But see also:

O. Benevides Rodrigues, MH, K. J. Kelly,
B. Littlejohn, P. A. N. Machado, I. Safa, T. Zhou
[PRL 135, 081801](#)

Roadmap

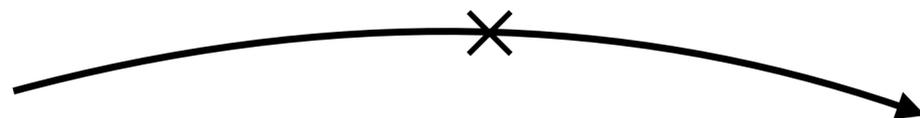


**STANDARD
MODEL**

DARK SECTOR

ν

N



Neutrino portal

Heavy Neutrinos

γ

γ_d



Photon portal

Dark photons

h

h_d



Higgs portal

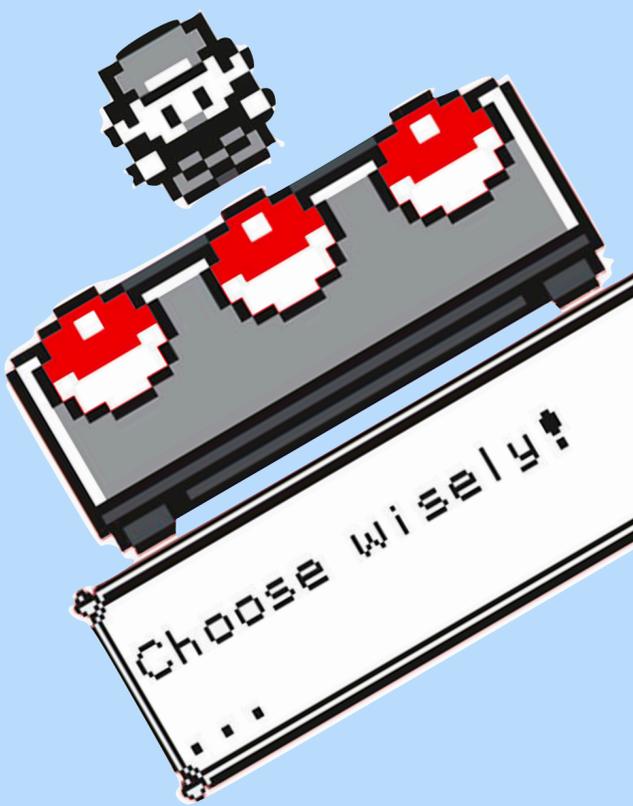
Dark Higgs



Axions and Axion-like-particles

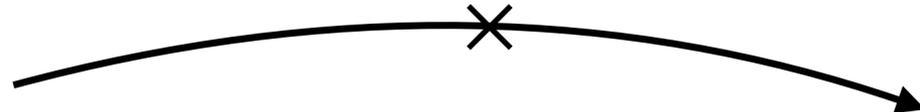
Axions (goldstone bosons)

STANDARD MODEL



ν

N



Neutrino portal

γ

γ_d



Photon portal

h

h_d



Higgs portal



Axions and Axion-like-particles

DARK SECTOR



Heavy Neutrinos



Dark photons



Dark Higgs



Axions (goldstone bosons)

Just the "starter" options.
"Bottom up" approach.



dark matter?

Dark Neutrinos

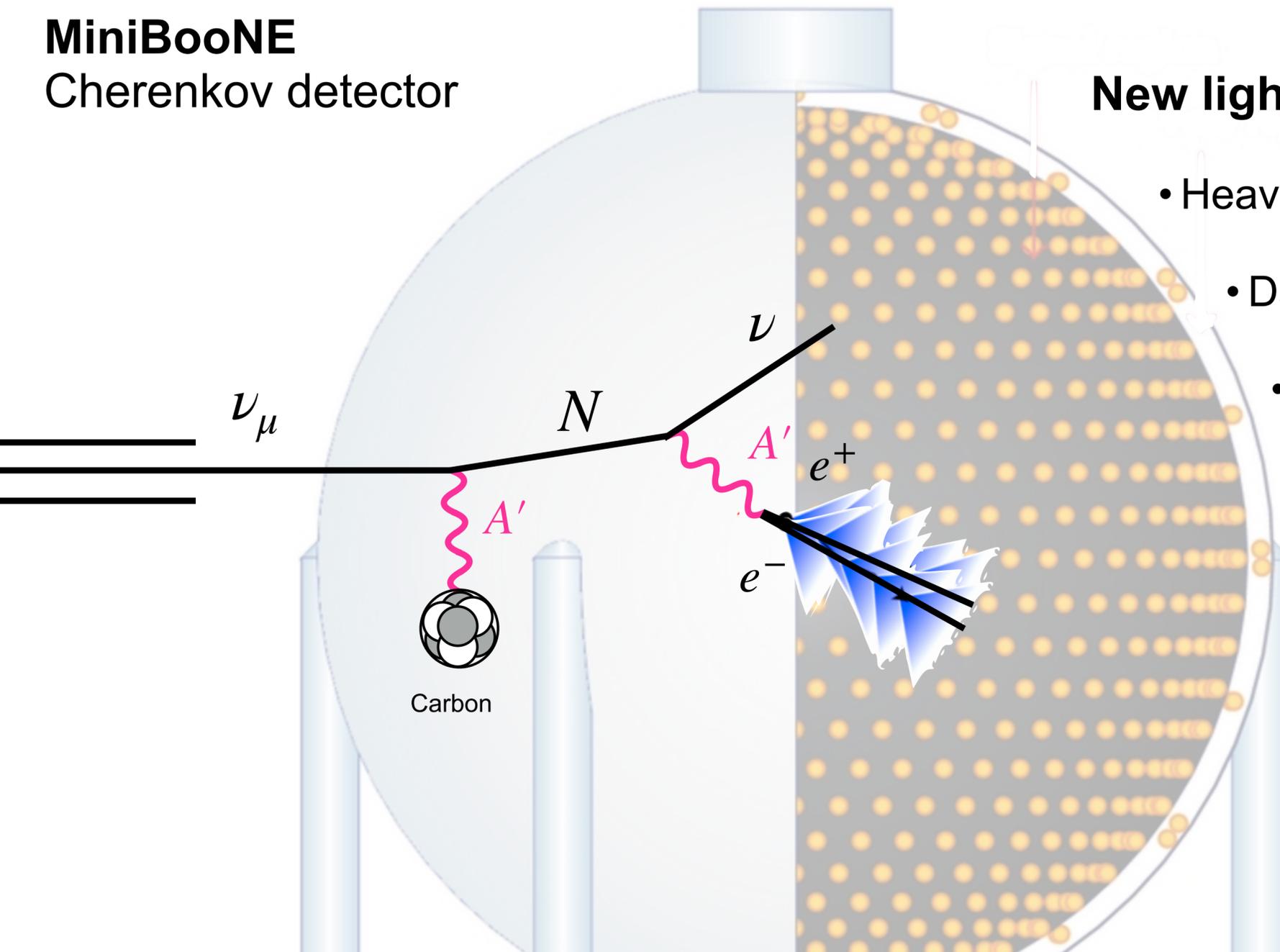
A multi-portal exploration of dark sectors

E. Bertuzzo, S. Jana, P. A. N. Machado, R. Z. Funchal, [PRL 121, 241801 \(2018\)](#)
 P. Ballett, M. Ross-Lonergan, S. Pascoli, [PRD 99, 071701 \(2019\)](#)

C. Argüelles, MH, Y. Tsai, [PRL 123, 261801 \(2019\)](#)
 P. Ballett, MH, S. Pascoli, [PRD 101, 115025 \(2020\)](#)
 A. Abdullahi, MH, S. Pascoli, [PLB 820 136531 \(2021\)](#)

MiniBooNE
 Cherenkov detector

MiniBooNE detector

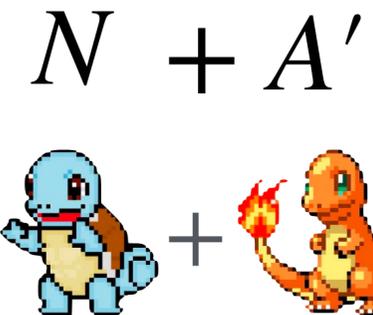


New light particles produced in neutrino interactions:

- Heavy Neutrinos?
- Dark Photons?
- Dark Higgses?

Multi-portal approach

Heavy neutrinos interacting through dark photons:



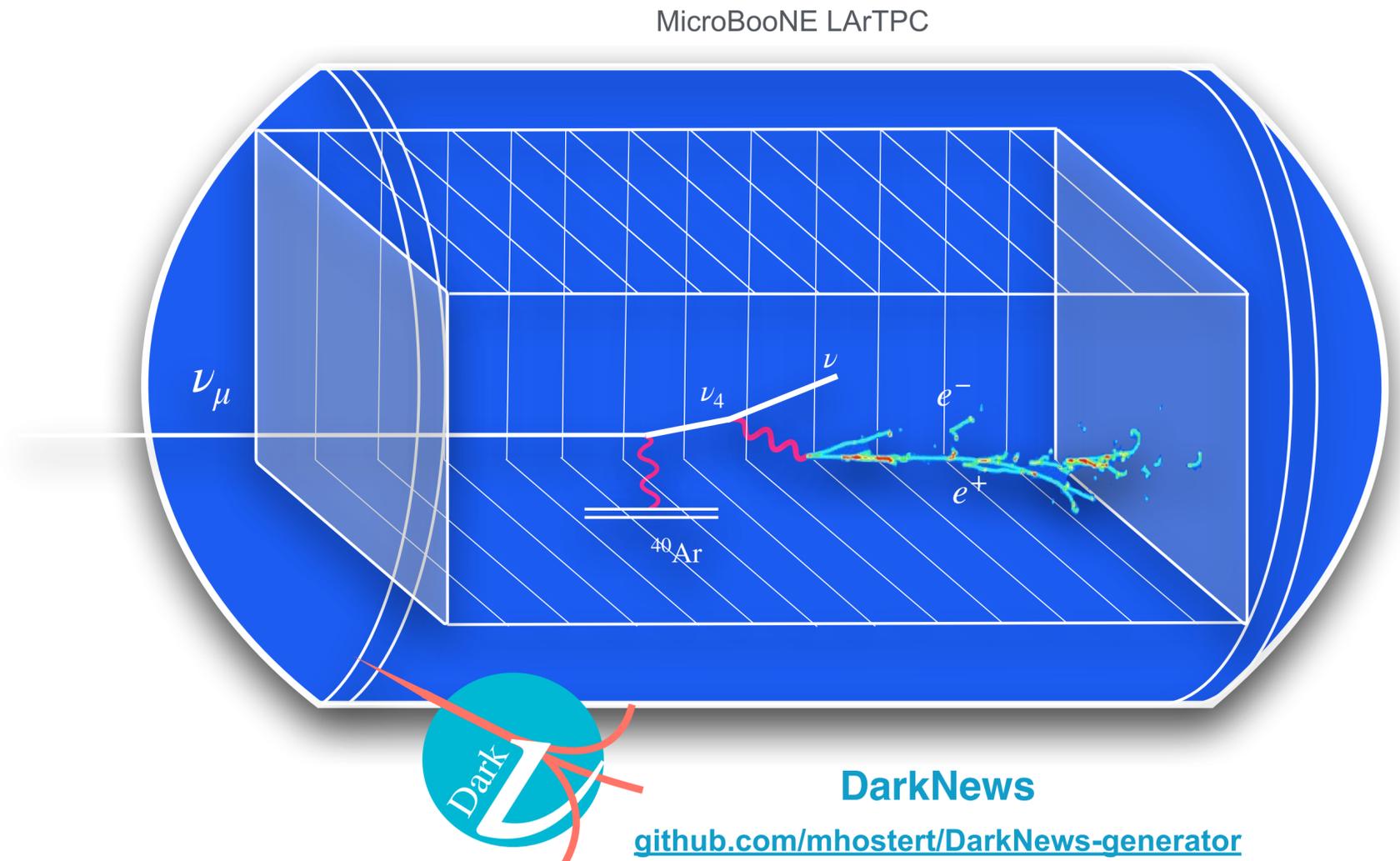
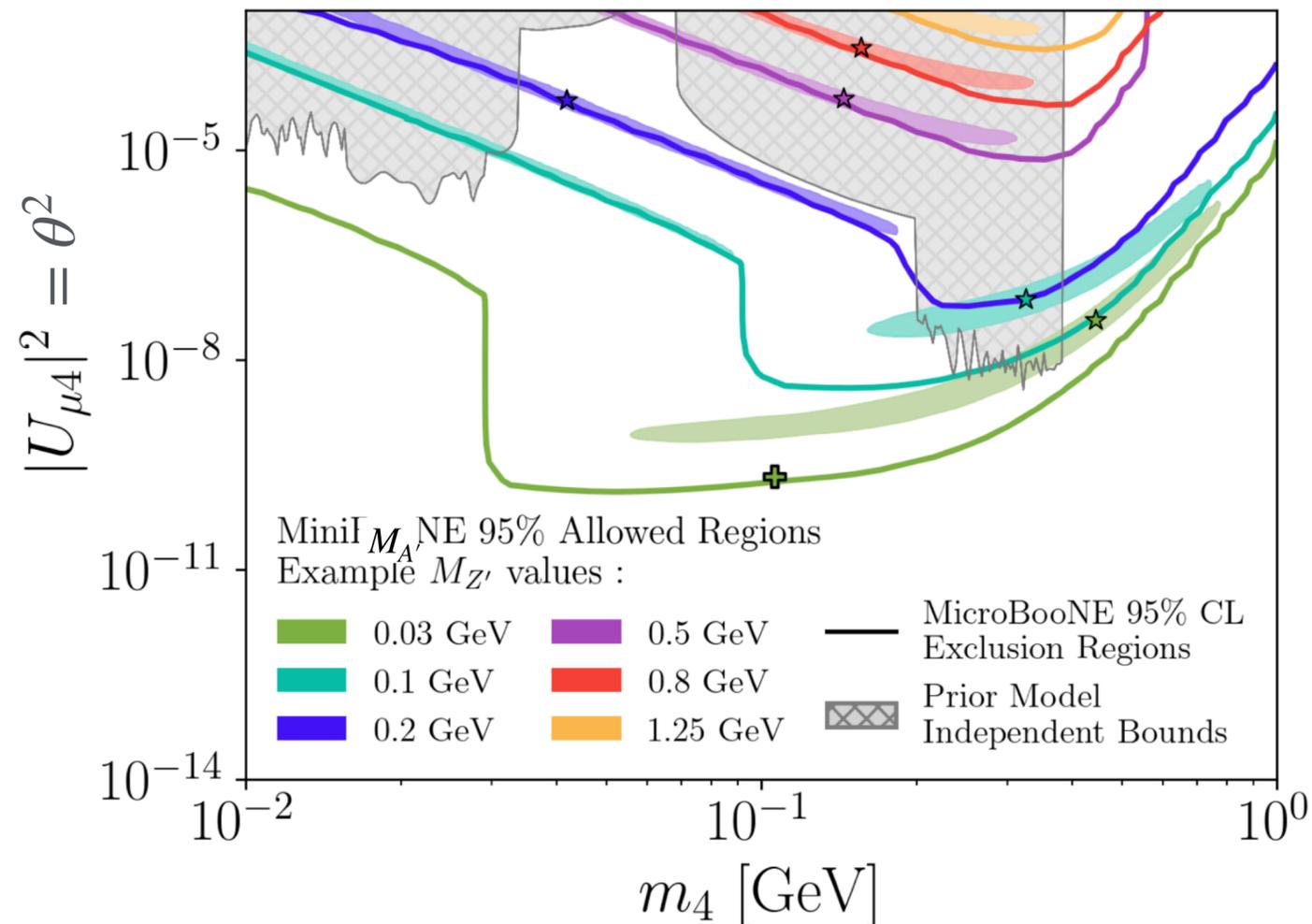
Dark Neutrinos @ MicroBooNE

First (modern) experimental limits on neutrino upscattering

MicroBooNE coll.
+
MOU with A. Abdullahi, MH, D. Massaro,
S. Pascoli, J. Zink.
arxiv.org/abs/2502.10900

First dedicated experimental search for neutrino-induced e^+e^- pairs:

MicroBooNE excludes minimal dark neutrino explanations to the MiniBooNE anomaly



Another example: axion-like-particles

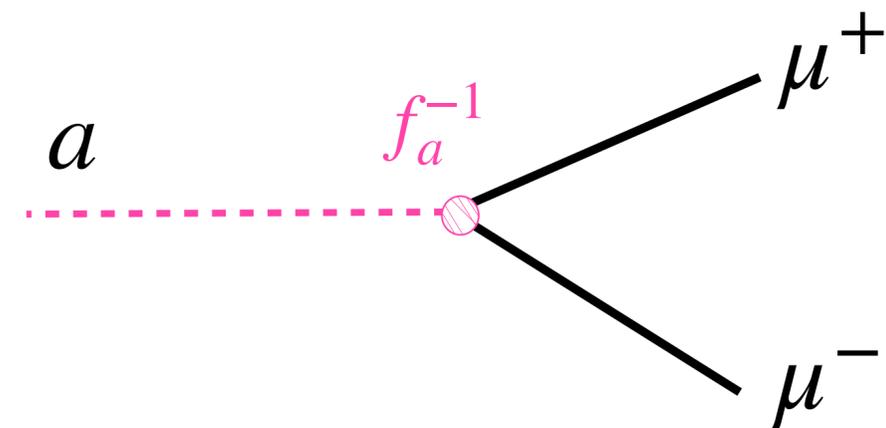
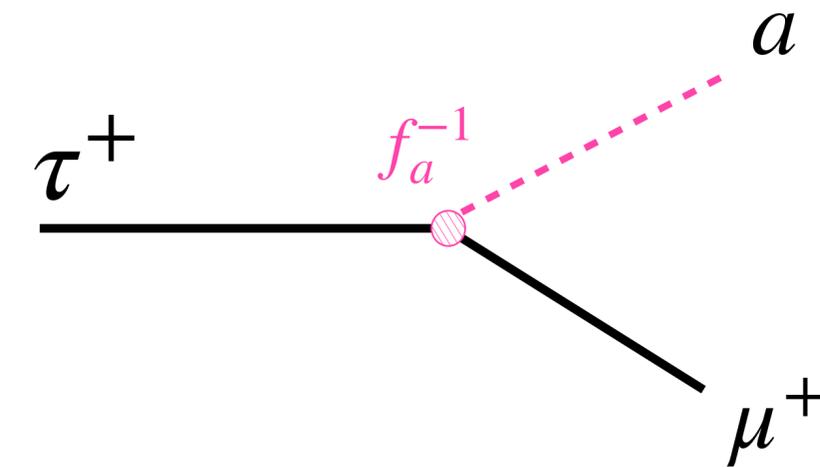
(pseudo-Goldstone bosons from, e.g., flavor symmetries)

$$\mathcal{L}_{\text{int}} = \frac{\partial_\mu a_\ell}{f_a} \sum_{i,j} c_{\ell_i \ell_j} \left[\bar{\ell}_i \gamma^\mu \gamma^5 \ell_j + \bar{\nu}_{\ell_i} \gamma^\mu \gamma^5 \nu_{\ell_j} \right]$$

Angular mode of some complex scalar.

Derivatively-coupled \rightarrow couples preferentially to 3rd family

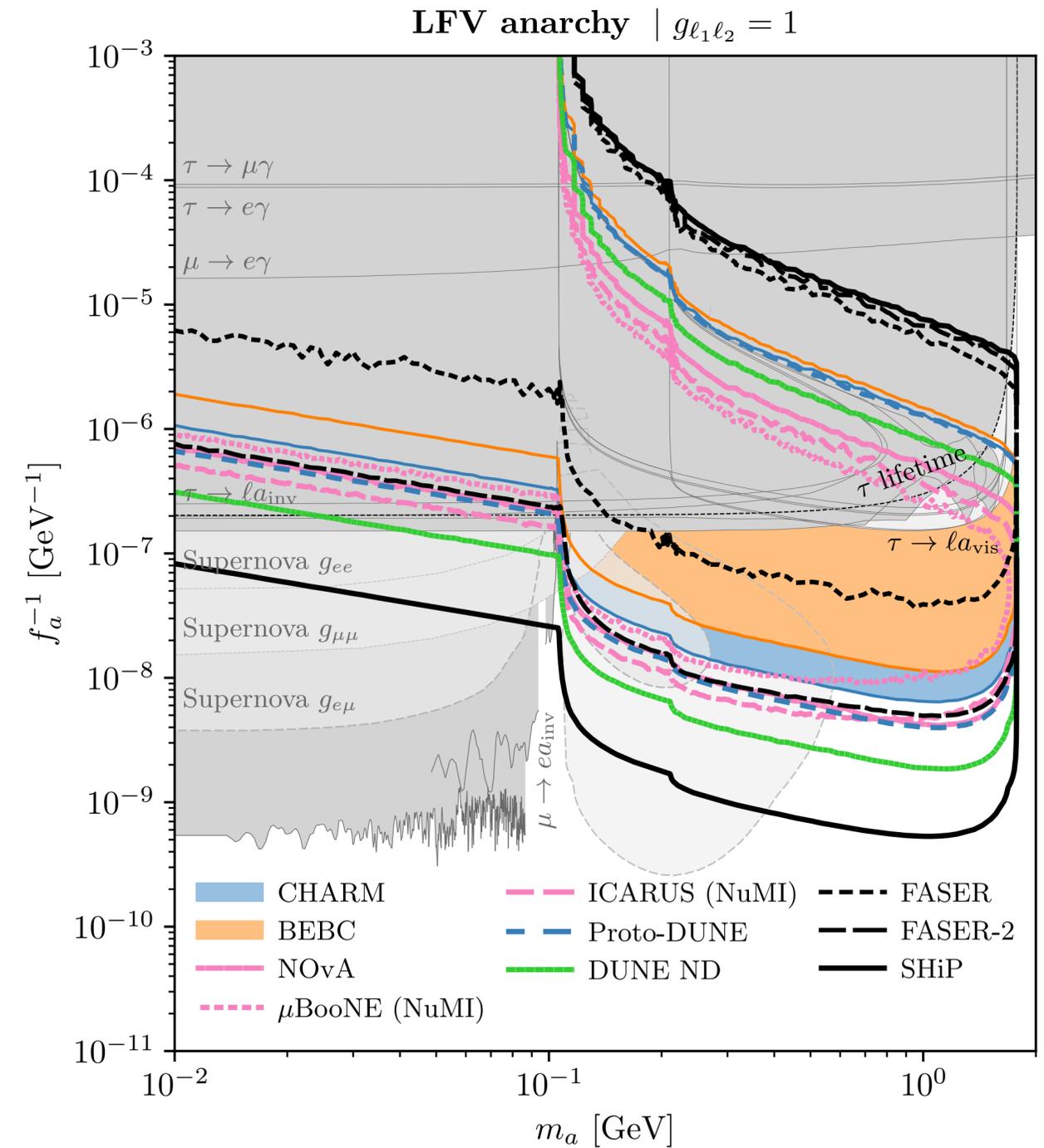
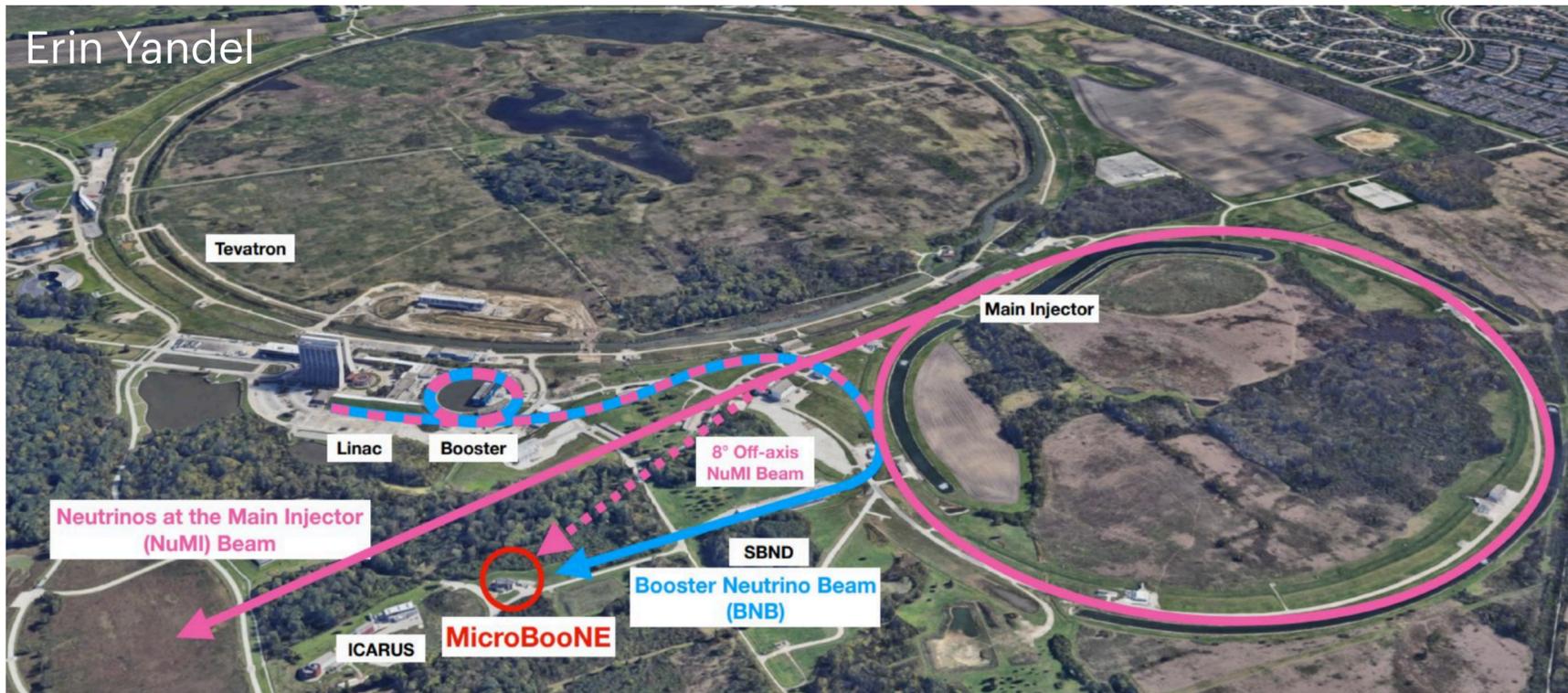
Messenger of new physics at energy scale f_a



Axion-Like-Particle

(pseudo-Goldstone bosons from lepton flavor symmetries)

Using the off-axis NuMI beam at **MicroBooNE** and **ICARUS**.



Axion-Like-Particle

(pseudo-Goldstone bosons from lepton flavor symmetries)

Search for Hidden Particles (SHiP)

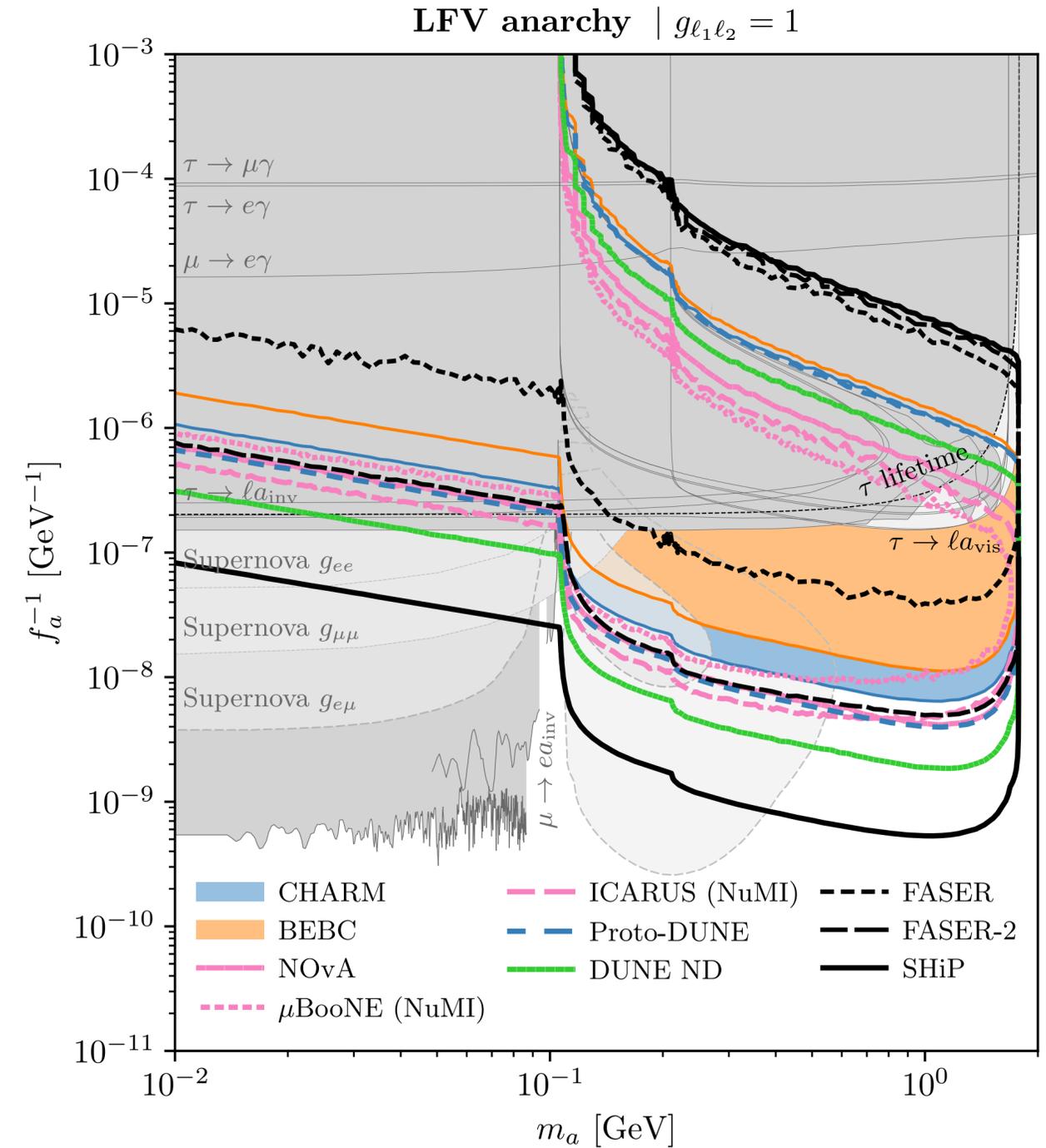
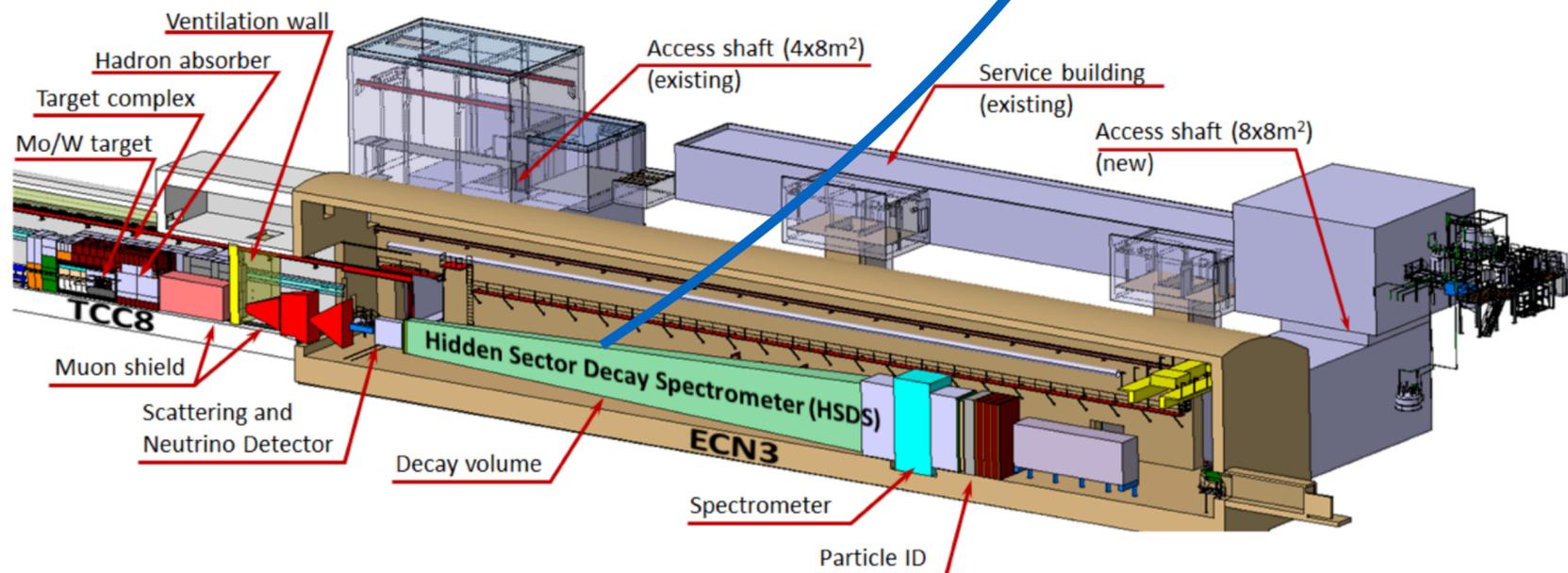
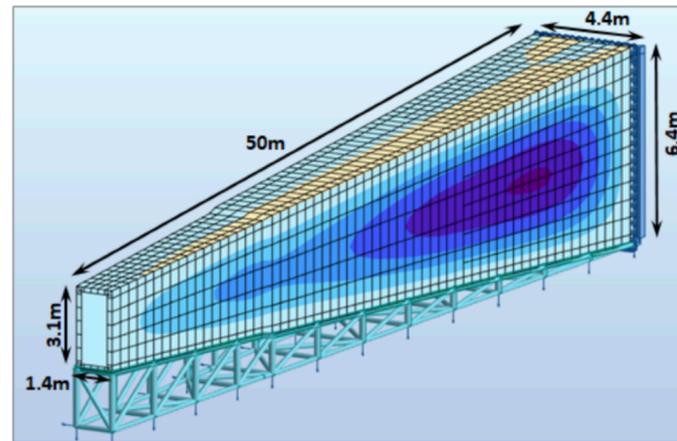
$E_p = 400 \text{ GeV}$ with $\mathcal{O}(10^{20})$ POT.

Largest acceptance.

“Commissioning and first data-taking planned for 2031-2033”

SHiP Collaboration & HI-ECN3 Project Team (ESPPU 2026)

“Pyramidal frustum”



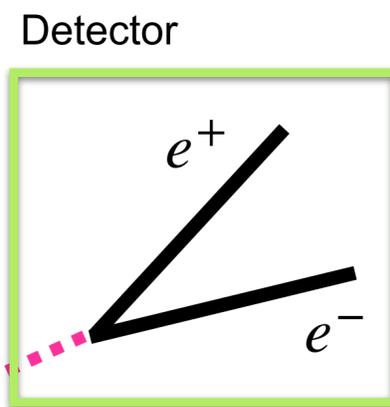
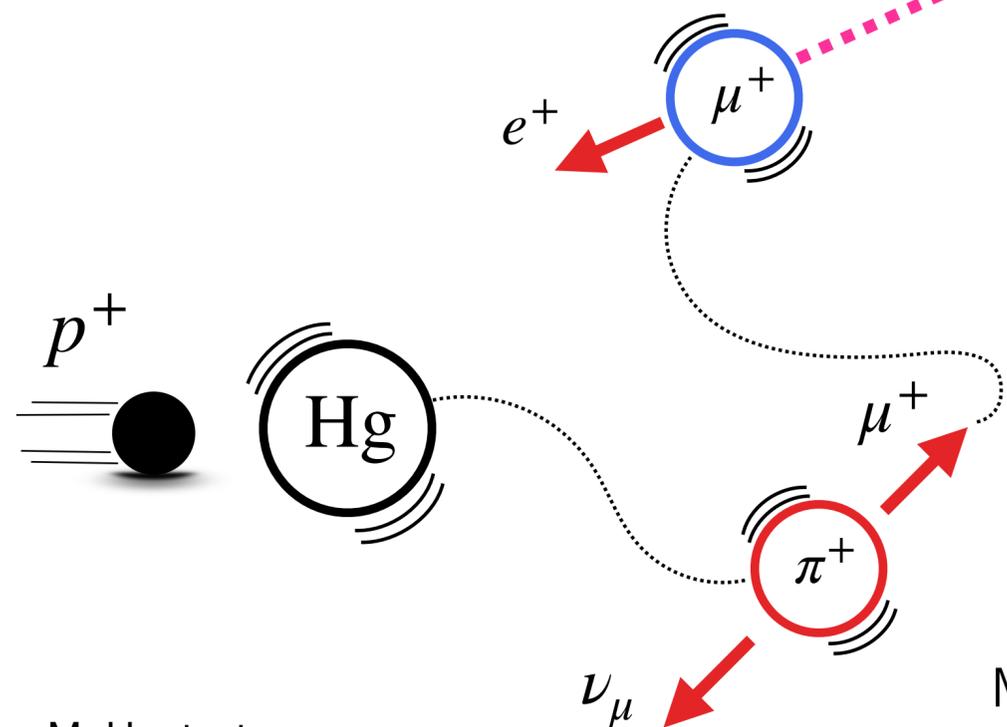
Axion-Like-Particle

(pseudo-Goldstone bosons from lepton flavor symmetries)

Spallation sources: $N_{\text{POT}} \lesssim N_{\text{Av0}}$

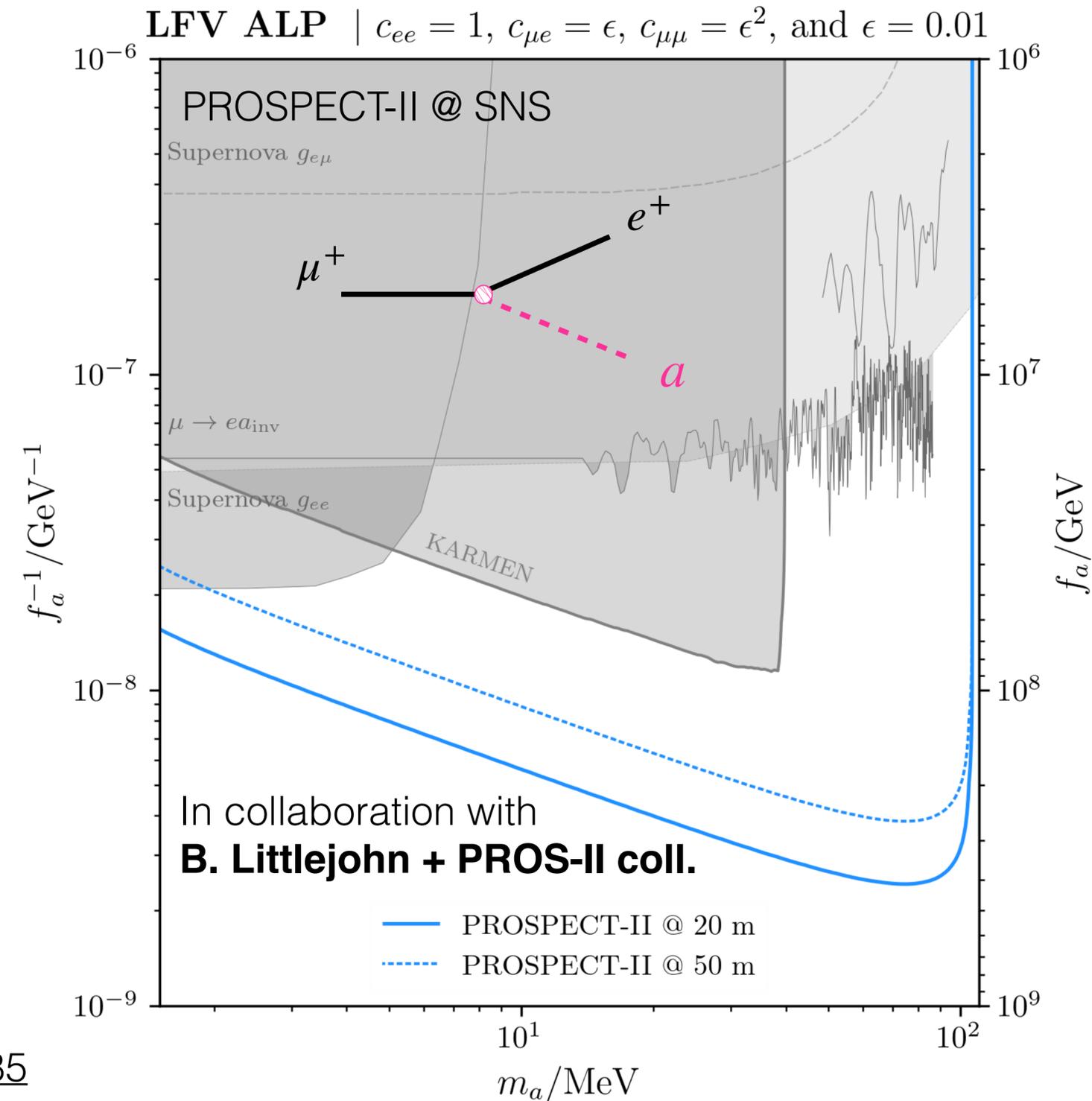
π^+ and μ^+ decay at rest.

Low-energy source, less boosted a .

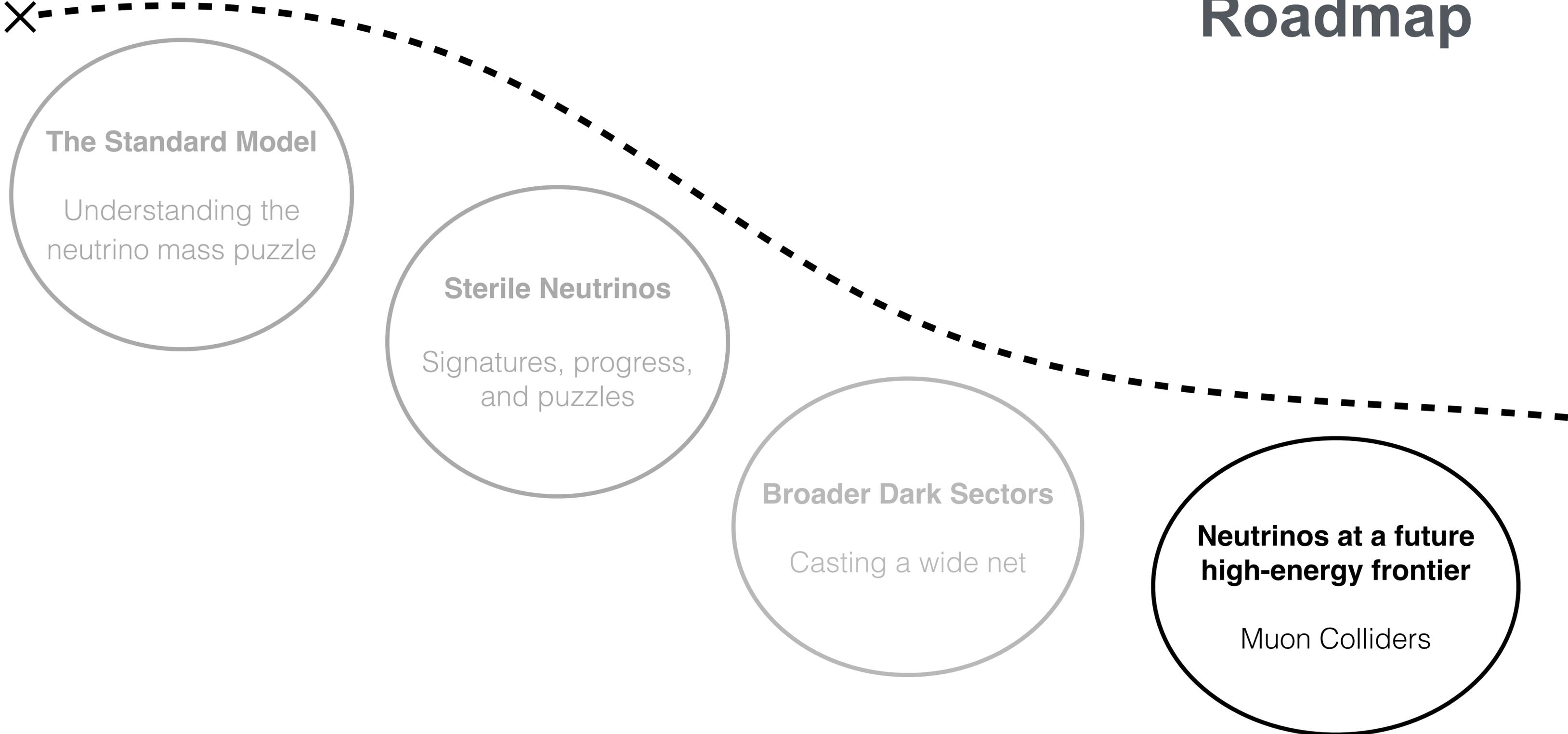


Just one example from

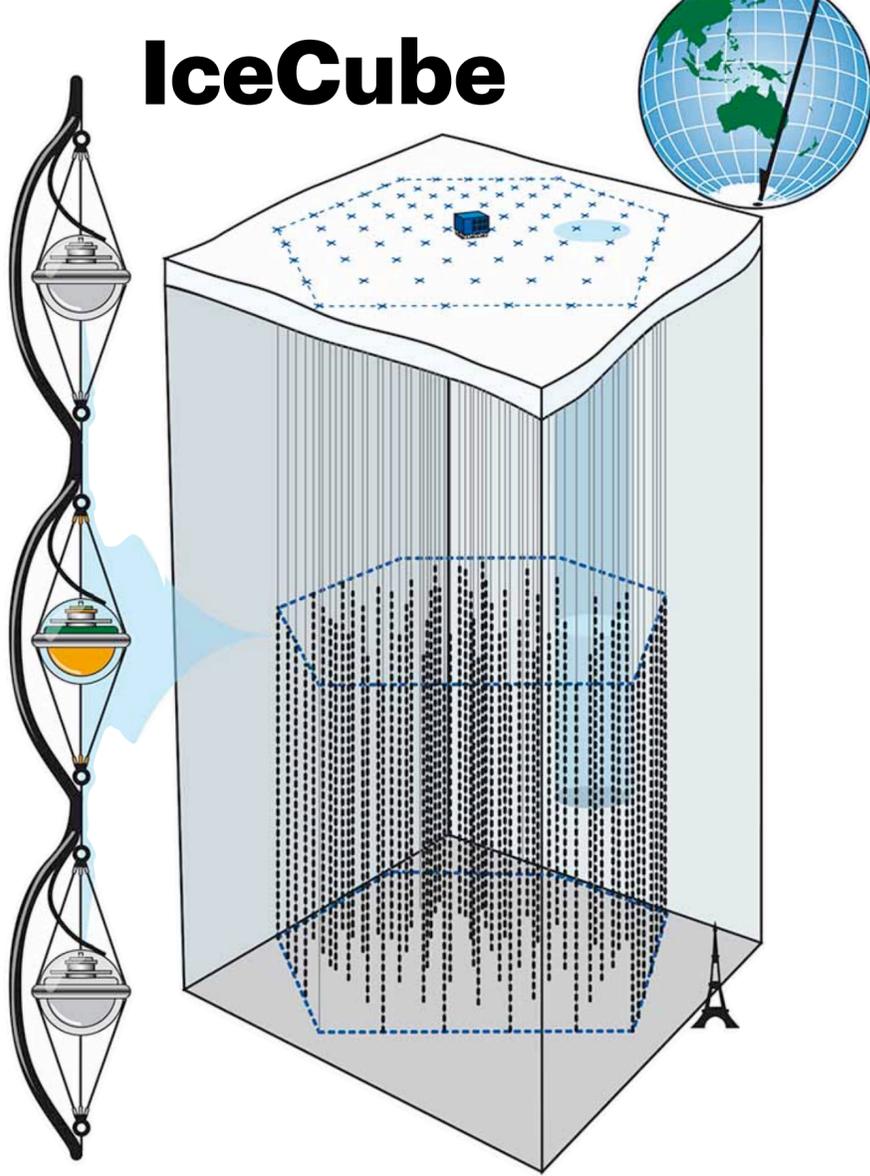
MH, S. Urrea, [arXiv:2509.14085](https://arxiv.org/abs/2509.14085)



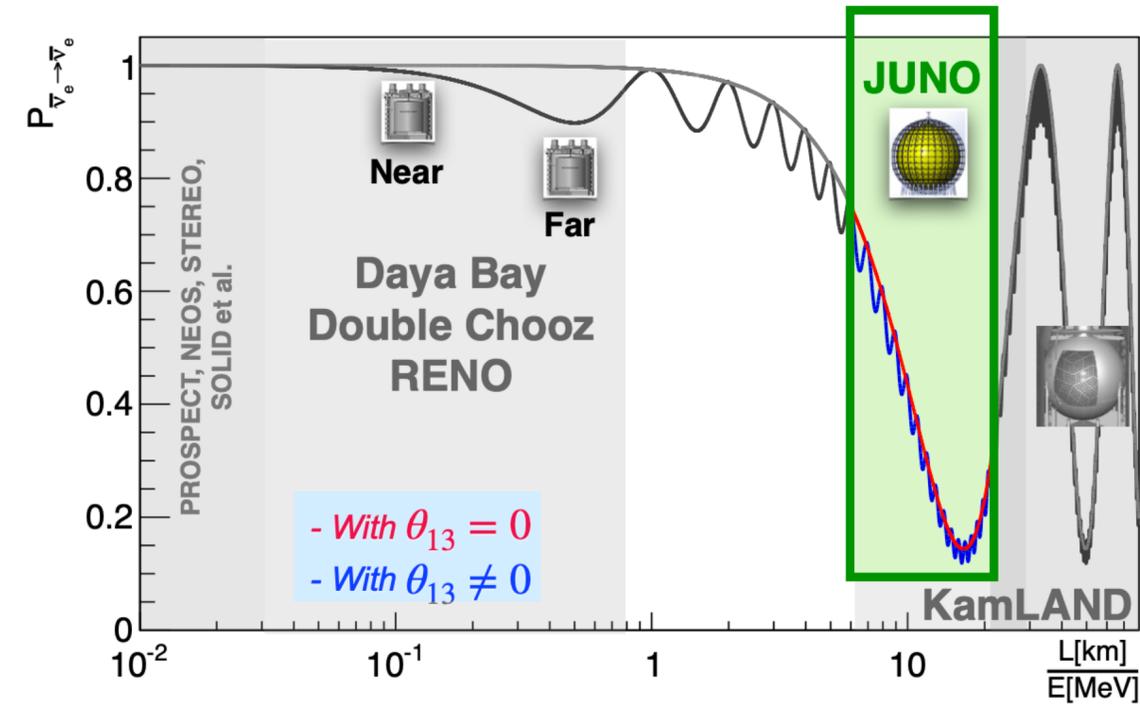
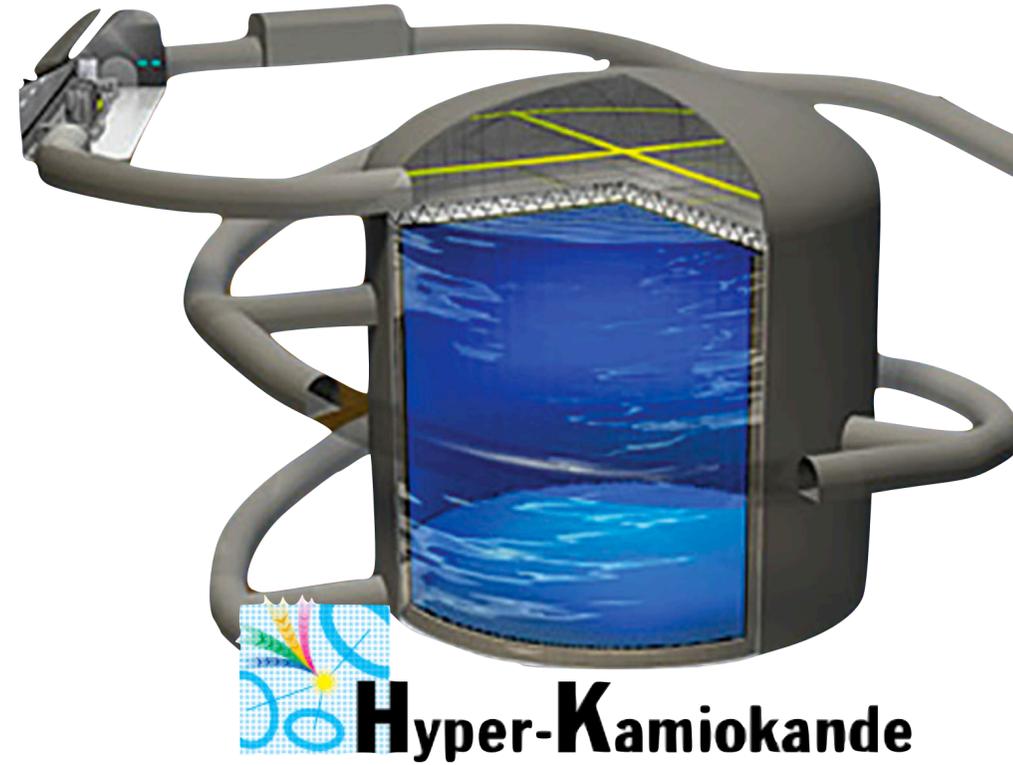
Roadmap



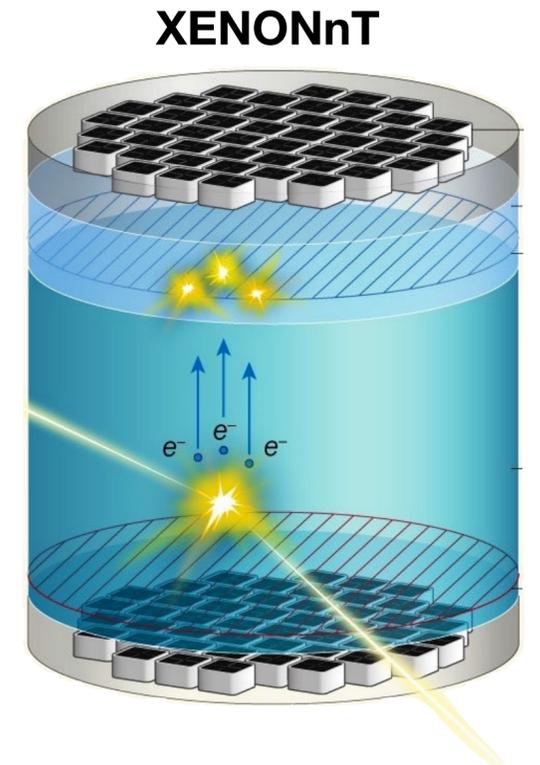
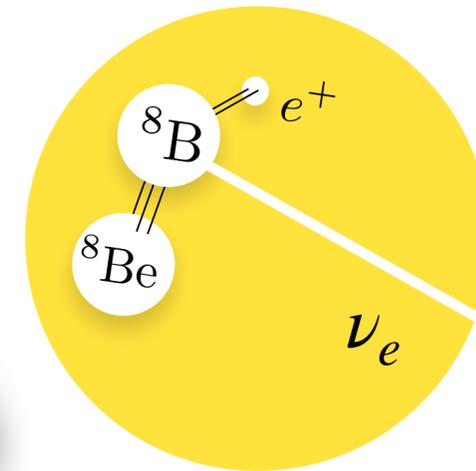
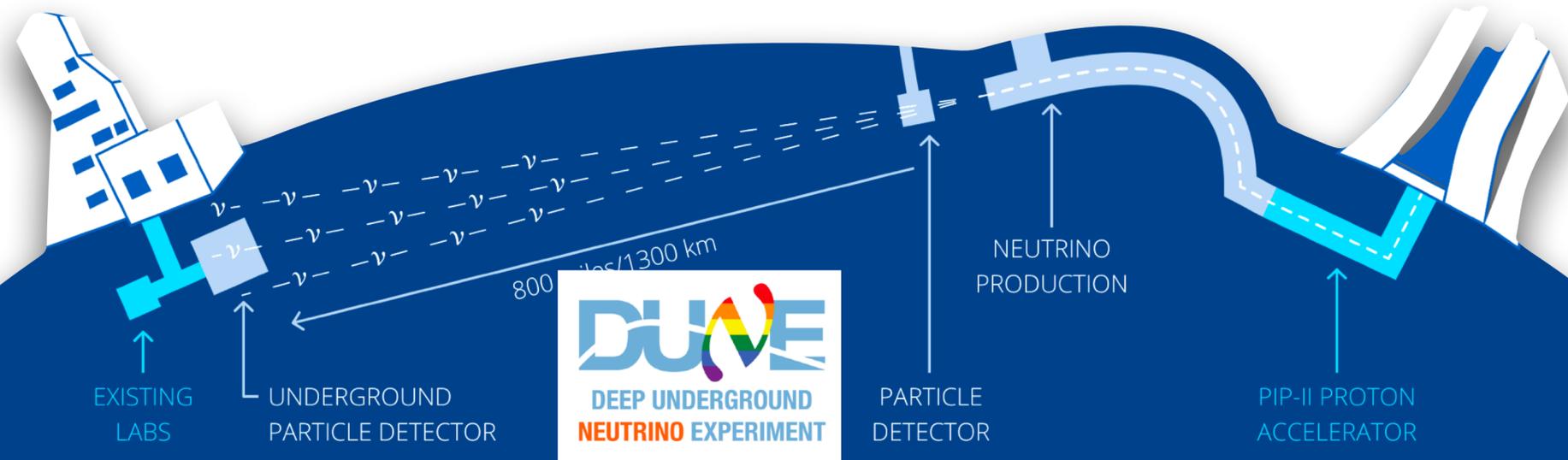
IceCube



Looking to the future



Adapted from P. Ochoa-Ricoux, NPN 2025



Next-generation accelerator neutrinos can boast about:

- 1) **Beam intensity:** Multi-MW beams (J-PARC, LBNF) ✓
- 2) **Detector size:** DUNE: 10's of kt, HyperK: 100's of kt. ✓
- 3) **Detector granularity:** LAr is a microscope on neutrino byproducts ✓
- 4) **Detector threshold:** dark matter detectors for ν 's ✓

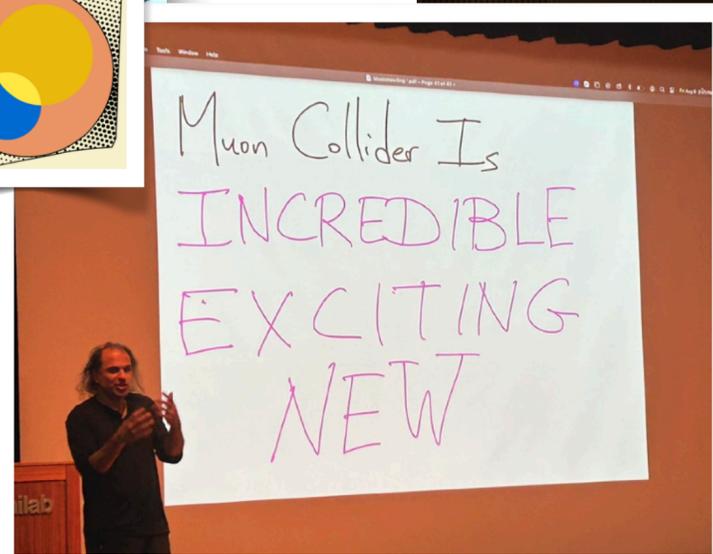
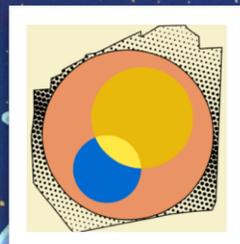
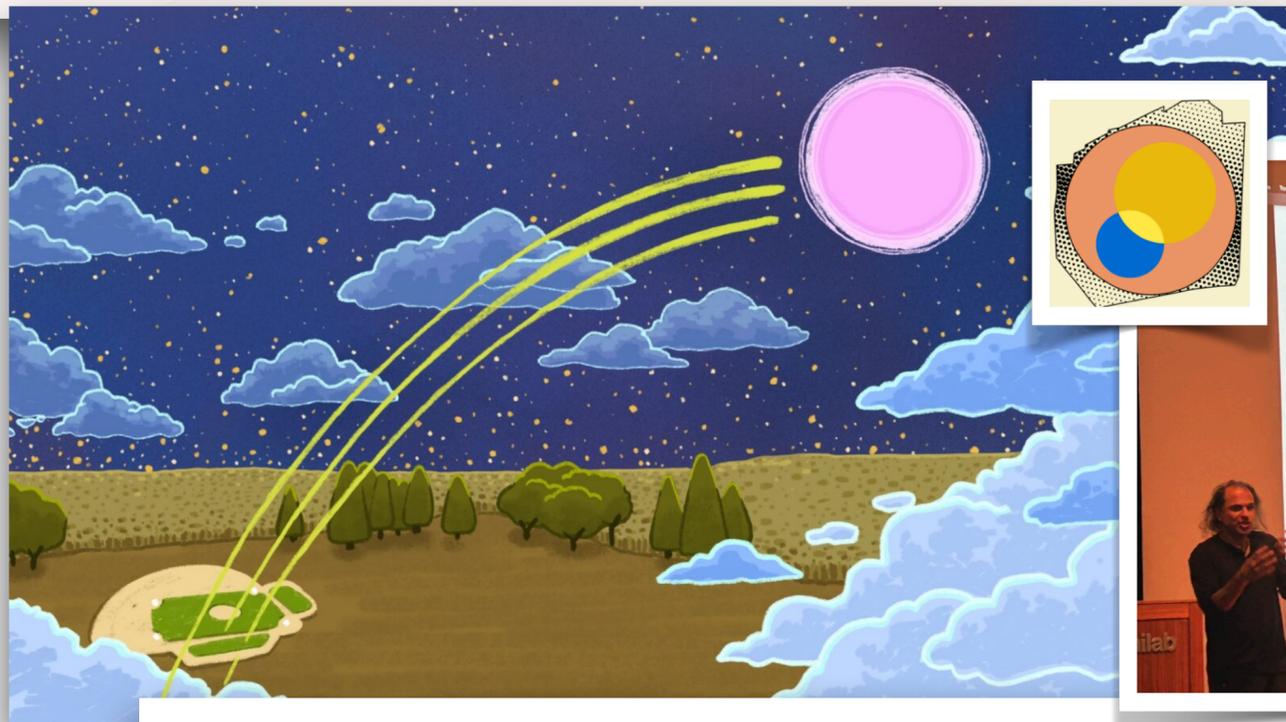
But we are still having to find work-around for **flux systematics**. ?

Beams from a parent particle we understand and can manipulate: Muons!

The Muon Collider

Symmetry Magazine

'This is our Muon Shot'



Science Magazine



Particle Physicists Agree on a Road Map for the Next Decade

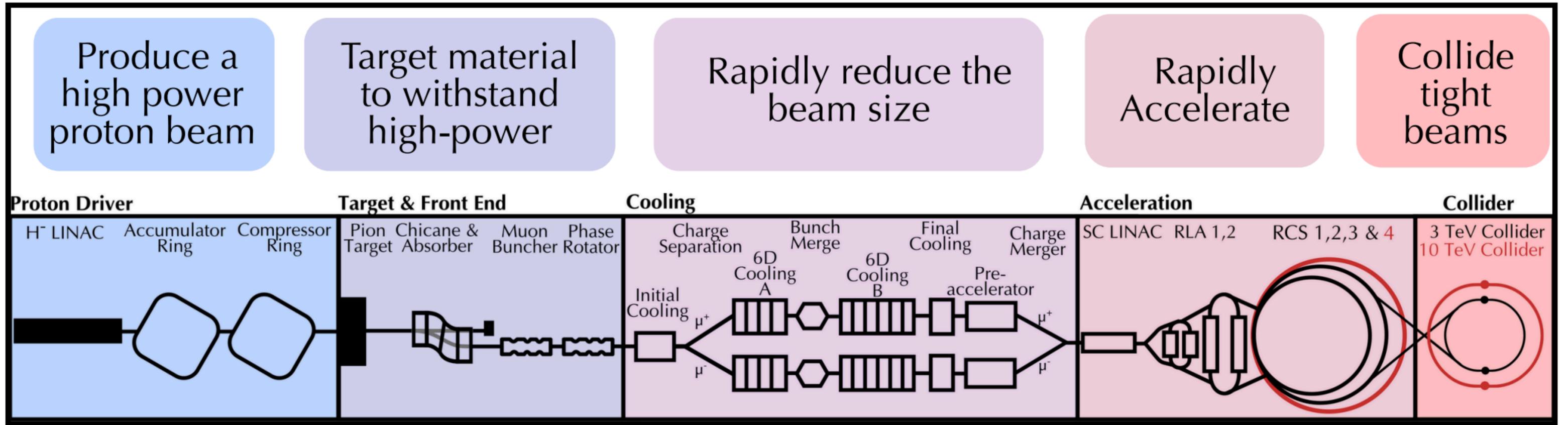
A “muon shot” aims to study the basic forces of the cosmos. But meager federal budgets could limit its ambitions.

“

The challenge, if you want to capture it in one word, is that the muon is unstable.

SERGO JINDARIANI | FERMI NATIONAL ACCELERATOR LABORATORY





Proton Target

multi-MW and ns-pulsed

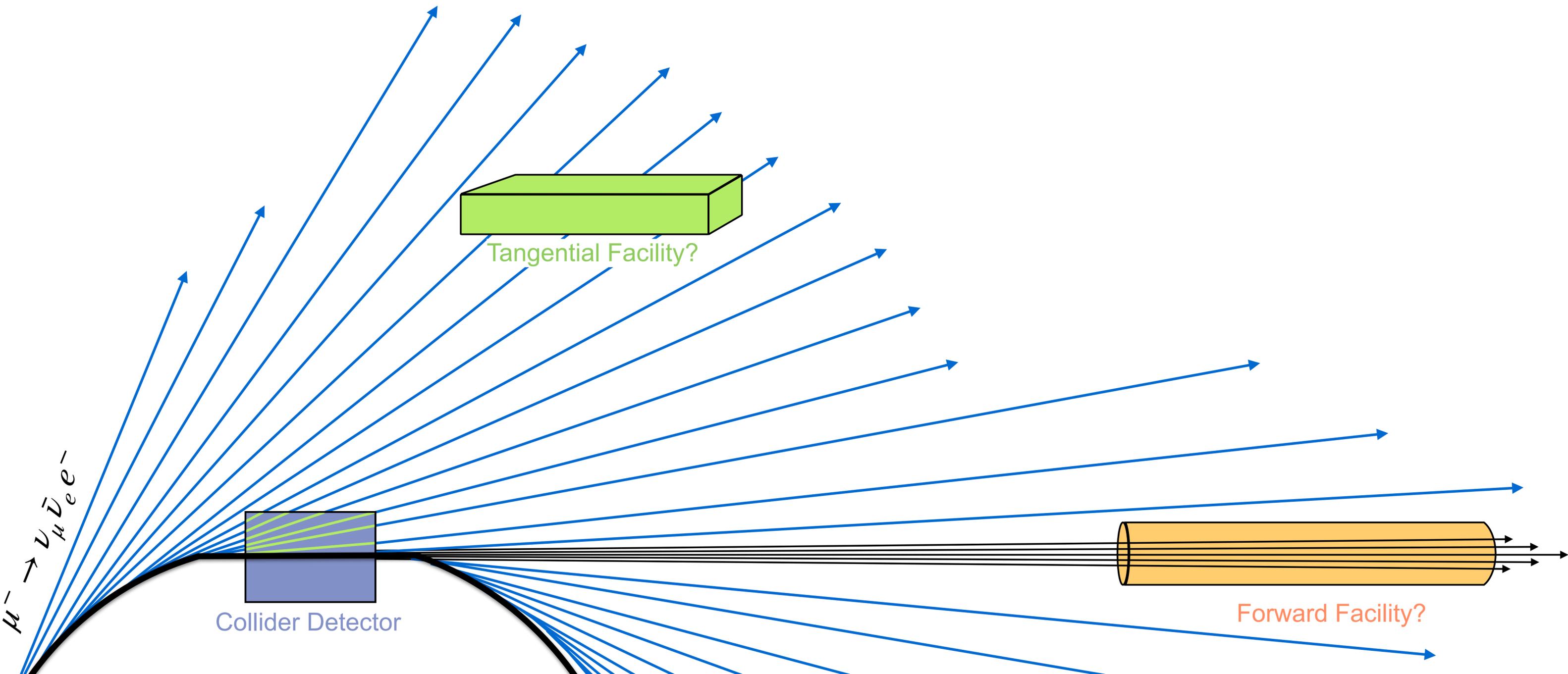
Muon cooling to increase luminosity

ionization in this case.

Low-energy acceleration

Muon Collider

Forward and **Tangential?** physics: leverage the neutrino byproducts of a MuC.

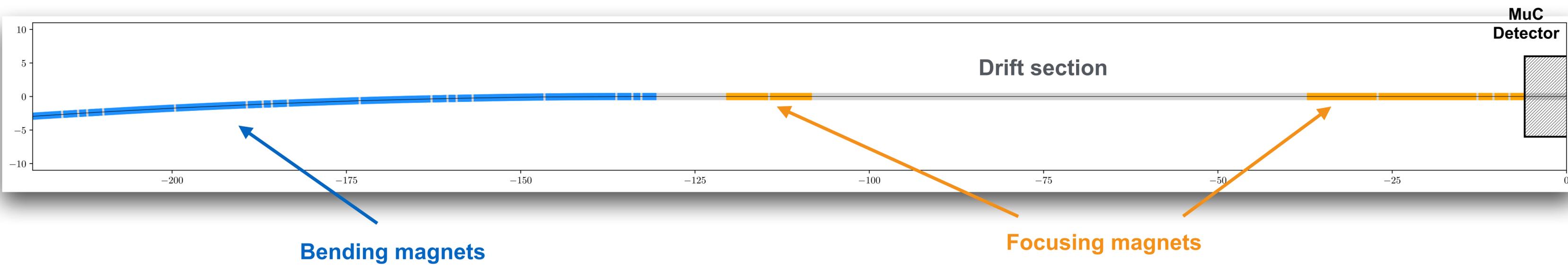
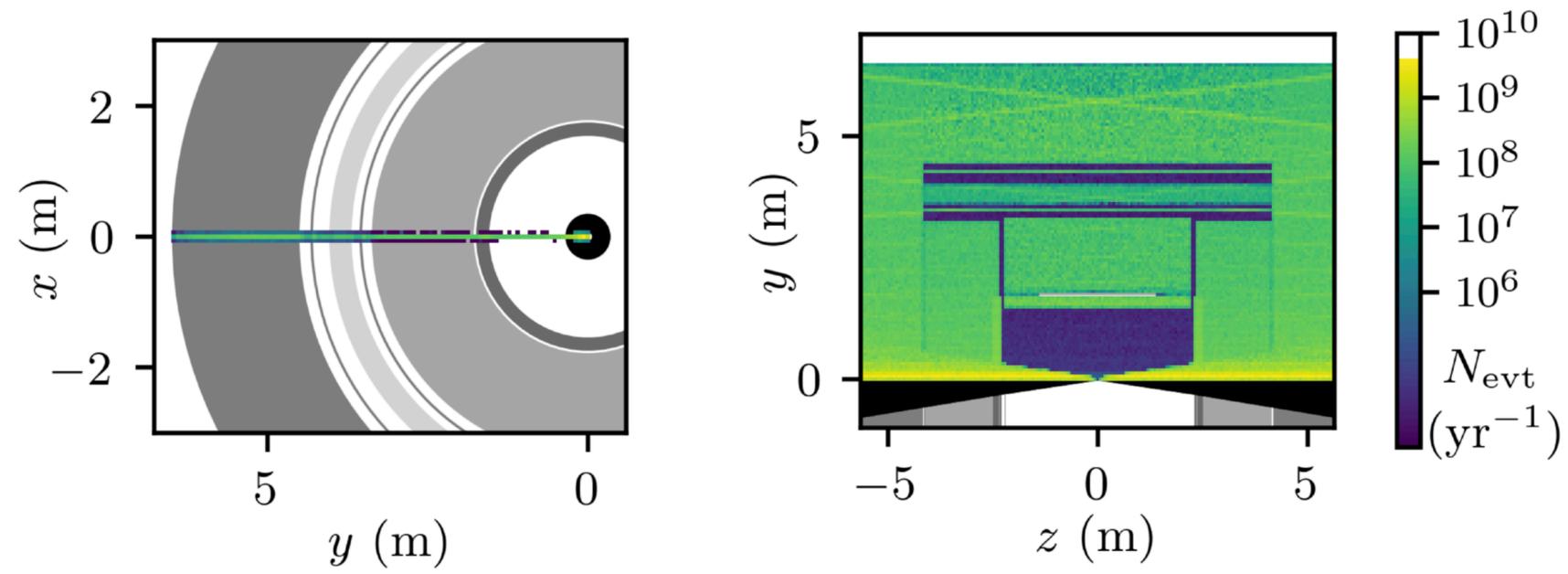


Beam-induced neutrinos

Tangential fluxes

L. Bojorquez-Lopez (Undergrad), **MH**, C. A. Argüelles, Z. Liu
The Neutrino Slice at Muon Colliders

Detector components following IMCC studies



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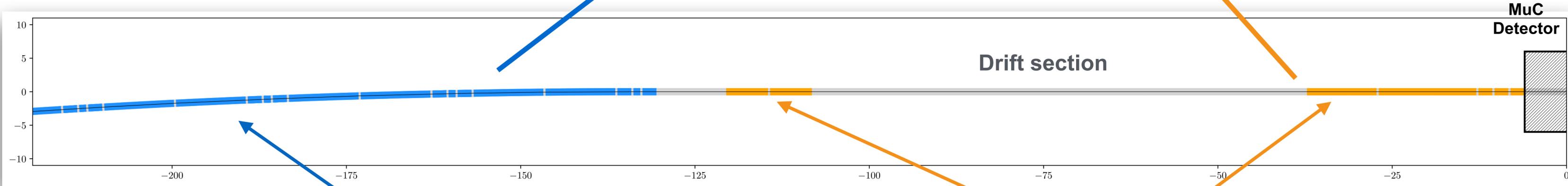
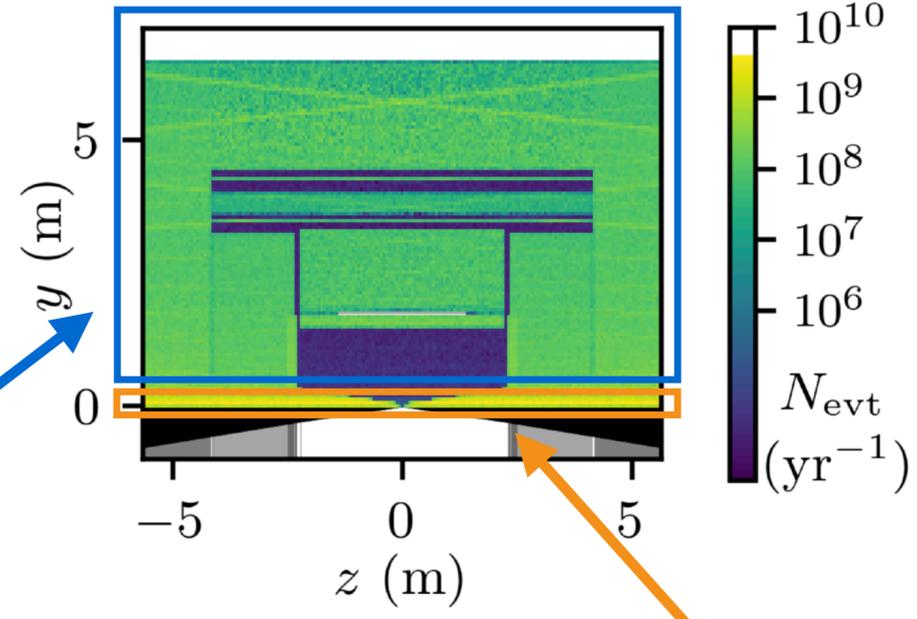
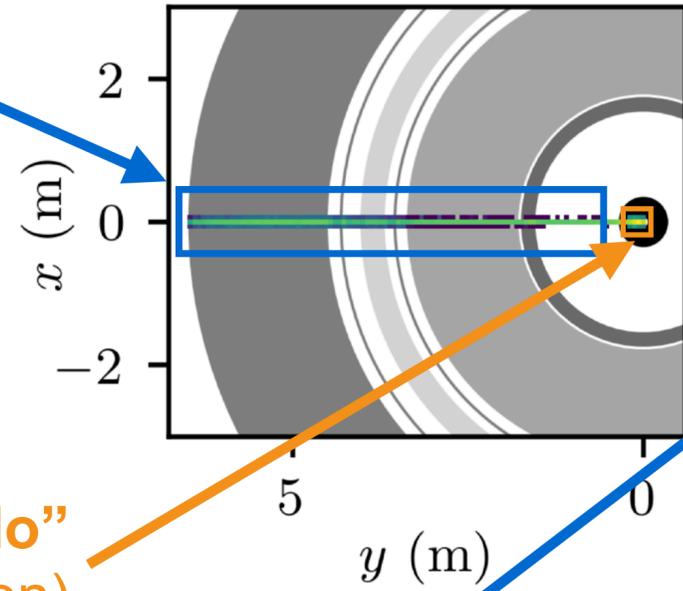
Detector components following IMCC studies

Neutrino "Slice"
(mostly from main ring)

Neutrino "Halo"
(mostly from straight section)

Bending magnets

Focusing magnets

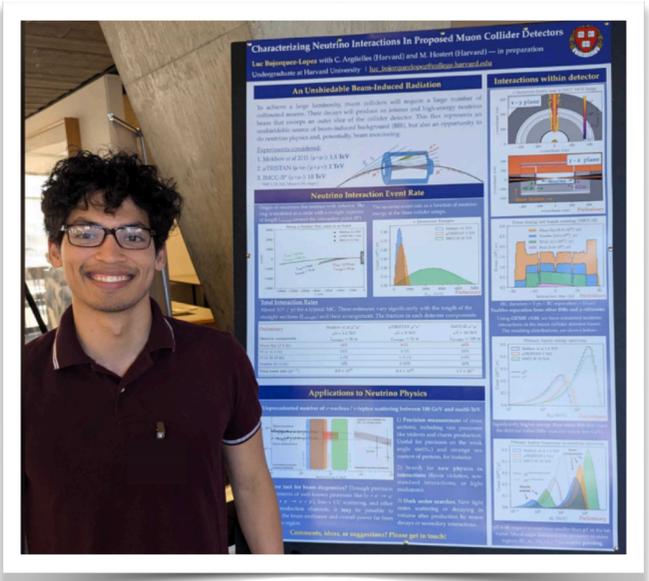


Muon Collider

Building toy models of muon collider

More neutrinos than we have ever detected by >3 orders of magnitude.

About 0.5 events per bunch crossing for 10 TeV



Luc Bojorques-Lopez (Harvard)
Inaugural US muon collider community meeting @ Fermilab

Collider	MuC 10 TeV	MuC 3 TeV	μ TRISTAN
Beams	$\mu^+ \mu^-$	$\mu^+ \mu^-$	$\mu^+ \mu^+$
Muons/bunch	1.8×10^{12}	1.8×10^{12}	1.4×10^{10}
bunches/cycle	1	1	40
f_{inj}	5 Hz	5 Hz	50 Hz
C	8.7 km	4.3 km	4.3 km

BIN exclusive reactions in HCAL and ECAL/year			
Total NC	1.5×10^9	4.6×10^8	3.4×10^9
Total ν_e CC	4.7×10^9	1.4×10^9	1.1×10^{10}
Total ν_μ CC	5.4×10^9	1.7×10^9	1.1×10^{10}

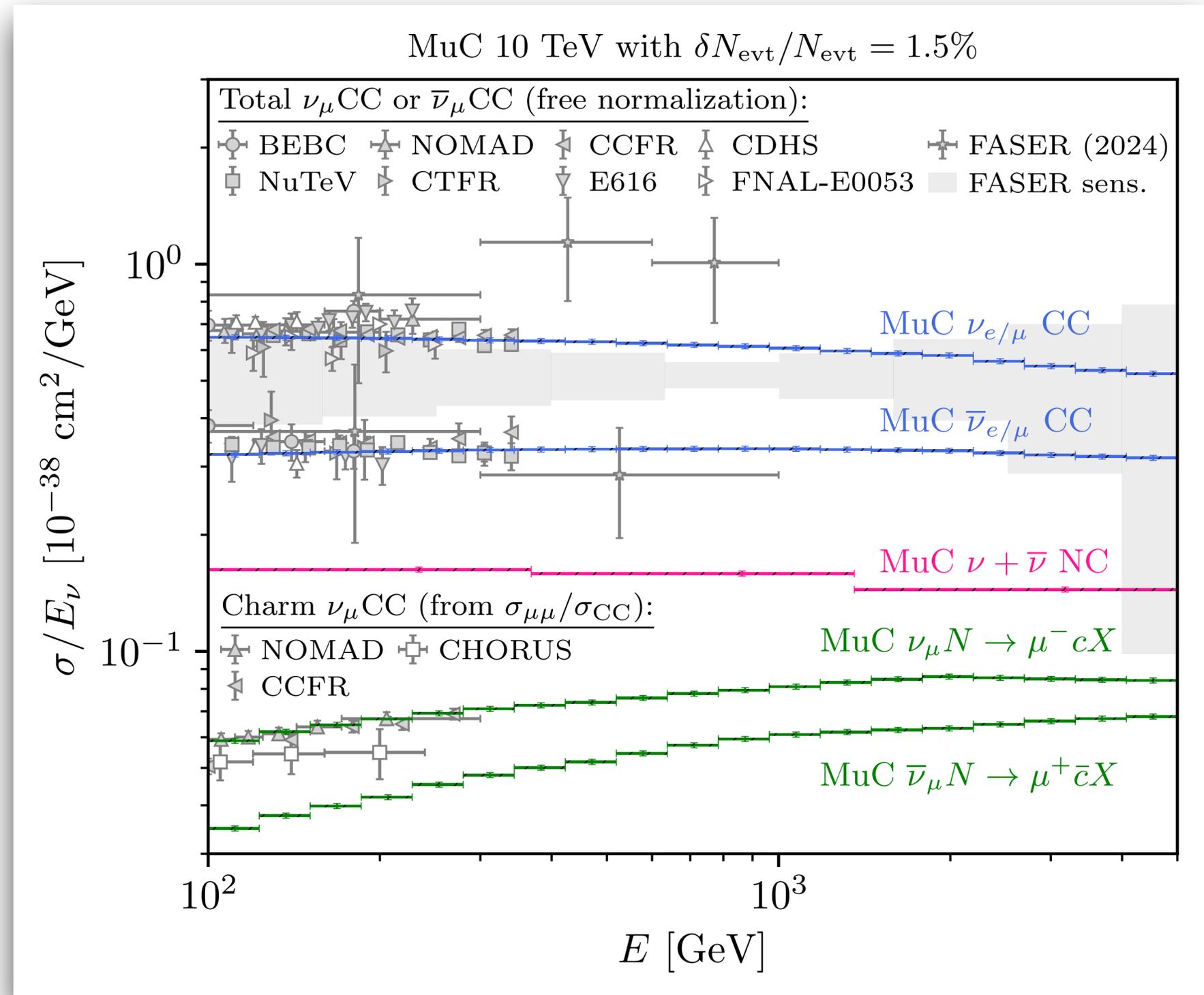
ES $\nu_\mu e \rightarrow \nu_\mu e$	3.8×10^5	1.1×10^5	0
ES $\bar{\nu}_e e \rightarrow \bar{\nu}_e e$	8.6×10^5	2.5×10^5	0
ES $\bar{\nu}_\mu e \rightarrow \bar{\nu}_\mu e$	3.4×10^5	9.9×10^4	1.9×10^6
QE $\nu n \rightarrow \ell^- p^+$	2.6×10^6	2.5×10^6	2.8×10^7
QE $\bar{\nu} p^+ \rightarrow \ell^+ n$	2.7×10^6	2.5×10^6	3.2×10^7
Coh π^0	3.0×10^5	2.9×10^5	3.5×10^6
Res $\bar{\nu}_e e \rightarrow \rho^-$	4.2×10^5	7.7×10^5	0
Res $\bar{\nu}_e e \rightarrow K^{*-}$	2.6×10^4	4.4×10^4	0
IMD $\nu_\mu e \rightarrow \nu_e \mu^-$	4.2×10^6	1.2×10^6	0
IMD $\bar{\nu}_e e \rightarrow \bar{\nu}_\mu \mu^-$	1.2×10^6	3.5×10^5	0
ITD $\bar{\nu}_e e \rightarrow \bar{\nu}_\tau \tau^-$	9.4×10^3	0	0
Trident $e^+ e^-$	1.2×10^6	2.9×10^5	1.7×10^6
Trident $\mu^\pm e^\mp$	2.9×10^6	6.7×10^5	5.0×10^6
Trident $\mu^\pm \mu^\mp$	7.5×10^5	1.6×10^5	1.3×10^6

BIN interactions

Neutrino-nucleus scattering measurement

This is what a 1.5 % uncertainty on TeV neutrino cross sections would look like:

Clearly more works required to understand if feasible at the neutrino slice, but such precision can undoubtedly be achieved with a forward detector if needed.



High Energy ν Interactions

MuC BINs would offer a high- Q^2 probe of fundamental matter with Weak interactions.

Access to many other rare exclusive processes.

Example of classes of exclusive processes

Deep Inelastic Scattering **CC** and NC

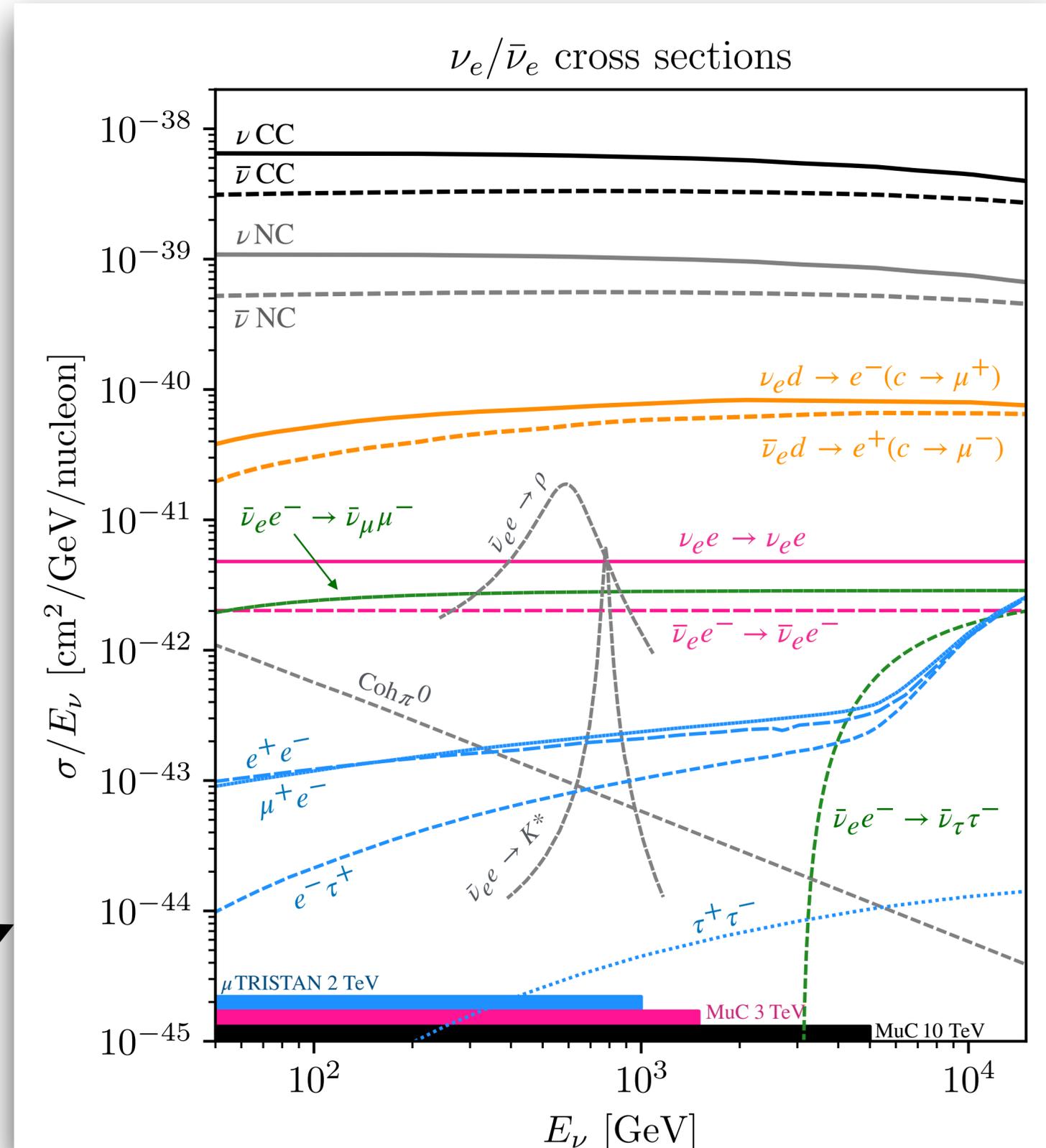
Charm production

Inverse muon(tau) decay

Elastic scattering on electrons

Resonant meson production

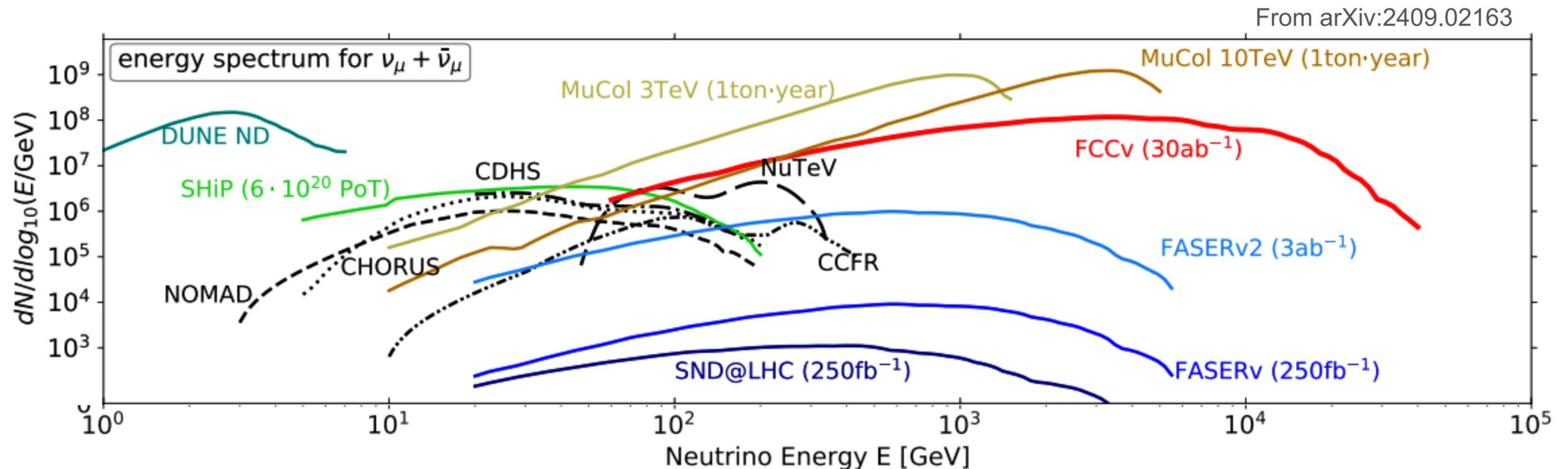
Neutrino trident production



MuC Neutrino Energies

High energy and extremely collimated neutrino beam: $\langle E_\nu^{\text{FASER}} \rangle \lesssim \langle E_\nu^{\text{MuC-10-TeV}} \rangle \lesssim \langle E_\nu^{\text{FCC}} \rangle$

Beam energies are not necessarily new, but flavor composition, sample size, and precision are.



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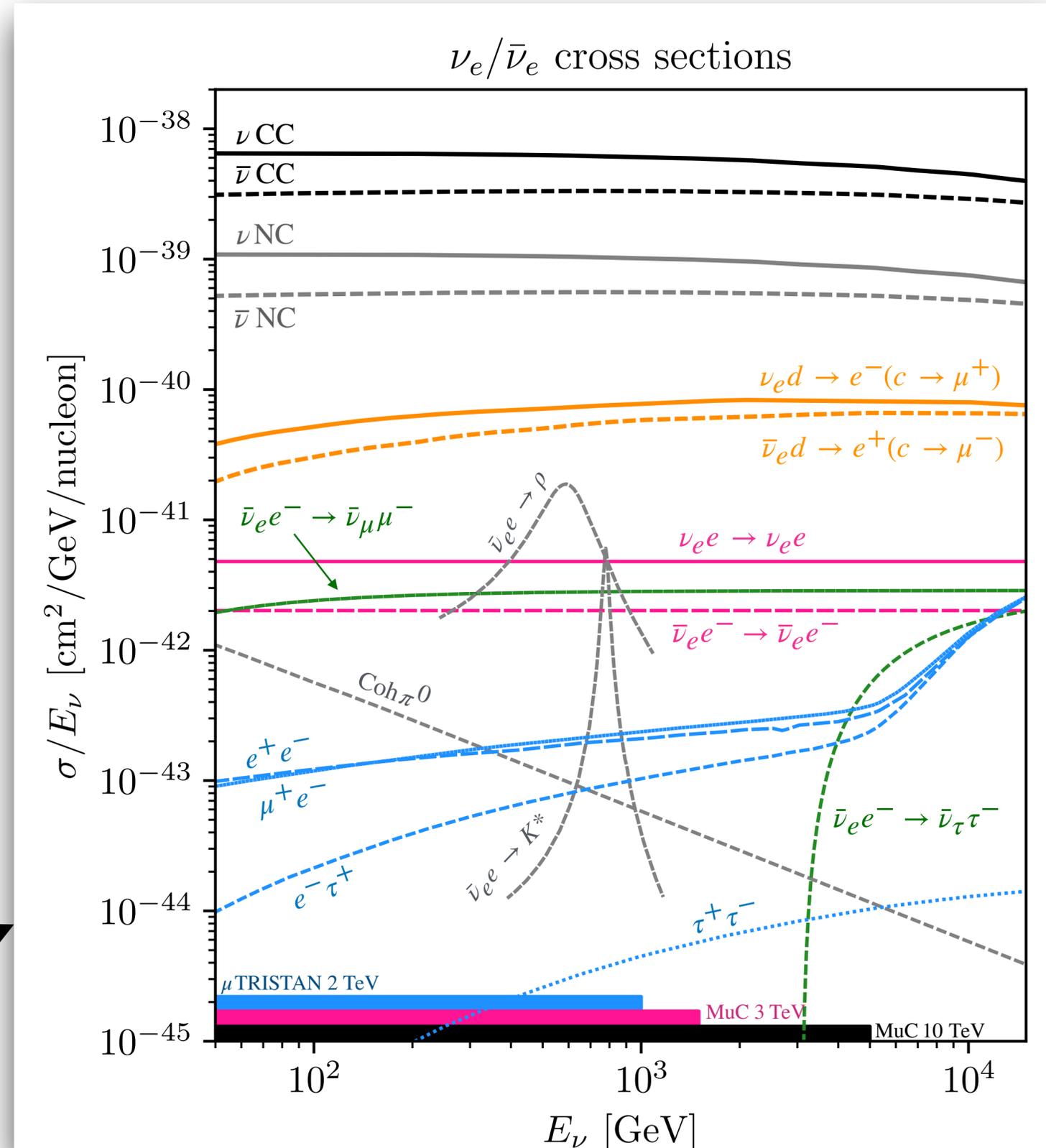
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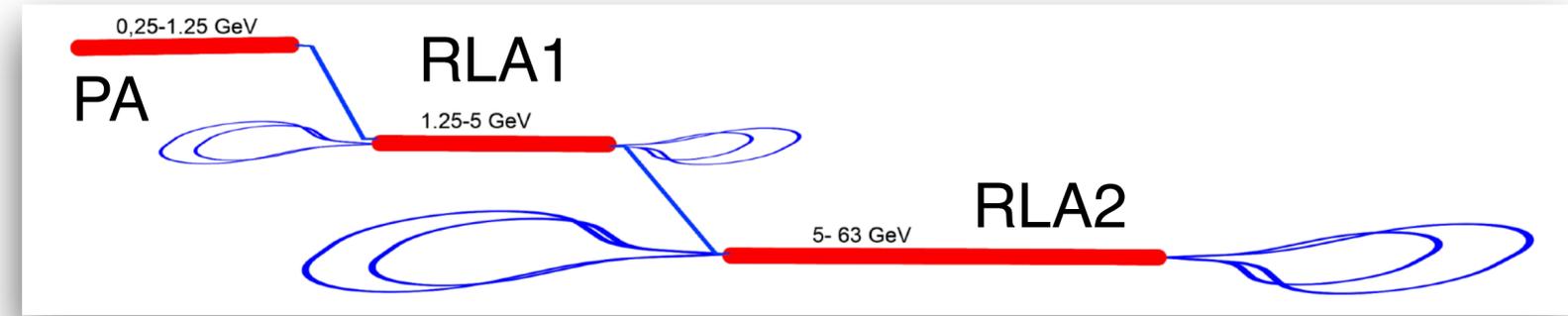
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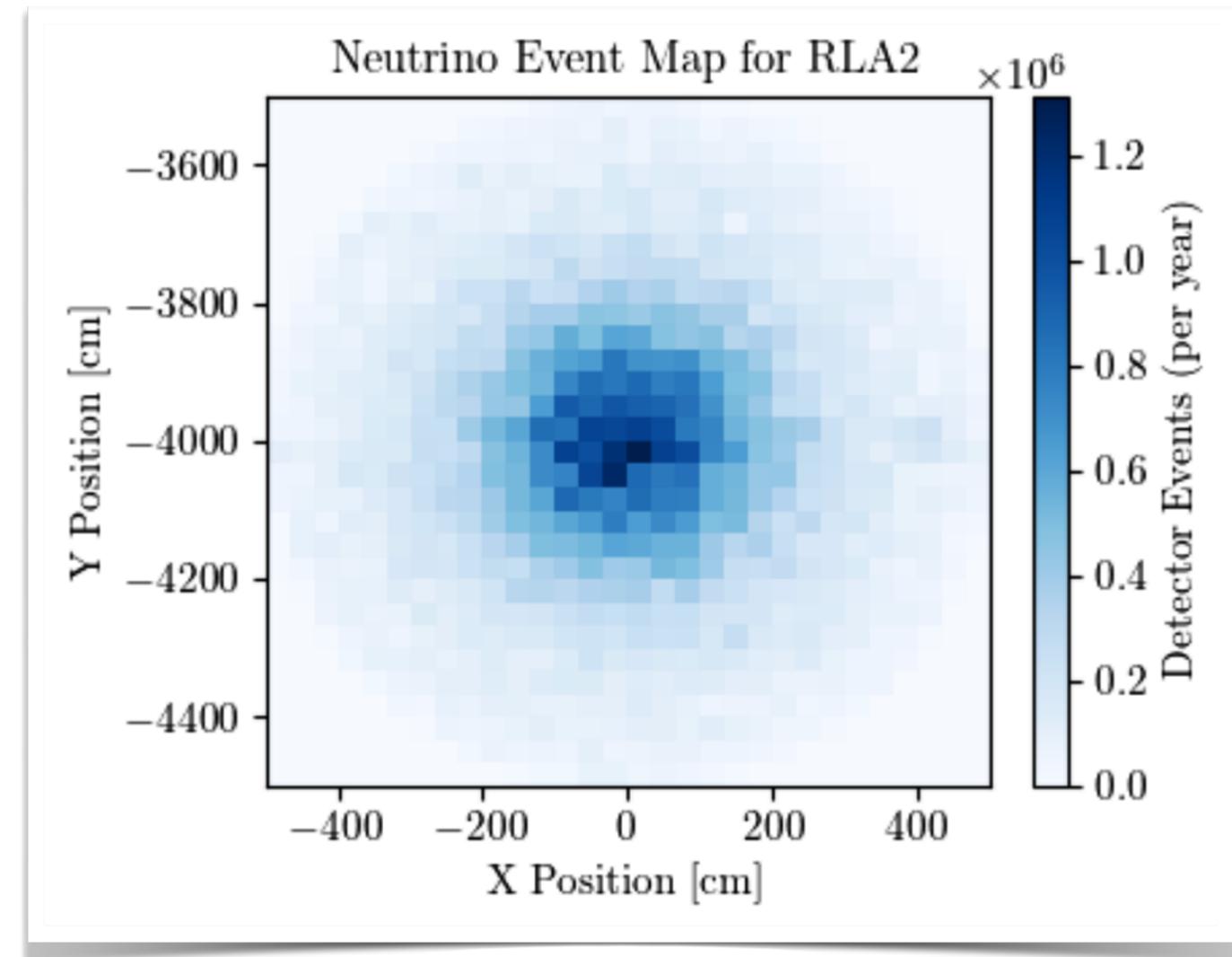
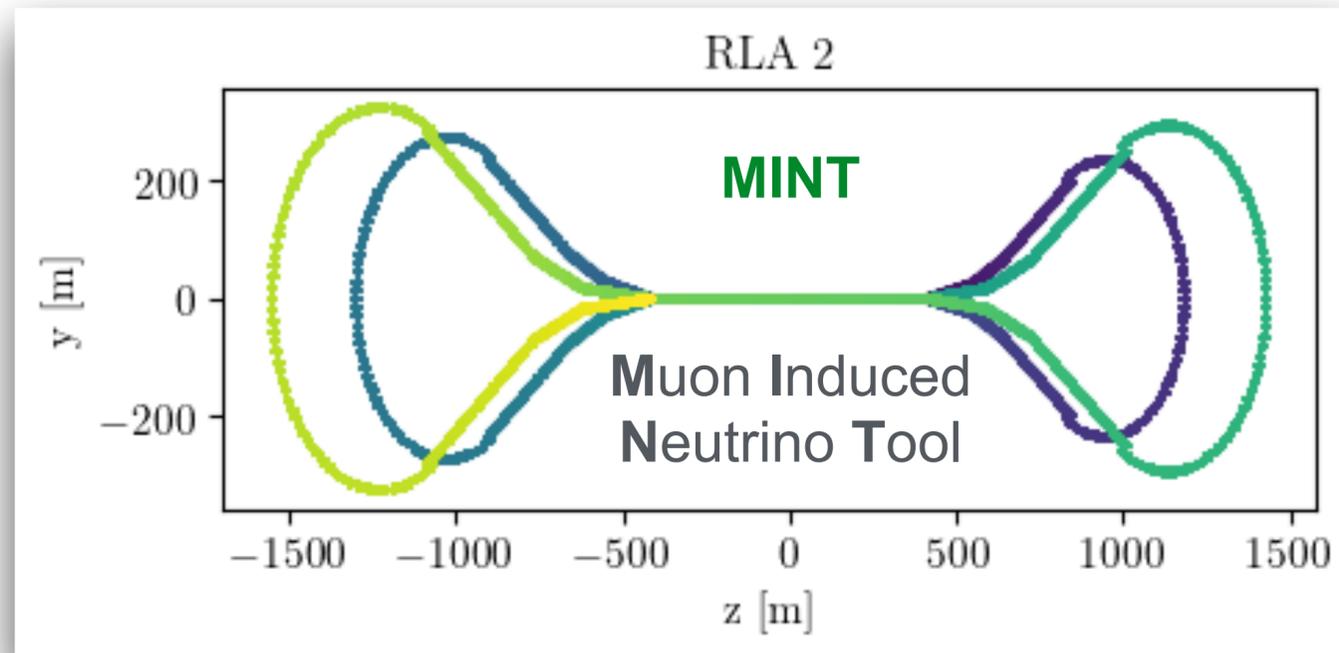
Neutrino trident production



Muon Acceleration Neutrinos?

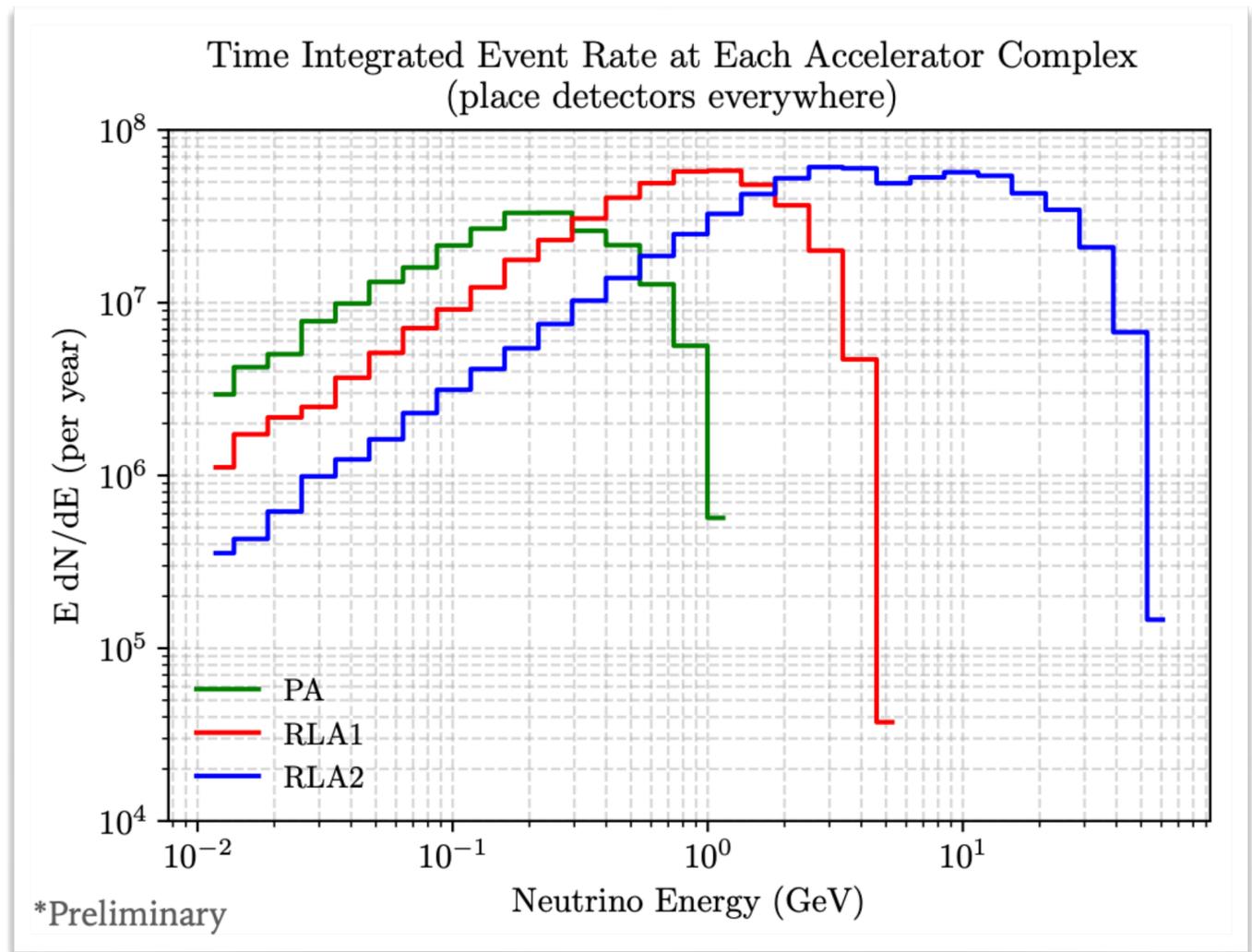
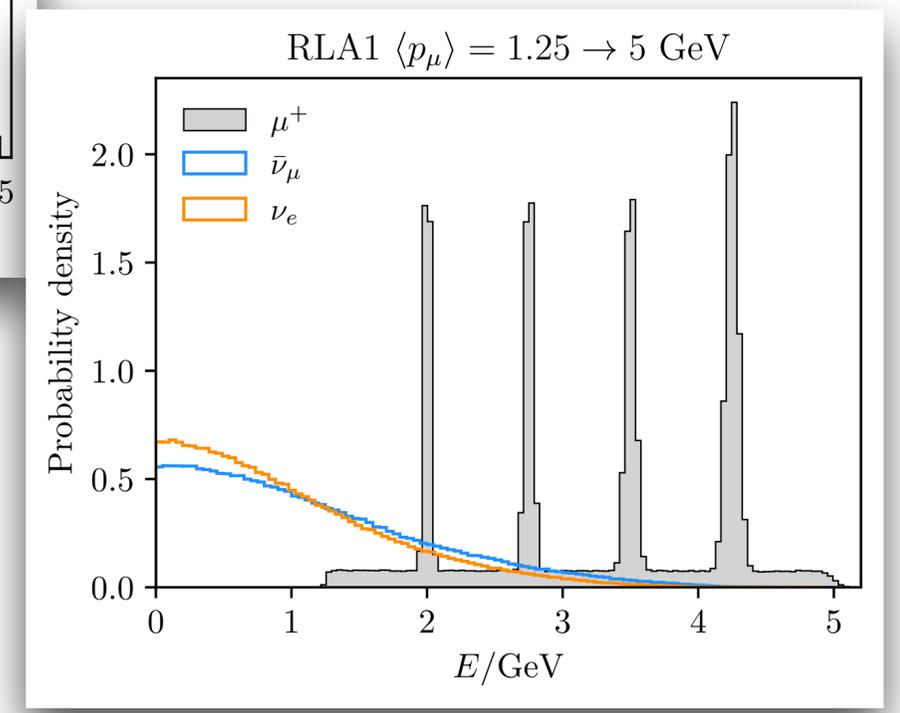
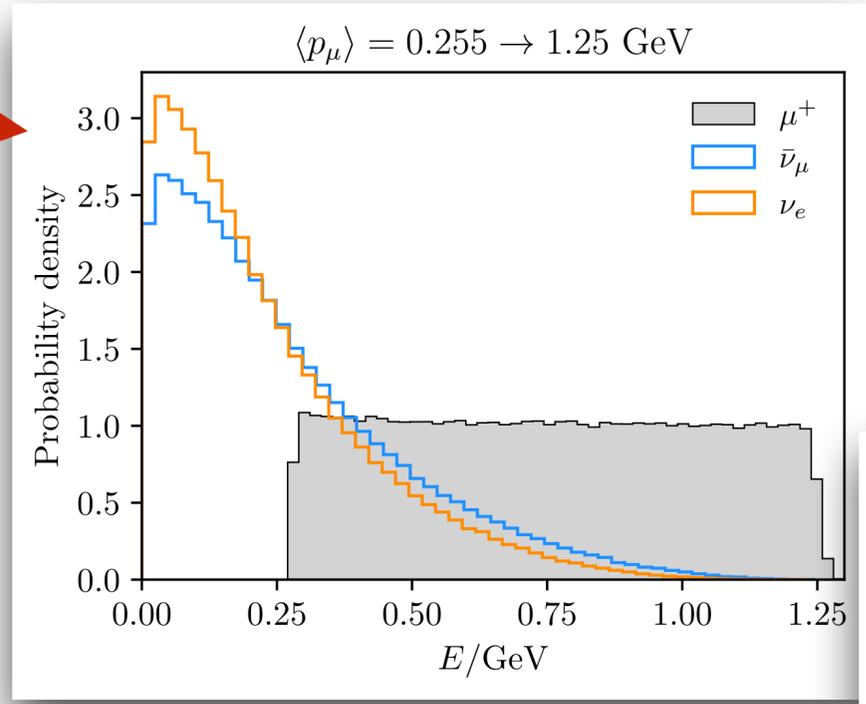
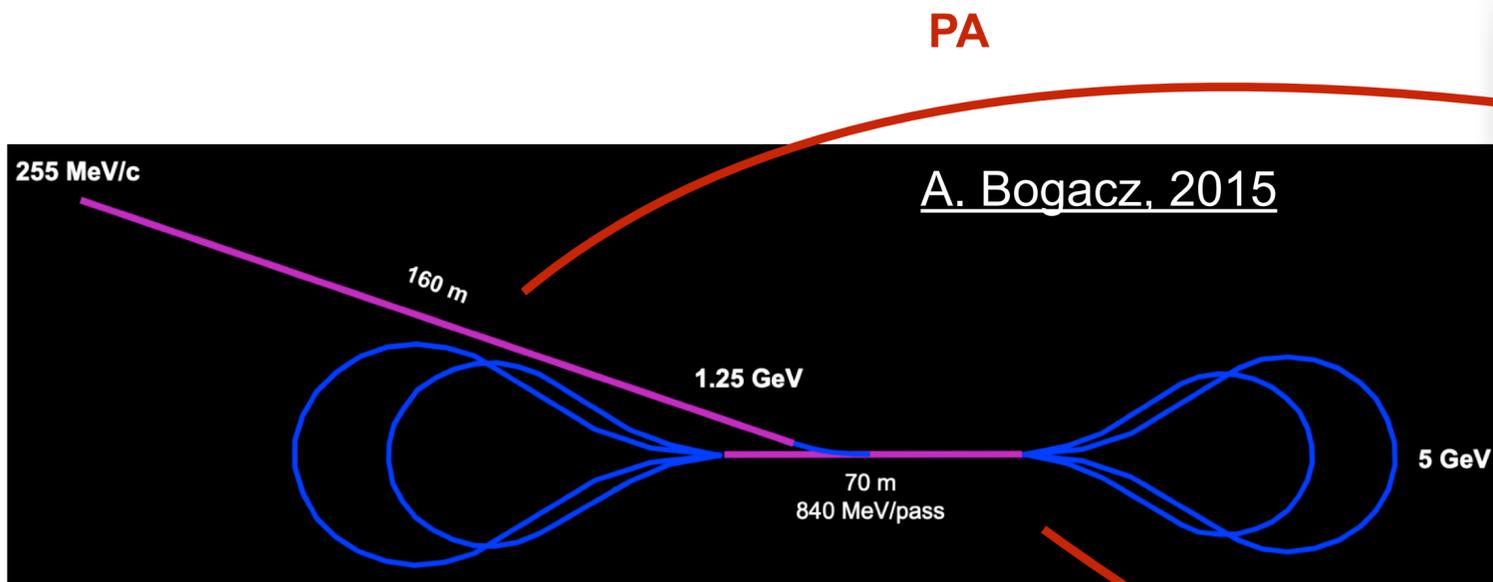


Currently investigating the entire low-energy acceleration stage of the Muon Collider complex.



Joel Choi
(U. Iowa)





RLA1

Spans a wide range of energies:

Physics from **100's of MeV to the TeV scale**
in the same machine.

In fundamental physics there are no guarantees:

A fast (idea \rightarrow result) turnover is usually the exception in Particle Physics.

Theory work can be fast, but building experiments/collecting data can take time.

Our job is to **explore different topics and energy scales**, learn about **new technologies**, and **advance large-scale projects in science**,

Having an active phenomenology community is a sign of a healthy stream of data and ideas.

The need to cast wide nets makes this a diverse field and that's exciting!



Take Home

There is “**reasonable doubt**” that the SM is incomplete already **at low energies.**

Many possibilities: the challenge is in covering them all. **Need data!**

Particle physics entered a new era of studying **rare phenomena** and phenomenology will be crucial to help us navigate it.

More neutrinos, but can also hope for “better” neutrinos.

A good time to ask: “Cool, what else can we do with this?”

